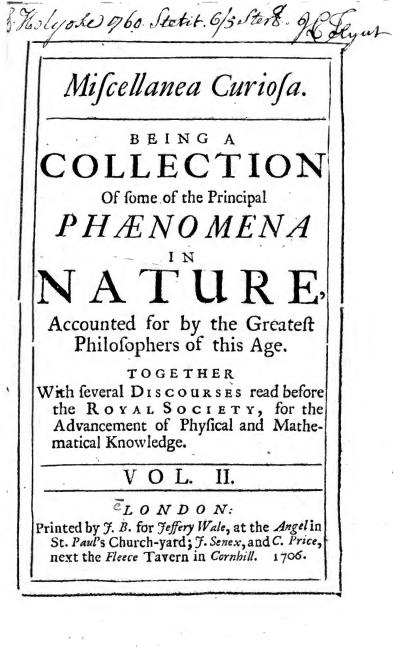


inguades Google

2 Soc 10.5 (2) 8827 LSn 10. LNVD and Howeve Fresiden avail bollege. 1764-66. ON.





LSoc 10.5 (2)



4

THE

CONTENTS

A Calculation of the Credibility of Humane Testimony. Page 1 A Letter from the Reverend Dr. Wallis, Professor of Geometry in the University of Oxford, and Fellow of the Royal Society, London, 10 Mr. Richard Norris, concerning the Colle-Etion of Secants; and the true Division of the Meridians in the Sea-Chart. 9

An easie Demonstration of the Analogy of the Logarithmick Tangents to the Meridian Line or fum of the Secants; with various Methods for computing the same to the utmost Exactness, by E. Halley. 20

A most compendious and facile Method for Conftructing the Logarithms, exemplified and demonstrated from the Nature of Numbers, without any regard to the Hyperbola, with a speedy Method for finding the Number from the Logarithm given. By E. Halley. 37

A Solution given by Mr. John Collins, of a Chorographical Problem, Proposed by Richard Townley, Esq; 53

The Solutions of three Chorographic Problems, by a Member of the Philosophical Society of Oxford. 58

An Arithmetical Paradox, concerning the Chances of Lotteries; by the Honourable Francis Rob-A 2 crts

The CONTENTS.

erts, Elg; Fellaw of the R. S. 65
A New, Exact and Easte Method, of finding the
Roots of any Equations generally, and that
without any previous Reduction. By Edm.
Halley. 70
A Differtation concerning the Construction of So-
lid Problems, or Fauations of the third or
lid Problems, or Equations of the third or fourth Power, by the help of one (given) Para-
bola and a Circle. By Edmund Halley. 09
A Discourse concerning the Number of Roots, in Solid and Biquadratical Equations, as also
in Solid and Biguadratical Equations, as alfo
of the Limits of them. By E. Halley. 101
Some Illustrious Specimens of the Dottrine of
Fluxions; or Examples by which is clearly
Shewn the Use and Excellency of that Method
in solving Geometrical Problems. By Ab.
De Moivre. 128
A Method of Squaring Jame Sorts of Curves,
or Reducing them to more simple Curves. By
A. De Moivre, R.S.S. 140
Two Problems; viz. concerning the Solid of
teaf Resistance, and the Curve of Swiftelt.
Descent. Solv'd by J. Craig. 159 The Quadrature of the Logarithmical Curve.
The Quadrature of the Logarithmical Curve.
Ry I Craig. 100
A Theorem concerning the Proportion of the lime
that a heavy Body penas in accenting this
a right line tothing two given Points, to the
(fhortest) Time, in which it passes from the
one to the other of these Points, by the Arch
of a Cycloid. 168
An Extract of a Letter from the Reverend Dr.
John Wallis, to Richard Waller, Elg; Se-
cretary to the Koyal Society, concerning the
Spaces in the Cycloid, which are perfectly
Quadrable, 171
The

ţ

The CONTENTS.

The Quadrature of a Portion of the Epycloid. By Mr. Cafwell. 175 A General Proposition, shewing the Dimension of the Areas in all those kinds of Curves which are describ'd by the Equable Revolution of a Circle upon any Basis, either a Rectilineal or a Circular one. By Edm. Halley. 177 A Method of Raising an infinite Multinomial to any given Power, or Extracting any given Root By Mr. A. De Moivre. of the same. 183 A Method of Extracting the Root of an Infinite Equation. By A. De Moivre, F.R.S. 191 An Experiment of the Refraction of the Air, made at the Command of the Royal Society, March 28. 1699. By J. Lowthorp, A.M. 196 A Discourse concerning a Method of Discovering the True Moment of the Sun's Ingress into the Tropical Sines. By E. Halley. 202 A Scale of the Degrees of Heat. 2 I I The Properties of the Catenaria. By David Gregory, M. D. Savilian Professor of Astronomy, and F. R. S. 219 Of the Quadratures of Geometrically irrational Figures, By . Craig. 251 Concerning the Apparent Magnitude of the Sun and Moon, or the apparent distance of two Stars when nigh the Horizon, and when higher elevated. 263 The Sentiments of the Reverend and Learned Dr. John Wallis R. S. Soc. upon the aforefaid Appearance, communicated in a Letter to the Publisher. 275 A Demonstration of an Error committed by common Surveyors in comparing of Surveys, taken at long Intervals of Time, arising from the Variation of the Magnetick, Needle, by William Moly-

The CONTENTS.

283 Molyneux. Esq; F. R. S. A Propofal concerning the Parallax of the fixed Stars, in Reference to the Earths Annual Orb. In Several Letters of May the 2d. June 29. and July 20. 1693. from Dr. John Wallis to William Molyneux E/g; 292 Why Bodies diffolved in Menstrua Specifically lighter than themselves, swim therein? 300 Of the weight of a Cubic foot of divers grains, &c. try'd in a Veffel of well-feason'd Oak, whose concave was an exact cubic foot. By the direction of the Philosophical Society at Oxford. 308 A Letter of Dr. Wallis to Dr. Sloane, concerning the Generation of Hail, and of Thunder and Lightning, and the Effects thereof. 315 A Synoplis of the Astronomy of Comets. I A Geometrical Differtation Concerning the Rainbow in which (by a direct Method) is shewn how to find the Diameter of each Bow, the Proportion of the Refraction being given : Together with the Solution of the Inverse Problem, or how to find the Ratio of the Refraction. the Diameter of the Iris being given. By Edm-Halley, F. R. S. 25

This refers to Page 80 in this Volume.

CZ = CA -2"=da"--2da .11 1 e---+ 6g a2 ce-1- 4ga 0/14 -----3 set 10ha2 et + : e+ 20ka : e+356 a 2 - Sha et-59 -6 ka es-+71 ae6+1e

Fabella Poteftatum

Ug and by Google

ADVERTISEMENT.

There is in the Prefs and will Speedily be publish'd, A Collection of the Travels, Antiquities and Natural Histories of Countries, as they have been deliver'd in to the Royal Society; collected from the Philosophical Transactions into one Volume in Octavo.

Miscellanea Curiosa.

PART II.

A Calculation of the Credibility of Humane Testimony.

Oral Certitude Absolute, is that in which the Mind of Man entirely acquiefces, requiring no further Affurance : As if one in whom I abfolutely confide, shall bring me word of 12001. accruing to me by Gift, or a Ship's Arrival, and for which therefore I would not give the least valuable Confideration to be Enfur'd.

Moral Certitude Incompleat, has its feveral Degrees to be estimated by the Proportion it bears to the *Absolute*. As if one in whom I have that degree of Confidence, as that I would not give above One in Six to be enfur'd of the truth of what he fays, shall inform me, as above, concerning 1200 l. I may then reckon that I have as good as the Absolute Certainty of a 1000 l. or five fixths of Absolute Certainty for the whole Summ. The. The Credibility of any Reporter is to be rated (1) by bis Integrity or Fidelity; and (2) by his Ability: and a double Ability is to be confidered; both that of Apprehending what is deliver'd, and also of Retaining it afterwards, till it be transmitted.

What follows concerning the Degrees of
Credibility, is divided into Four Propositions.
The Two First, respect the Reporters of the
Narrative; as they either Transmit Successiveiy, or Attest Concurrently: the Third, the Subjest of it; as it may consist of feveral Articles:
and the Fourtb, joins those three Considerations together, exemplifying them in Oral and
in Written Tradition.

PROP. I.

Concerning the Credibility of a Report, made by Single Succeflive Reporters, who are equally Credible.

Let their Reports have, each of them, five Sixths of Certainty; and let the first Reporter give me a Certainty of 1000 l. in 1200 l. it is plain, that the Second Reporter, who delivers that Report, will give me the Certainty but of $\frac{2}{5}$ ths of that 1000 l. or the $\frac{5}{5}$ th of $\frac{2}{5}$ ths of the full Certainty of the whole 1200 l. And fo a Third Reporter, who has it from the fecond, will transmit to me but $\frac{2}{5}$ ths of that Degree of Certainty, the Second would have deliver'd me, . $\mathfrak{G}^{*}c$.

That is, if, a, be put for the Share of Affurance a fingle Reporter gives me; and, c, for that which is wanting to make that Affurance compleat; and I therefore fuppos'd to have

a of Certainty from the First Reporter; I fhall have from the Second, =; from the $a^{a^{3}}_{a+c^{3}}$.

And accordingly, if, a, be = 100, and c=6, (the number of Pounds that an 100% put out to Intereft, brings in at the Year's end;) and confequently my Share of Certainty from One Reporter, be = $\frac{1}{16}$; which is the prefent value of any Summ to be paid a Year hence : The Proportion of Certainty coming to me from a Second, will be $\frac{1}{16}$ multiplyed by $\frac{1}{16}$ (which is the prefent Value of Mony to be paid after two Years;) and that from a Third-hand Reporter, = $\frac{1}{16}$: thrice multiplied into it felf; (the Value of Mony payable at the end of Three Years) &c.

Corollary.

And therefore, as at the Rate of 6 per Cent. Interest, the present Value of any Summ payable after Twelve Years, is but half the Summ So if the Probability or proportion of Certitude transmitted by each Reporter, be $\frac{1}{1+\tau}$; the Proportion of Certainty after Twelve such Transmissions, will be but as a half; and it will grow by that Time an equal Lay, whether the Report be true or no. In the same manner, if the Proportion of Certainty be set at $\frac{1}{1+\tau}$ it will come to half from the 70th Hand: And if at $\frac{1}{1+\tau}$, from the 695th.

B 2

PROP.

3

4

PROP. II.

Concerning Concurrent Testifications.

If Two Concurrent Reporters have, each of them, as $\frac{5}{6}$ ths of Certainty; they will both give me an Affurance of $\frac{3}{5}$ ths, or of 35 to one: If Three; an Affurance of $\frac{3}{21}\frac{5}{6}$, or of 215 to one.

For if one of them gives a Certainty for 1200 l. as of $\frac{1}{2}ths$, there remains but an Aflurance of $\frac{1}{2}th$, or of 200 l. wanting to me, for the whole. And towards that the Second Attester contributes, according to his Proportion of Credibility: That is to $\frac{1}{2}ths$ of Certainty beforehand, he adds $\frac{1}{2}ths$ of the $\frac{1}{2}th$ which was wanting: So that there is now wanting but $\frac{1}{2}th$ of a $\frac{1}{2}th$, that is $\frac{1}{3}ths$; and confequently I have, from them both, $\frac{3}{3}ths$ of Certainty. So from Three, $\frac{21t}{3}$, $\mathfrak{S}^{*}c$.

That is, if the first Witness gives me $\frac{a}{a+c}$ of Certainty, and there is wanting of it $\frac{c}{a+c}$; the Second Attester will add $\frac{a}{a+c}$ of that $\frac{c}{a+c}$; and confequently leave nothing wanting but $\frac{c}{a+c}$ of that $\frac{a}{a+c} = \frac{c^2}{a+c}$. And in like manner the third Attester adds his $\frac{a}{a+c}$ of that

 c^2 and leaves wanting only $\frac{c^3}{a+c^4}$. Corol-

Corollary.

Hence it follows, that if a fingle Witnefs fhould be only fo far Credible, as to give me the Half of a full Certainty; the Second of the fame Credibility, would (joined with the firft) give me $\frac{1}{4}ths$; a Third, $\frac{7}{8}ths$, $\mathfrak{S}c$. So that the Coattestation of a Tenth, would give me $\frac{1}{10}\frac{1}{2}\frac{1}{4}ths$ of Certainty; and the Coattestation of a Twentieth, $\frac{1097}{10077000}$ or above Two Millions to one, $\mathfrak{S}c$.

PROP. III.

Concerning the Credit of a Reporter for a Particular Article of that Narrative, for the whole of which he is Credible in a certain Degree.

Let there be Six Particulars of a Narrative equally remarkable: If he to whom the Report is given, has $\frac{1}{6}ths$ of Certainty for the whole, or Summ, of them; he has 35 to one, against the Failure in any One certain Particular.

For he has Five to One, there will be no Failure at all. And if there be, he has yet another Five to One, that it falls not upon that fingle Particular of the Six. That is, he has it hs of Certainty for the whole: and of the it wanting, he has likewife it hs, or is the of the whole more; and therefore that there will be no Failure in that fingle Particular, he has it hs and is the of Certainty, or is of the

In General, if $\frac{n}{a+c}$ be the Proportion of Certainty for the whole; and $\frac{m}{m+k}$ be the chance of the reft of the particular Articles *m*, against B 3

fome one, or more of them n; there will be nothing wanting to an abfolute Certitude, against the not failing in Article, or Articles, n, but

only $\frac{nc}{m+nx_{a}+c}$

6

PROP. IV.

Concerning the Truth of either Oral or Written Tradition, (in Whole, or in Part,) Succeffively transmitted, and also Coattested by several Succeffions of Transmittents.

(1) Supposing the Transmission of an Oral and Narrative to be fo performed by a Succeffion of Single Men, or joined in Companies, as that each Transmission, after the Narrative has been kept for Twenty Years, impairs the Credit of it a th part; and that confequently at the Twelfth Hand, or at the end of 240 Years, its certainty is reduced to a Half; and there grows then an even Lay (by the Corollary of the fecond Proposition) against the Truth of the Relation : Yet if we further fuppofe, that the fame Relation is Coattested by Nine other feve ral Succeffions, transmitting alike each of them; the Credibility of it when they are all found to agree, will (by the Corollary of the first Proposition) be as 1213 of Certainty, or above a Thousand to one; and if we suppose a Coattestation of Ninetcen, the Credibility of it will be, as above Two Millions to One.

(2) In Oral Tradition as a Single Man is fubject to much Cafuality, fo a Company of Men cannot be fo eafily fuppos'd to join; and therefore the Credibility of 1_{33} ths, or about

. .

12ths, may poffibly be judged too high a Degree for an Oral Conveyance, to the distance of Twenty Years. But in Written Tradition, the Chances against the Truth or Confervation of a fingle Writing, are far lefs; and feveral Copies may also be easily supposed to concur; and those fince the Invention of Printing exactly the fame: feveral alfo diffinct Successions of fuch Copies may be as well fuppos'd, taken by different Hands, and preferv'd in different Places or Languages.

And therefore it Oral Tradition by any one Man or Company of Men might be suppos'd to be Credible, after Twenty Years at 12ths of Certainty ; or but ?-ths ; or #ths : a Written Tradition may be well imagin'd to continue, by the Joint Copies that may be taken of it for one Place, (like the feveral Copies of the fame Impression) during the space of a 100, if not 200 Years; and to be then Credible at $\frac{1}{2}$ or $\frac{1}{2}$ or $\frac{1}{2}$ the of Certainty, or at the Proportion of a Hundred to one. And then, feeing that the Succeffive Transmissions of this 12 of Certainty, will not diminish it to a Half, until it paffes the Sixty ninth Hand; (for it will be near Seventy Years before the Rebate of Mony, at that Interest, will fink it to half:) It is plain, that written Tradition, if preferv'd but by a fingle Succession of Copies, will not lose half of its full Certainty, until 70 times a Hundred (if not two Hundred) Years are paft; that is, Seven Thousand, if not Fourteen thousand Years; and further, that, if it be likewise preserv'd by Concurrent Successions of fuch Copies, its Gredibility at that Diftance may be even encreas'd, and grow far more certain from the feveral agreeing Deliveries at the end of Seventy

8

venty Succeffions, than it would be at the very first from either of the Single Hands.

(3) Laftly, in ftating the Proportions of Credibility for any Part or Parts of a Copy, it may be obferv'd; that in an Original not very long, good Odds may be laid by a careful Hand, that the Copy fhall not have fo much as a Literal Fault: But in one of greater Length, that there may be greater Odds againft any Material Error, and fuch as fhall alter the Senfe; greater yet; that the Senfe fhall not be alter'd in any confiderable Point; and ftill greater, if there be many of thofe Points, that the Error lights not upon fuch a fingle Article; as in the Third Proposition. A Letter from the Reverend Dr. Wallis, Professor of Geometry in the University of Oxford, and Fellow of the Royal Society, London, to Mr. Richard Norris, concerning the Collection of Secants; and the true Division of the Meridians in the Sea Chart.

A N old enquiry, (about the Sum or Aggregate of Secants) having been of late moved a-new; I have thought fit to trace it from its Original : with fuch folution as feems proper to it: Beginning first with the general Preparation; and then applying it to the Particular Cafe.

General Preparation.

1. Becaufe Curve lines are not fo eafily managed as Straight lines: the Ancients, when they were to confider of Figures terminated (at leaft on one fide) by a Curve line (Convex or Concave) as *AFKE*, Fig.1.2. Tab 1. did oft make use of some such expedient as this following, (but diversity varied as occasion requir'd.) Namely,

2. By Parallel Straight lines, as AF, BG, CH, &c. (at equal or unequal diftances, as there was occasion,) they parted it into fo many ny Segments as they thought fit; (or supposed it to be fo parted.)

3. These Segments were so many wanting one, as was the number of those Parallels.

4. To each of these Parallels, wanting one; they fitted Parallelograms, of such breadths as were the Intervals (equal or unequal) between each of them (respectively) and the next following. Which formed an Adscribed Figure made up of those Parallelograms.

5. And, if they began with the Greatest (and therefore neglected the least) such Figure was Circumscribed, (as Fig. 1.) and therefore Bigger than the Curvilinear proposed.

6. If with the Least (neglecting the greateft;) the Figure was Inferibed (as Fig, 2.) and therefore Less than that proposed.

7. But, as the number of Segments was increafed, (and thereby their breadths diminifhed;) the difference of the Circumfcribed from the Infcribed (and therefore of either from that proposed) did continually decrease, fo as at last to be less than any affigned.

8. On which they grounded their Method of Exhauftions.

9. In cafes wherein the Breadth of the Parallelograms, or Intervals of the Parallels, is not to be confidered, but their length only; (or, which is much the fame, where the Intervals are all the fame, and each reputed = 1.) Archimedes (inftead of Inferibed and Circumferibed Figures) used to fay, All except the Greateft, and All except the Leaft. As Prop. 11. Lin. Spiral.

Particular Cafe.

10. Though it be well known, that, in the Terrestrial Globe, all the Meridians meet at the

Tug and a Google

the Pole, (as *E P. E P*, Fig 3.) whereby the Parallels to the Equator, as they be nearer to the Pole, do continually decrease.

11. And hereby a degree of Longitude in fuch Parallels, is lefs than a degree of Longitude in the Equator, or a degree of Latitude. 12. And that, in fuch proportion, as is the Co-fine of Latitude (which is the femidiamiter of fuch Parallel,) to the Radius of the Globe, or of the Equator.

13. Yet hath it been thought fit (for fome reafons) to reprefent these Meridians, in the Sea Chart, by Parallel straight lines; as EP, EP.

14, Whereby, each Parallel to the Equator (as LA) was reprefented in the Sea-Chart, (as la,) as equal to the Equator EE: and a degree of Longitude therein, as large as in the Equator.

15. By this means, each degree of Longitude in fuch Parallels, was increafed, beyond its just proportion, at fuch rate as the Equator (or its Radius) is greater than fuch Parallel, (or the Radius thereof.)

16. But, in the Old Sea-Charts, the degrees of Latitude were yet reprefented (as they are in themfelves) equal to each other; and, to those of the Equator.

17. Hereby, amongst many other Inconveniencies, (as Mr. Edward Wright observes, in his Correction of Errors in Navigation, first publissed in the Year 1599,) the representation of Places remote from the Equator, was so distorted, in those Charts, as that (for instance) an Island in the Latitude of 60 degrees, (where the Radius of the Parallel is but half so great as that of the Equator) would have its Length (from ×

(from East to West) in comparison of its Breadth (from North to South) represented in a double proportion of what indeed it is.

18. For rectifying this in fome measure (and of fome other inconveniences)Mr.Wright advifeth; that (the Meridians remaining Parallel, as before) the degrees of Latitude, remote from the Equator, should at each Parallel, be protracted in like proportion with those of Longitude.

19. That is; As the Co-Sine of Latitude, (which is the Semi-diameter of the Parallel) to the Radius of the Globe, (which is that of the Equator:) fo fhould be a degree of Latitude, (which is every where equal to a degree of Longitude in the Equator,) to fuch a degree of Latitude fo protracted (at fuch diftance from the Equator;) and fo to be reprefented in the Chart.

20. That is, every where, in fuch proportion as is the refpective Secant (for fuch Latitude) to the Radius. For, as the Co-fine, to the Radius; fo is the Radius to the Secant (of the fame Arch or Angle;) as Fig. 4. $\Sigma \cdot R$:: R. f.

21. So that (by this means) the polition of each Parallel in the Chart, fhould be at fuch diftance from the Equator, compared with fo many Equinottial Degrees or Minutes, (as are those of Latitude,) as are all the Secants (taken at equal distances in the Arch) to fo many times the Radius.

22. Which is equivalent, (as Mr. Wright there notes) to the Projection of the Spherical furface (fuppoling the Ey at the Center) on the concave furface of a Cylinder, crected at right Angles to the Plain of the Equator.

23. And

Digitation of Google

23. And the division of Meridians, repre fented by the furface of a Cylinder erected (on the Arch of Latitude) at right Angles, to the Plain of the Meridian (or a portion thereof.) The Altitude of fuch Projection, (or portion of fuch Cylindrick furface) being (at each point of fuch Circular bafe) equal to the fecant (of Latitude) anfwering to fuch point. As Fig. 5.

24. This Projection (or portion of the Cylindrick furface) if expanded into a Plain, will be the fame with a Plain Figure, whofe bafe is equal to a Quadrantal Arch extended (or a portion thereof) on which (as ordinates) are erected Perpendiculars equal to the Secants, anfwering to the refpective points of the Arch fo- extended : The leaft of which (anfwering to the Equinoctial) is equal to the Radius; and the reft continually increasing, till (at the Pole) it be infinite. As at Fig. 6.

25. So that, as $E R \int L$. (a Figure of Secants erected at right Angles on EL, the Arch of Latitude extended,) to E R R L, (a rectangle on the fame bafe, who's altitude ER is equal to the Radius;) fo is E L (an Arch of the Equator equal to that of Latitude,) to the diftance of fuch Parallel, (in the Chart) from the Equator.

26, For finding this diffance, answering to each degree and Minute of Latitude; Mr. Wright (as the most obvious way) adds all the Secants (as they are found calculated in the Trigonometrical Canon) from the beginning, to the degree or Minute of Latitude propofed.

27. The fum of all which, except the Greateft, (anfwering to the Figure Inferibed) is too Little: The fum of all except the Leaft, (anfwering 14

fwering to the Circumfcribed,) is too Great, (which is that he follows :) And it would be nearer to the Truth than either, if (omitting all thefe) we take the intermediates; for Min. $\frac{1}{2}$, $1\frac{1}{2}$, $2\frac{1}{2}$, $3\frac{1}{2}$, $\mathcal{O}c$. or (the doubles of thefe) Min. 1, 3, 5, 7, Gc. Which yet (becaufe on the Convex fide of the Curve) would be fomewhat too Little.

28. But any of these ways are exact enough for the use intended, as creating no sensible difference in the Chart.

29. If we would be more exact ; Mr. Oughtred directs (and fo had Mr. Wright done before him) to divide the Arch into parts yet fmaller than Minutes, and calculate Secants fuiting thereunto.

30. Since the Arithmetick of Infinites introduced, and (in pursuance thereof) the Doctrine of Infinite feries (for fuch cafes as would not, without them, come to a determinate proportion;) Methods have been found for fquaring fome fuch Figures; and (particularly(the Exterior Hyperbola (in a way of continual approach) by the help of an Infinite feries. As. in the Philosophical Transactions, Numb. 38, (for the Month of August, 1668,) And my Book. De Motu, Cap. 5. Prop. 31.

31. In Imitation whereof, it hath been defired (I find) by fome, that a like Quadrature for this Figure of Secants (by an Infinite feries fitted thereunto) might be found.

32. In order to which, put we for the Radius of a Circle, R; the right Sine of an Arch or Angle, S, the Verfed Sine; V, the Co-Sine (or Sine of the Complement) $\Sigma = R - V = \sqrt{2}$: Rq-Sq: the Secant, J; the Tangent, T. Fig. 4.

Downday Google

33. Then is, $\Sigma \cdot \mathbb{R} :: \mathbb{R} \cdot f$. That is, $\Sigma \in \mathbb{R}^2$ (S= $\frac{\mathbb{R}^2}{\Sigma}$; the Secant.

34. And Σ . S :: R.T. That is, Σ) S R (T=^{SR} the Tangent.

35. Now, if we suppose the Radius C P, Fig. 7. divided into equal Parts, (and each of them $=_{\overline{ab}} R$;) and, on these, to be crected the Co-Sines of Latitude L A:

36. Then are the Sines of Latitude in Arithmetick Progression.

37. And the Secants answering thereunto, $L \int = \frac{R^2}{\Sigma}$.

38. But these Secants, (answering to right Sines in Arithmetical Progression) are not those that stand at equal distance on the Quadrantal Arch extended, Fig. 6.

39. But ftanding at unequal diftances (on the fame extended Arch;) Namely, on those points thereof, whose right Sines (whilst it was a Curve) are in *Arithmetical* Progression. As Fig. 8.

40. To find therefore the magnitude of REL, Fig. 6. Which is the fame with that of Fig. 8. (fuppofing EL of the fame length in both; however the number of Secants therein may be unequal:) we are to confider the Secants, tho' at unequal diffances: Fig. 8. to be the fame with those at equal diffances in Fig. 7. answering to Sines in Arithmetical Progreffion.

41. Now these Intervals, (or portions of the base) in Fig. 8. are the same with the intercepted Arches (or portions of the Arch) in Fig. 7. For this base is but that Arch extended.

42. And

15

16

42. And these Arches (in parts infinitely finall) are to be reputed equivalent to the portions of their respective Tangents intercepted between the fame ordinates. As in Fig. 7. 9.

43. That is, equivalent to the portions of the Tangents of Latitude.

44. And these portions of Tangents are, to the Equal intervals in the base, as the Tangent (of Latitude) to its Sine.

45. To find therefore the true Magnitude of the Parallelograms (or fegments of the Figure;) we must either protract the equal fegments of the base, Fig. 7. (in such proportion as is the respective Tangent to the Sine) to make them equal to those of Fig. 8.

46. Or elfe (which is equivalent) retaining the equal intervals of Fig. 7. protract the Secants in the fame proportion. (For, either way the Intercepted Rectangles or Parallelograms will be equally encreased) As L M Fig. 9.

47. Namely; As the Sine (of Latiude) to its Tangent; fo is the Secant to a Fourth; which is to ftand (on the Radius equally divided) inftead of that Secant.

S. $\frac{SR}{\Sigma}(:: \ge .R):: \frac{R^2}{\Sigma} \frac{R^3}{\Sigma^2 = R^2 \cdot S^2} = L M$, Fig.9.

48. Which therefore are as the Ordinates in (what I call Arith. Infin. Prop. 104) Reciproca Secundinorum: supposing \geq^{2} to be squares in the order of Secundanes.

R²

Mifeellanea Curiofa. 17

R'-S'). R'(R, $\frac{S_2}{R}, \frac{S_1}{R}, \frac{1}{R}, \frac{1}{R}$

49. This because of R3-S2R $\sum {}^{2} = R^{2} - S^{2}$, & the Sines + S R - $\frac{54}{R}$ S, in Arithmetical Pro-. -1- S4 R greffion) is reduced (by division) into this Infinite Series. $R \cdot \left\{ -\frac{S_2}{R} + \frac{S_4}{K_3} + \frac{S_6}{K_5} \right\} \&c.$ 50. That is, (putting R=1.) -1- R3

I.1-S'.1-S'-1-S', &c.

st. Then (according to the Arithmetick of Infinites) we are to interpret S, fucceffively, by 1 S, 2 S, 3 S, &c. till we come to S, the greateft, Which therefore reprefents the number of All.

52. And because the first Member doth reprefent a Series of Equals; the fecond of Secundans; the third, of Quartans, Oc. Therefore the first Member is to be multiplied by S; the fecond, by $\frac{1}{5}$ S; the third, by $\frac{1}{5}$ S; the fourth, by $\frac{1}{5}$ S; &c.

53. Which makes the Aggregate, S'-; S'-; S'-; S'-; S', &c.=ECLM, Fig. 9.

34. This (because S is always less than R = i) may be fo far continued, till fom: power of S become fo finall as that it (and all which follow it) may be fately neglected.

55. Now (to fit this to the Sea-Chart, according to Mr. Wright's defiga:) Having the propoproposed Parallel (of Latitude) given; we are to find (by the Trigonometrical Canon) the Sine of fuch Latitude; and take, equal to it, CL=S. And, by this, find the magnitude of ECLM, Fig. 9; that is, of RELf, Fig. 8. that is, of RELf, Fig. 6. And then, as R RLE (or fo many times the Radius,) to RELf (the Aggregate of all the Secants;) fo must be a like Arch of the Equator (equal to the Latitude proposed,) to the distance of fuch Parallel, (representing the Latitude in the Chart) from the Equator. Which is the thing required.

56. The fame may be obtained, in like manner, by taking the Verfed Sines in Arithmetical Progreffion. For if the right Sines (as here) beginning at the Equator, be in Arithmetical Progreffion, as 1, 2, 3, Ge. Then will the Verfed Sines, beginning at the Pole, (as being their complements to the Radius) be fo alfo.

The Collection of Tangents.

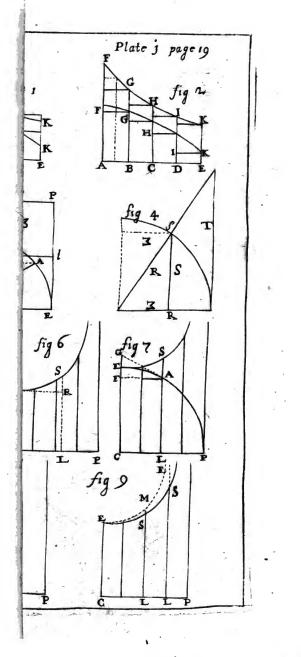
57. The fame may be applied in like manner, (though that be not the prefent businefs,) to the Aggregate of Tangents, (answering to the Arch divided into equal parts.)

58. For, those answering to the Radius fo divided, are $\frac{SR}{\Sigma}$; (taking S in Arithmetical Progreffion.)

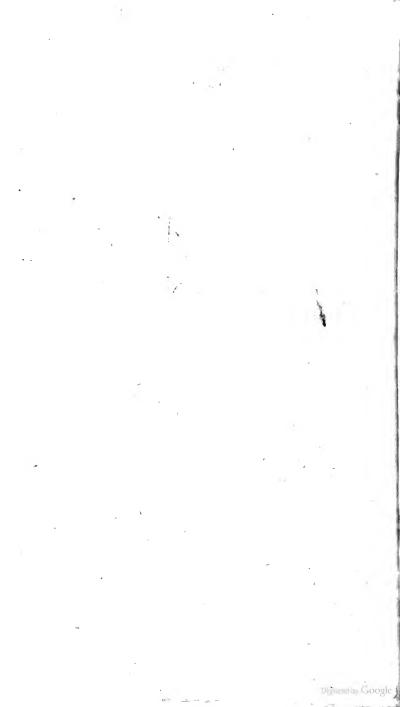
59. And then, inlarging the Bafe, (as in Fig. 8.) or the Tangent (as in Fig. 9.) in the proportion of the Tangent to the Sine.

S.

Digitized by Google



Ingineer of Google



S. $\frac{SR}{\Sigma}$ (:: $\Xi \cdot R$):: $\frac{SR}{\Xi} \cdot \frac{SR^2}{\Xi^2 = R^2 - S^2}$. 60. We have (by Division) this Series, $R^{\frac{1}{2}}S^{\frac{1}{2}}$) SR^2 (S, $+\frac{S_3}{R^2}, +\frac{S_5}{R^4}, +$ $S - \frac{S_3}{R^2} + \frac{S_5}{R^4} + \frac{S_7}{R^6} + \frac{S_9}{R^8}$ &c. $\frac{SR_2 - S_3}{S_3}$ 61. That is(putting R=1) $+\frac{S_3}{S_2} - \frac{S_5}{R^2}$ 62. Which (multiplying $+\frac{S_5}{R^2}$ 62. Which (multiplying $+\frac{S_5}{R^2}$ 62. Which (multiplying $+\frac{S_5}{R^2}$ 62. Which (multiplying $+\frac{S_5}{R^2}$ 63. Which (multiplying $+\frac{S_5}{R^2}$ 64. That is(S, $\frac{1}{2}S, \frac{1}{2}S, \frac{1}$

Which is the Aggregate of Tangents to the Arch, whole right Sine is S.

63. And this method may be a pattern for the like process in other cases of like nature. An easse Demonstration of the Analogy of the Logarithmick Tangents to the Meridian Line or sum of the Secants; with various Methods for computing the same to the utmost Exactness, by E. Halley.

T is now near 100 Years fince our Worthy Countryman, Mr. Edward Wright, published his Correction of Errors in Navigation, a Book well deferving the perufal of all fuch as defign to use the Sea. Therein he confiders the Course of a Ship on the Globe, ftearing obliquely to the Meridian; and having fhewn, that the Departure from the Meridian, is in all cafes lefs than the Difference of Longitude, in the ratio of Radius to the secant of the Latitude, he concludes, That the fum of the Secants of each point of the Quadrant being added fucceffively would exhibit a line divided into Spaces, fuch as the intervals of the parallels of Latitude ought to be in a trueSea-Chart, whereon theMeridians are made parallel Lines, and the Rhombs or Oblique Courfes reprefented by right Lines. This is commonly known by the name of the Meridian Line, which tho' it generally be called Mercator's, was yet undoubtedly Mr. Wright's Invention, (as he has made it appear in his Preface.) And the Table thereof is to be met with in moftBooks treating of Navigation, computed with fufficient exactness for the purpose. It

Digitated by Google

21

It was first discovered by Chance, and as far as I can learn, first publish'd by Mr. Henry Bond, as an addition to Norwood's Epitome of Navigation, about 50 Years fince, that the Me- . ridian Line was analogous to a Scale of Logarithmick Tangents of half the Complements of the Latitudes. The difficulty to prove the truth of this Proposition, feemed such to Mr. Mercator, the Author of Logarithmotechnia, that he proposed to wager a good fum of Mony, against . whofo would fairly undertake it, that he fould not demonstrate either, that it was true or falfe: And about that time Mr. John Collins, holding a Correspondence with all the Eminent Mathematicians of the Age, did excite them . to this enquiry

The first that demonstrated the faid Analogy, was the excellent Mr. James Gregory in his Exertitationes Geometrica, published Anno 1668. which he did, not without a long train of Confequences and Complication of Proportions, whereby the evidence of the Demonstration is . in a great measure loft, and the Reader wearied before he attain it. Nor with lefs work and apparatus hath the celebrated Dr. Barrow, . in his Geometrical Lectures (Lect. XI. App. 1.) proved, that the Sum of all the Segants of any arch is analogous to the Logarithm of the ratio of Radius + Sine to Rad. - Sine, or, which, is all one, that the Meridional parts answering, to any degree of Latitude, are as the Logarithms of the rationes of the Versed Sines of the diftances from both the Poles. . Since which the incomparable Dr. Wallis (on occasion of a Paralogifm committed by one Mr. Tvarris in this matter) has more fully and clearly handled this Argument, as may be feen in Num 176. of. the C 3

the Transactions. But neither Dr. Wallis, nor Dr. Barrow, in their faid Treatifes, have any where touched upon the aforefaid relation of the Meridian-line to the Logarithmick Tangent; nor hath any one, that I know of, yet discovered the Rule for computing independently the interval of the Meridional parts answering to any two given Latitudes.

Wherefore having attained, as I conceive, a very facile and natural demonstration of the faid Analogy, and having found out the Rule for exhibiting the difference of Meredional parts, between any two parallels of Latitude, without finding both the Numbers whereof they are the difference: I hope I may be entituled to a fhare in the Improvements of this useful part of Geometry. And first, let us demonstrate the following Proposition.

The Meridian Line is a Scale of Logarithmick. Tangents of the half Complements of the Latitudes.

For this Demonstration, it is requisite to premise these four Lemmata.

Lemma. I. In the Stereographick Projection of the Sphere upon the plain of the Equinoctial, the diftances from the Center, which in this cafe is the Pole, are laid down by the Tangents of half those diftances, that is, of half the Complements of the Latitudes. This is evident from Eucl. 3. 20.

Lem. II. In the Stereographick Projection; the Angles under which the Circles interfect each other, are in all cafes equal to the Spherical Angles they represent: Which is perhaps as valuable a property of this Projection, as that of all the Circles of the Sphere thereon appearing

22

ing Circles: But this not being vulgarly known, must not be assumed without a Demonstration.

Let EBPL in Fig. 1. Tab. 2. be any great circle of the Sphere, E the Eye placed in its Circumference, C its Center, P any point thereof, and let FCO be supposed a plain erected at right Angles to the Circle EBPL, on which FCO we defign the Sphere to be projected. Draw E P croffing the Plain F CO in p, and p shall be the Point P projected. To the point P draw the Tangent A P G and on any point thereof, as A, erect a perpendicular AD, at right angles to the plane EBPL, and draw the lines PD, AC, DC: and the Angle APD shall be equal to the Spherical Angle contained between the plains AP C, DPC. Draw alfo AE, DE, interfecting the plain FCO in the points a and d; and joyn ad, pd: I fay the Triangle ad p is fimu-lar to the triangle ADP. And the Angle apd equal to the Augle APD. Draw PL, AK, parallel to FO, and by reason of the parallels, a p will be to a d as AK to AD: But (by Eucl. 3. 32.) in the triangle AKP, the angle AKP= LPE is alfo equal to APK=EPG, wherefore the fides AK, AP, are equal, and 'twill be as ap to ad fo AP to AD. Whence the angles DAP, dap being right, the angle APD will be equal to the angle apd; that is, the Spherical Angle is equal to that on the Projection, and that in all Cafes. Which was to be proved.

This Lemma I lately received from Mr. Ab. de Moivre, though I fince understand from Dr. Hook, that he long ago produced the fame thing before the Society. However the demonstration and the rest of the Discourse, is my own.

C4

24

Lemma III. On the Globe, the Rumb Lines make equal angles with every Meridian, and by the aforegoing Lemma, they must likewife make equal angles with the Meridians in the Stereographick Projection on the plain of the Fquator': They are therefore, in that Projection, Proportional Spirals about the Pole Point.

Lem. IV. In the Proportional Spiral (Fig. 2.) it is a known property, that the angles BPC, or the arches BD, are Exponents of the rationes of BP to PC: for if the arch BD be divided into innumerable equal parts, right lines drawn fromthem to the Center P, shall divide the Curve B. ccC, into an infinity of proportionals; and all the lines Pc shall be an infinity of proportionals between FB and PC; whole number is equal to all the points d, d, in the arch BD : Whence and by what I have deliver'd'in the next enfuing Difcourfe it follows, that as BD to Bd, or as the angle BPC to the angle BPc, fo is the Logarithm of the ratio of PB to PC, to the Logarithm of the ratio of PB to Pc. 1 2 2 2 12. 27

From these Lemmara our Proposition is very clearly demonstrated 'For by the first, PB; Pc, PC are the Tangents of half the Complements of the Latitudes in the Stereographick Projection: and by the last of them, the differences of Longitude, or angles at the Pole between them, are Logarithms of the values of those Tangents one to the other. But the Nautical Meridian Line, is no other than a Table of the Longitudes, answering to each minute of Latitude, on the Rhumb-line, making an angle of 45 degrees with the Meridian Wherefore the Meridian Line is no other than a Scale of Logarithmick Tangents of the half-Complements

ments of the Latitudes. Quod erat demonstrandum.

Coroll. 1. Becaufe that in every point of any Rhum Line, the difference of Latitude is to the Departure, as the Kadius to the Tangent of the angle that Rhumb makes makes with the Meridian; and those equal Departures are every where to the differences of Longitude, as the Radius to the Secant of the Latitude; it follows, that the differences of Longitude are, on any Rhumb, Logarithms of the fame Tangents, but of a differing Species; being proportioned to one another as are the Tangents of the angles made with the Meridian.

Coroll. 2. Hence any Scale of Logarithm Tangents, (as those of the Vulgar Tables made after Briggs's form; or those made to Napier's, or any other form whatfoever) is a Table of the differences of Longitude, to the feveral Latitudes, upon fome determinate Rhumb or other: And therefore, as the Tangent of the angle of fuch Rhumb, to the Tangent of any other Rhumb: So the difference of the Logarithms of any two Tangents, to the difference of Longitude, on the proposed Rhumb, intercepted between the two Latitudes, of whose half Complements you took the Logarithm Tangents:

- And fince we have a very compleat Table of Logarithm Tangents of Briggs's form, published by Vlacq, Anna 1633, in his Canon Magnus. Triangulorum Logarithmicus, computed to ten Decimal places of the Logarithm, and to every ten Seconds of the Quadrant (which feems to be more than fufficient for the niceft Calculator) I thought fit to enquire the Oblique angle, with which that Rhumb Line croffes the Metridian, 26

ridian, whereon the faid Canon of Valcq precifely answers to the differences of Longitude, putting Unity for one minute thereof, as in the Common Meridian Line. Now, the momentary augment or fluxion of the Tangent Line at 45 degrees, is exactly double to the fluxion of the arch of the Circle, (as may eafily be proved) and the Tangent of 45 being equal to Radius, the fluxion also of the Logarithm Tangent will be double to that of the arch, if the Logarithm be of Napier's form : but for Briggs's form, it will be as the fame doubled arch, multiplied into 0, 43429, &c. or divided by 2, 30258, &c. Yet this must be understood only of the addition of an indivisible arch, for it ceafes to be true, if the arch have any determinate magnitude.

Hence it appears, that if one minute be fuppofed Unity, the length of the arch of one minute being ,000290888208665721596154, &c. in parts of the Radius, the proportion will be as Unity to 2,908882, &c. fo Radius to the Tangent of 71° 1' 42'' whofe Logarithm is 10. 46372611720718325204, &c. and under that angle is the Meridian interfected by that Rhumb Line, on which the differences of Napier's Logarithm Tangents of the half Complements of the Latitudes are the true differences of Longitude, cflimated in minutes and parts, taking the first Four Figures for Integers. But for *Vlacq*'s Tables, we must fay.

As .2302585, &c. to 2908882, &c. So Radius to 1,26331143874244569212, &c. which is the Tangent of 51° 38' 9", and its Logarithm 10,101510428507720941162, &c. wherefore in the Rhumb Line, which makes an angle of 51° 38' 9" with the Meridian, *Vlacq*'s Logarithm Tan-

Dationally Google

Tangents are the true differences of Longitude. And this compared with our fecond *Corollary* may fuffice for the use of the Tables already computed.

27

But if a Table of Logarithm Tangents be made by extraction of the root of the Infiniteth power, whofe Index is the length of the arch you put for Unity, (as for minutes the ,000 2908882th,&c. power) which we will call a; fuch a Scale of Tangents fhall be the true Meridian Line, or fum of all the Secants taken infinitely many. Here the Reader is defired to have recourfe to my little Treatife of *Logarithms*, in the enfuing Difcourfe that I may not need to repeat it. By what is there delivered, it will follow, that putting t for the excefs or defect of any Tangent above or under the *Radius* or *Tangent* of 45; the Logarithm of the ratio of *Radius* to fuch Tangent will be

into t-itt -- itt- ittt- ittt -- it5,&c.

when the arch is greater than 45^{gr}, or

m intor + itt + 1t3 + it + + - it', &c.

when it is lefs than 45gr. And by the fame doftrine putting T for the Tangent of any arch, and r for the difference thereof from the Tangent of another arch, the Logarithm of their ratio will be

 $\int_{m}^{1} \operatorname{into} \frac{t}{T} + \frac{tt}{2TT} + \frac{t^{3}}{3T^{3}} + \frac{t^{3}}{4T^{4}} + \frac{t}{5T^{7}} \&c.$ when T is the greater Term, or

 $\frac{1}{m}$ into $\frac{t}{T} - \frac{tt}{2TT} + \frac{t^3}{3T_3} + \frac{t_4}{4T_7} + \frac{t_7}{5T_7}$, &c

when T is the leffer Term :

And if m be supposed ,0002908882, &c. =

28

a, its reciprocal $\frac{r}{a}$ will be, 3437,74677078493 02526, &c. which multiplied into the aforefaid Series, fhall give precifely the difference of Meridional parts, between the two Latitudes, to whofe half complements the affumed Tangents belong. Nor is it material from whether Pole you estimate the Complements, whether the clevated or depressed; the Tangents being to one another in the fame ratio as their Complements, but inverted.

In the fame Difcourfe I alfo fhewed, that the Series might be made to converge twice as fwift, all the even powers being omitted : and putting τ for the fum of the two Tangents, the fame Logarithm would be

 $\frac{2}{m}$ or $\frac{2r}{a}$ into $\frac{t}{\tau} + \frac{t}{3\tau^3} + \frac{t}{5\tau^4} + \frac{t}{7\tau^7} + \frac{t^9}{9\tau^9} \&c.$

but the ratio of τ to t, or of the fum of two Tangents to their difference, is the fame as that of the fine of the fum of the arches, to the fine of their difference. Wherefore, if S be put for the fine Complement of the Middle Latitude, and s for the fine of half the difference of Latitudes, the fame Series will be

 $\frac{2r}{a}$ into $\frac{s}{S} + \frac{s^3}{3S^3} + \frac{s^7}{5S^7} \frac{s^7}{7S^7} + \frac{s^9}{9S^9}$ &c.

wherein, as the differences of Latitude are fmaller, fewer fteps will fuffice. And if the Equator be put for the middle Latitude, and confequently S = R, and s to the *fine* of the Latitude, the Meridional parts reckoned from the Equator will be

which is coincident with Dr. Wallis's folution in Numb. 176. of the Philosophical Transactions. And this fame Series being half the Logarithm of the ratio of R-1-s to R--s, that is, of the Verfed-fines of the distances from both Poles, does agree with what Dr. Barrow had shewn in his XIth. Lecture.

The fame ratio of τ to t may be expressed also by that of the Sum of the Co-fines of the two Latitudes, to the fine of their difference : As likewise by that of the Sine of the Sum of the two Latitudes, to the difference of their Co-fines : Or by that of the Versed-fine of the Sum of the Co-latitudes, to the difference of the fines of the Latitudes: Or as the fame difference of the fines of the Latitudes, to the Versed-fine of the difference of the Latitudes; to the Versed-fine of the difference of the Lattudes; all which are in the fame ratio of the Co-fine of the middle Latitude, to the Sine of half the difference of the Latitudes. As it were case to demonstrate, if the Reader were not supposed capable to do it himself, upon a bare inspection of a Scheme duly representing these Lines.

This variety of Expression of the fame ratio I thought not fit to be omitted, because by help of the rationality of the Sine of 30gr. in all cases where the Sum or difference of the Latitudes is 30gr. 60gr. 90gr. 120gr. or 150 degrees, some one of them will exhibit a simple feries, wherein great part of the Labour will be faved : And besides I am willing to give the Reader his choice which of these equippolent methods to make use of; but for his exercise shall

Milcellanea Curiofa.

20

fhall leave the profecution of them, and the compendia arifing therefrom, to his own Industry. Contenting my felf to confider only theformer, which for all uses feems the most convenient, whether we defign to make the whole Meridian Line, or any part thereof, viz.

 $\frac{2r}{a}$ into $\frac{s}{S} + \frac{s^{3}}{3S^{3}} + \frac{s^{3}}{5S^{3}} + \frac{s^{7}}{7S^{7}} + \frac{s^{9}}{9S^{9}} \&c.$

Wherein *a* is the length of any Arch which you defign fhall be the Integer or Unity in your Meridional Parts, (whether it be a Minute, League, or Degree, or any other,) S the Cofine of the Middle Latitude, and s the Sine of half the difference of Latitudes; but the Secants being the Reciprocals of the Co-fines,

 $\frac{s}{S}$ will be equal to $\frac{fs}{rr}$ putting f for the Secant of the Middle Latitude; and $\frac{2r}{a}$ into $\frac{s}{S}$ will be $= \frac{2fs}{ar}$ This multiplied by $\frac{ss}{3SS}$ that is by $\frac{ffss}{3rrrr}$, will give the fecond ftep: and that again by $\frac{3ffss}{5rrrr}$, the third ftep; and so forward, till you have compleated as many places as you defire. But the fquares of the Simes being in the fame ratio with the Verfed-fines of the double Arches, we may inftead of $\frac{ss}{3SS}$ affume for our Multiplicator $\frac{v}{3V}$, or the Verfed-fine of the

the difference of the Latitudes, divided by thrice the Verfed-fine of the fum of the Co-latitudes, &c. which is the utmost *Compendium* 1 can think of for this purpose, and the fame *feries* will become,

 $\frac{2sr}{aS} into 1 + \frac{v}{3V} + \frac{v^2}{5V^2} + \frac{v^3}{7V_1} + \frac{v^4}{9V^2}$

Hereby we are enabled to estimate the default of the method of making the Meridian line, by the continued addition of the Secants of æquidifferent Arches, which as the difference of those Arches are finaller, does still nearer and nearer approach the Truth. If we allume, as Mr. Wright did, the Arch of one minute to be Unity, and one minute to be the common difference of a rank of Arches : It will be in all cafes, as the Arch of one Minute, to its Chord :: So the Secant of the middle Latitude, to the first step of our series. This by reason of the near equality between a and 2 s, which are to one another in the ratio of Unity to 1-0, 0000000352566457713, &c. will not differ from the Secant f but in the ninth Figure; being lefs than it in that proportion. The next ftep being $+\frac{2\int_{3}^{3}s^{3}}{ar^{5}}$ will be equal to the Cube of the Secant of the middle Latitude multipli ed into $\frac{2555}{3\,arr}$ = 0,00000000705132908715; which therefore unless the Secant exceed ren times Radius, can never amount to 1 in the fifth place. These two steps suffice to make the Meridian Line, or Logarithm Tangent to far more places than any Tables of Natural Secafes.

cants yet extant, are computed to; but if the third ftep be required, it will be found to be f^{s} into, $\frac{2s^{s}}{5ar^{4}} = 0,000000000000089498;$ By all which it appears, that Mr. Wright's Table does no where exceed the true Meridian Parts by fully half a Minute: which finall difference arifes by his having added continually the Secants of 1', 2', 3', &c. inftead of 01', 11, 2'1, 3'1, &c. But as it is, it is abundantly fufficient for Nautical Vles. That in Sir Jonas Moor's New System of the Mathematicks, is much nearer the Truth, but the difference from Wright is fcarce fenfible till you exceed those Latitudes where Navigation ceafes to be practicable, the one exceeding the Truth by about half a MInute, the other being a very finall matter deficient therefrom.

For an Example easie to be imitated by whofo pleases, I have added the true Meridional Parts to the first and last Minutes of the Quadrant; not so much that there is any occasion for such occurrancy, as to shew that I have obtained, and laid down herein, the full Doctrine of these spiral Rhumbs, which are of so great concern in the Art of Navigation.

The first Minute is, 1.00000001410265862178 The Second, 2,00000005641063806707 The Last, or 89° 59 is 30374,9634311414228643

and not 32348, 5279 as Mr. Wright has it; by adding the Secants of every whole Minute: Nor 30249,8 as Mr. Oughtred's Rule makes it, by adding the Secants of every other half Minute. Nor 30364,3 as Sir Jonas Moor had concluded it

Tug and by Google

it by I know not what Method, tho' in the reft of his Table he follows Oughtred.

32

And this may fuffice to fhew how to derive the true Meridian Line from the Sines, Tangents, or Secants fuppofed ready made; but we are not defitute of a Method for deducing the fame independently, from the Arch it felf. If the Latitude from the Equator be estimated by the length of its Arch *A*, Radius being Unity, and the Arch put for an *Integer* be *a*, as before; the Meridional parts answering to that Latitude, will be

 $\frac{1}{a} \operatorname{into} A \cdot |_{-\frac{1}{6}} A^{3} \cdot |_{-\frac{1}{2}} A^{5} + \frac{1}{70} A^{7} \operatorname{or} \frac{5}{70} \frac{4}{40} A^{7} + \frac{2}{1} \frac{1}{77} A^{7} \operatorname{or} \frac{1}{30} \frac{1}{40} A^{7} + \frac{1}{1} \frac{1}{177} A^{5} \operatorname{or} \frac{1}{30} \frac{1}{10} \frac{1}{10} A^{5} + \frac{1}{10} \frac{1}{10} A^{5} + \frac{1}{10} \frac{1}{10} A^{5} + \frac{1}{10} \frac{1}{10} \frac{1}{10} \frac{1}{10} A^{5} + \frac{1}{10} \frac{1}{1$

which converges much fwifter than any of the former Series, and befides has the advantage of A encreasing in Arithmetical progression, which would be of great ease, if any should undertake de novo to make the Logarithm Tangents, or the Meridian Line to many more places than now we have them. The Logarithm Tangent to the Arch of $45 + \frac{1}{2}A$ being no other than the aforefaid Series $A + \frac{1}{6}A^3 + \frac{1}{27}A^5$, &c. in Napeir's form, or the fame multiplied into 0,43429, &c. for Brigg's.

But becaufe all thefe Series toward the latter end of the Quadrant do converge exceeding flowly, fo as to render this Method almoft ufelefs, or at leaft very tedious: It will be convenient to apply fome other Arts, by affuming the Secants of fome intermediate Latitudes; and you may for s or the Sine of α the Arch of half the difference of Latitudes, fublitute $\alpha - \frac{1}{2}\alpha^3 + \frac{1}{112}\alpha^{45} + \frac{1}{2}\alpha^{47} + \frac{1}{2}\alpha^{47}\alpha^{47}$, &c. according

34

ing to Mr. Newton's Rule for giving the Sine from the Arch: And if a be no more than a Degree, a very few steps will fuffice for all the accuracy that can be defired.

And if α be commenfurable to α , that is, if it be a certain number of those Arches with which you make your *Integer*, then will $\frac{\alpha}{\alpha}$ be that number: which if we call n, the parts of the Meridional Line will be found to be,

f ⁿ into 4	$\begin{bmatrix} 1 & + \frac{3}{3r^{4}} & \frac{5}{5r^{8}} & \frac{1}{7r^{12}} \\ -\frac{\alpha a}{6rr} & \frac{3}{6r^{6}} & -\frac{5}{6r^{6}} \\ -\frac{1}{20r^{4}} & -\frac{1}{360r^{8}} \\ -\frac{1}{360r^{8}} & \frac{1}{360r^{8}} \\ -\frac{1}{300r^{4}} & \frac{1}{300r^{8}} \\ -\frac{1}{5040r^{8}} \\ -\frac{1}{5040r^{8}$
1	5040 r *

In this, the first two steps are generally sufficient for Nautical uses, especially when neither of the Latitudes exceed 60 degrees, and the difference of Latitudes doth not pass 30 degrees.

But I am fenfible I have already faid too much for the Learned, tho' too little for the Learner; to fuch I can recommend no better Treatife, than Dr. Wallis's precedent Difcourfe, wherein he has with his ufual brevity, and that perspecuity peculiar to himself, handled this Subject from the first Principles, which here for the most part we suppose known.

I need not fhew how, by regreffive work, to find the Latitudes from the Meridional Parts, the Method being fufficiently obvious. I fhall only conclude with the proposal of a Problem which

which remains to make this Doctrine compleat, and that is this.

A Ship fails from a given Latitude, and having run a certain number of Leagues, has alter'd her Longitude by a given angle, it is required to find the Courfe fteared. The folution hereof would be very acceptable, if not to the Publick, at least to the Author of this Tract, being likely to open fome further Light into the Mysteries of Geometry.

To conclude, I shall only add, That Unity being Radius; the Co-fine of the Arch A, according to the fame Rules of Mr. Newton, will be

1-- 1 A2-1- 1 A1-- 1 A6-1- 1 A8-- 1 2855 A10 &c.

from which and the former Series exhibiting the Sine by the Arch, by division, it is easie to conclude, that the Natural Tangent of the Arch A, is

 $A + \frac{1}{3} A^3 + \frac{2}{13} A^5 + \frac{17}{315} A^7 + \frac{25}{315} A^9$, &c.

and the Natural Secant to the fame Arch

1+1 A2-+ 1 A++ 61 A6+ 277 A8, &c.

and from the Arithmetick of Infinites, the Number of these Secants being the Arch A, it follows, that the sum Total of all the Infinite Secants on that Arch, is

A+ A'+ A'+ 1 A'+ 1 AT+ 1 AT+ 1777 A', &c.

D 2

e 19. 1 1

26

the which, by what foregoes, is the Logarithm Tangent of Napeir's form, for the Arch of 45gr. +1 A, as before.

And Collecting the Infinite Sum of all the Natural Tangents on the faid Arch A, there will arife

which will be found to be the Logaritm of the Secant of the fame Arch A.

A most compendious and facile Method for Constructing the Logarithms, exemplified and demonstrated from the Nature of Numbers, without any regard to the Hyperbola, with a speedy Method for finding the Number from the Logarithm given. By E. Halley.

HE Invention of the Logarithms is justly efteemed one of the most Useful Difcoveries in the Art of Numbers, and accordingly has had an Universal Reception and Applause; and the great Geometricians of this Age, have not been wanting to cultivate this Subject with all the Accuracy and Subtilty a matter of that confequence doth require; and they have demonstrated feveral very admirable Properties of these Artificial Numbers, which have rendred their Construction much more facile than by those operose Methods at first used by their truly Noble Inventor, the Lord Napeir, and our worthy Country-man Mr. Briggs.

But notwithstanding all their Eedeavours, I find very few of those who make constant use of Logarithms, to have attained an adequate Notion of them; to know how to make or examine them; or to understand the extent of D_3 the

use of them · Contenting themselves with the Tables of them as they find them, without daring to question them, or caring to know how to rectifie them, should they be found amiss; being I suppose under the apprehension of some great difficulty therein. For the fake of such the following Tract is principally intended, but not without hopes however to produce something that may be acceptable to the most knowing in these matters.

But first, it may be requisite to premise a definition of Logarithms, in order to render the enfuing Difcourfe more clear, the rather because the old one Numerorum proportionalium aqui differentes comites, feems too fcanty to define them fully. They may more properly be faid to be Numeri Rationum Exponentes : Wherein we confider ratio as a Quantitas sui generis, beginning from the ratio of equality, or 1 to 1=0; being Affirmative when the ratio is increafing, as of Unity to a greater Number, but Negative when decreasing; and these rationes we suppose to be measured by the Number of ratiuncula contained in each. Now thefe ratiuncula are fo to be understood as in a continued Scale of Proportionals infinite in Number between the two terms of the ratio, which infinite Number of mean Proportionals is to that infinite Number of the like and equal ratiuncula between any other two terms, as the Logarithm of the one ratio is to the Logarithm of the other. Thus, if there be suppofed hetween 1 and 10 an infinite Scale of mean Proportionals, whofe Number is 100000, &c. in infinitum; between 1 and 2 there shall be 30102, Oc. of fuch Proportionals, and between

I

Dig and Google

29

1 and 3 there will be 47712 &c. of them; which Numbers therefore are the Logarithms of the rationes of 1 to 10, 1 to 2, and 1 to 3; and not fo properly to be called the Logarithms of 10; 2 and 3.

But if inftead of fuppofing the Logarithms composed of a number of equal Ratiuncule, proportional to each ratio, we shall take the ratio of Unity to any number to confift always of the fame infinite number of Ratiuncula, their magnitude, in this cafe, will be as their number in the former; wherefore if between Unity and any Number proposed, there be taken any infinity of mean Proportionals, the infinitely little augment or decre-ment of the first of those means from Unity, will be a ratiuncula, that is, the momentum or Fluxion of the ratio of Unity to the faid Number: And feeing that in these continual Proportionals all theratiuncula are equal, their Sum, or the whole ratio will be as the faid momentum is directly; that is, the Logarithm of each ratio will be as the Fluxion thereof. Wherefore if the Root of any infinite Power be extracted out of any Number, the differentiola of the faid Root from Unity, shall be as the Logarithm of that Number. So that Logarithms thus produced may be of as many forms as you please to assume infinite Indices of the Power whole Root you feek : as if the Index be fuppofed 100000 c. infinitely, the Roots shall be the Logarithms invented by the Lord Napeir ; but if the faid Index were 2302585, Oc. Mr. Briggs's Logarithms would immediately be produced. And if you please to stop at any number of Figures, and not to continue them on, it will D4 fuffice

fuffice to affume an *Index* of a Figure or two more than your intended Logarithm is to have, as Mr. *Briggs* did, who to have his Logarithms true to 14 places, by continual extraction of the Square Root, at last came to have the Root of the 140737488355328th Power; but how operofe that Extraction was, will be easily judged by whofo shall undertake to examine his *Calculus*.

Now, though the Notion of an Infinite Power may feem very ftrange, and to those that know the difficulty of the Extraction of the Roots of High Powers, perhaps impracticable; yet by the help of that admirable Invention of Mr. Newton, whereby he determines the Uncia or Numbers prefix'd to the Members compofing Powers (on which chiefly depends the Do-Arine of Series) the Infinity of the Index contributes to render the Expression much more eafie: For if the Infinite Power to be refolved be put (after Mr. Newton's Method) $\overline{p+pq}, \overline{p+pq}$ in or $\overline{1+q|m}$, inftead of $1+\frac{1}{m}q+\frac{1}{m}$ $+\frac{1-m}{2mm} qq + \frac{1-3m+2mm}{6m^3} q^3 + \frac{1-6m+11mm-6m^3}{24m^4} q^4$ &c. (which is the Root when m is finite) becomes $1 + \frac{1}{m}q - \frac{1}{2m}qq + \frac{1}{3m}q^3 + \frac{1}{4m}q^4 + \frac{1}{5m}q^5$, &c. m m being infinite infinite, and confequently whatever is divided thereby vanishing. Hence it follows that $\frac{1}{m}$ multiplied into $q - \frac{1}{2} qq + \frac{1}{3}$ $qqq - \frac{1}{3}q^4 - \frac{1}{3}q^5$ &c. is the augment of the first

of our mean Proportionals between Unity and 1+q, and is therefore the Logarithm of the ratio of 1 to 1+q; and whereas the Infinite Index

dex m may be taken at pleasure, the feveral Scales of Logarithms to fuch Indices will be as $\frac{1}{m}$ or reciprocally as the Indices. And if the Index be taken 10000, &c. as in the cafe of Napeir's Logarithms, they will be fimply $q_{-\frac{1}{2}}$ $qq+\frac{1}{3}qq-\frac{1}{4}q^4+\frac{1}{3}q^5-\frac{1}{4}q^5$ &c. Again, if the Logarithm of a decreafing ratio be fought, the infinite Root of 1-9 or $\frac{1}{1-q!} = \frac{1}{m} \frac{1}{m}$ $q^{\pm} - \frac{1}{\delta m} q^{\delta}$ &c. whence the decrement of the first of our infinite Number of Proportionals will be $\frac{1}{m}$ into $q + \frac{1}{2} q g + \frac{1}{2} q^3 + \frac{1}{4} q^4 + \frac{1}{2} q^5 + \frac{1}{4} q^6$ &c. which therefore will be as the Logarithm of the ratio of Unity to 1-q. But if m be put 10000, &c. then the faid Logarithm will be 9-1-199+193+191+191+196, &c. Hence the terms of any ratio, being a and b. g becomes $\frac{b-a}{a}$ or the difference divided by the leffer term, when 'tis an increasing ratio; or $\frac{b}{b}$ when 'tis decreasing, or as b to a. Whence the Logarithm of the fame ratio may be doubly express, for putting x for the difference of the terms a and b, it will be either $\frac{1}{m} \operatorname{into} \frac{x}{b} + \frac{x^2}{2bb} + \frac{x^3}{2bb} + \frac{x^4}{4bb} + \frac{x^7}{5bb} + \frac{x^5}{5bb} \& c. \text{ or }$

 $\frac{1}{m} into \frac{x}{a} - \frac{x^2}{2 a a} + \frac{x^3}{3 a^3} - \frac{x^4}{4 a^4} + \frac{x^7}{5 a^5} - \frac{x^6}{6 a^5} & \text{C.}$ But

But if the ratio of a to b be fuppofed divided into two parts, viz. into the ratio of a to the Arithmetical Mean between the terms, and the ratio of the faid Arithmetical Mean to the other term b, then will the Sum of the Logarithms of those two rationes be the Logarithm of the ratio of a to b; and substituting $\frac{1}{2}z$ instead of $\frac{1}{2}a + \frac{1}{2}b$ the faid Arithmetical Mean, the Logarithms of those rationes will be by the foregoing Rule,

 $\frac{1}{m}\frac{x}{7} + \frac{xx}{277} + \frac{x^7}{37^2} + \frac{x^4}{471} + \frac{x^5}{575} + \frac{x^6}{67^6}$ &c. and

 $\frac{1}{m} \ln \frac{x}{2} - \frac{xx}{237} + \frac{x^{7}}{37^{3}} - \frac{x^{4}}{47^{4}} + \frac{x^{7}}{575} - \frac{x^{6}}{67^{6}} \&C.$

the Sum 1/m in $\frac{2x}{7} * + \frac{2x^3}{37^3} * + \frac{2x^5}{57^3} * \frac{2x^7}{77^7}$ &c. will

be the Logarithm of the ratio of a to b, whofe difference is x and Sum z. And this Series converges twice as fwift as the former, and therefore is more proper for the Practice of making Logarithms: Which it performs with that expedition, that where x the difference is but the hundredth part of the Sum, the first step $\frac{2x}{z}$ fuffices to seven places of the Logarithm, and the second step to twelve: But if Briggs's first Twenty Chiliads of Logarithms be supposed made, as he has very carefully computed them to fourteen places, the first step alone, is capable to give the Logarithm of any intermediate Number true to all the places of those Tables.

After the fame manner may the difference of the faid two Logarithms be very fitly applied

Un and Google

43

plied to find the Logarithms of Prime Numbers, having the Logarithms of the two next Numbers above and below them: For the difference of the ratio of a to $\frac{1}{2}$ z and of $\frac{1}{3}$ z to b is the ratio of a b to $\frac{1}{2}$ zz, and the half of that ratio, is that of \sqrt{ab} to $\frac{1}{2}$ z, or of the Geometrical Mean to the Arithmetical. And confequently the Logarithm thereof will be the half difference of the Logarithms of those rationes, viz.

$$\frac{1}{m} \text{ into } \frac{xx}{2zz} + \frac{x_1}{4z^4} + \frac{x^5}{6z^6} + \frac{x^8}{8z^8} \&c.$$

Which is a Theorem of good difpatch to find the Logarithm of $\frac{1}{2}z$. But the fame is yet much more advantageoufly performed by a Rule derived from the foregoing, and beyond which, in my Opinion, nothing better can be hoped. For the ratio of $a b \tan \frac{1}{2} z \operatorname{or} \frac{1}{4} a a - \frac{1}{2} a b + \frac{1}{4}$ b b, has the difference of its terms $\frac{1}{4} a a - \frac{1}{2} a b + \frac{1}{4}$ b b, or the Square of $\frac{1}{2} a - \frac{1}{2} b = \frac{1}{4} x x$, which in the prefent cafe of finding the Logarithms of Prime Numbers, is always Unity, and calling the Sum of the terms $\frac{1}{4} z z - |-a b = -\frac{1}{2}b$ or $\frac{1}{2}$ z will be found

 $\frac{1}{m} \frac{1}{yy} + \frac{1}{3y^6} + \frac{1}{5y^{10}} + \frac{1}{7y^{1+}} + \frac{1}{9y^{18}} \&c.$

which converges very much faster than any Theorem hitherto published for this purpose. Here note $\frac{1}{m}$ is all along applied to adapt these Rules

Rules to all forts of Logarithms. If *m* be 10000 &c. it may be neglected, and you will have *Napeir's* Logarithms, as was hinted before; but if you defire *Briggs's* Logarithms, which are now generally received, you must divide your Series by

2,302585092994045684017991454684364207 601101488628772976033328 or multiply it by the reciprocal thereof, viz.

0,4342944819032518276511289189166050822 94397005803666566114454

But to fave fo operofe a Multiplication (which is more than all the reft of the Work) it is expedient to divide this Multiplicator by the Powers of z or y continually, according to the Direction of the Theorem, effectially where x is finall and Integer, referving the proper Quotes to be added together, when you have produced your Logarithm to as many Figures as you defire: Of which Method I will give a Specimen.

If the Curiofity of any Gentleman that has leifure, would prompt him to undertake to do the Logarithms of all Prime Numbers under 100000, to 25 or 30 Figures, I dare affure him, that the facility of this Method willinvite him thereto; nor can any thing more eafie be defired. And to encourage him, I here give the Logarithms of the first Prime Numbers under 20 to 60 places, computed by the accurate Pen of Mr. Abraham Sharp, (from whofe Industry and Capacity the World may in time expect great Performances) as they were communicated to me by our common Friend Mr. Euclid Speidall.

Numb:

Tig and a Google

Logarithm. Numb. 0,30102999566398119521373889472449 2 3026768189881462108541310427 0,47712125471966243729502790325511 3 5309200128864190695864829866 0,84509804001425683071221625859263 7 6193483572396323965406503835 1,04139268515822504075019997124302 II 4241706702190466453094596539 1,11394335230683776920654189502624 13 6254561189005053673288598083 1,23044892137827302854016989432833 17 7030007567378425046397380368 1,27875360095282896153633347575692 19

The next Prime Number is 23, which I will take for an Example of the foregoing Doctrin; and by the first Rules, the Logarithm of the ratio of 22 to 23, will be found to be either

9317951129337394497598906819

I	IL	I	I	-	1
22	968	31944	937024	2576	1 8160 & c.or
23	1 1058	3650	1 11193	64	1 32181715
~~~					

As likewife that of the ratio of 23 to 24 by a like Process.

 $\frac{\frac{1}{23}}{1058} + \frac{1}{36501} + \frac{1}{1119364} + \frac{1}{32181715} & \text{&c.}$ or,

 $\frac{1}{24} + \frac{1}{1152} + \frac{1}{41472} + \frac{1}{1327104} + \frac{1}{39813120}$ &c. And

And this is the Refult of the Doctrine of Mercator, as improved by the Learned Dr. Wallis. But by the fecond Theorem, viz.

 $\frac{2x}{z} + \frac{2x^3}{3z^3} + \frac{2x^5}{5z5}$  &c. The fame Logarithms

are obtained by fewer steps. To wit,

 $\frac{2}{45} + \frac{2}{273375} + \frac{2}{922640625} + \frac{2}{2615686171875}$ &c. and

 $\frac{2}{47} + \frac{2}{311469} + \frac{2}{1146725035} + \frac{2}{3546361813211} \&c.$ 

which was invented and demonstrated in the Hyperbolick Spaces Analogous to the Logarithms, by the Excellent Mr. James Gregory, in his Exercitationes Geometrica, and fince further profecuted by the aforefaid Mr. Speidall, in a late Treatife, in English, by him published on this Subject. But the Demonstration as I conceive, was never till now perfected without the confideration of the Hyperbola, which in a matter purely Arithmetical as this is, cannot be fo properly applied. But what follows I think I may more juftly claim as my own, viz. That the Logarithm of the ratio of the Geometrical Mean to the Arithmetical between 22 and 24, or of  $\sqrt{528}$  to 23 will be found to be either.

 $\frac{1}{1058} + \frac{1}{1119364} + \frac{1}{888215334} + \frac{1}{626487882248}$ &c. or

 $\frac{1}{1057} + \frac{1}{3542796579} + \frac{1}{659676558485285} \&c.$ Au

All theife Series being to be multiplied into 0,4342944819 &c. if you defign to make the Logarithm of Briggs. But with great Advantage in refpect of the Work, the faid 434294 4819, &c. is divided by 1057 and the Quotient thereof again divided by three times the Square of 1057, and that Quotient again by  $\frac{5}{3}$ of that Square, and that Quotient by 7 thereof, and fo forth, till you have as many Figures of your Logarithm as you defire. As for Example; the Logarithm of the Geometrical Mean, between 22 and 24, is found by the Logarithms of 2, 3 and 11 to be

> 1057)43429 &c. 3 in 1117249)41087 &c. 3 in 1117249)12258 &c. 3 in 1117249)65832 &c. 3 in 1117249)42088 &c.

1.36131696126690612945009172669805 ( 41087462810146814347315886368 ( 12258521544181829460074 ( 6583235184376175 ( 4208829765 ( ________2930

Summa.

1.36172783601759287886777711225117

Which is the Logarithm of 23 to thirty two places, and obtained by five Divisions with very finall *Divisors*; all which is much lefs Work than fimply multiplying the *Series* into the faid Multiplicator 43429, &c.

Before I pass on to the converse of this Problem, or to shew how to find the Number appertaining 48

pertaining to a Logarithm affigned, it will be requisite to advertise the Reader, that there is a fmall mistake in the aforefaid Mr. James -Gregory's Vera Quadratura Circuli & Hyperbola, published at Padua Anno 1667. wherein he applies his Quadrature of the Hyperbola to the making the Logarithms; In pag. 48. he gives the Computation of the Lord Napeir's Logarithm of 10, to five and twenty places, and finds it 2302585092994045624017870 inftead of 2302585092994045684017991, erring in the eighteenth Figure, as I was affured upon my own Examination of the Number I here give you, and by comparison thereof with the fame wrought by another hand, agreeing therewith to 57 of the 60 places. Being defirous to be fitisfied how this difference arole, I took the no fmall trouble of Examining Mr. Gregory's Work, and at length found, that in the infcribed Polygon of 512 Sides, , in the eighteenth Figure, was a o instead of 9, which being re-Aified, and the fubfequent Work corrected therefrom, the refult did agree to a Unite with our Number. And this I propose not to Cavil at an easie mistake in managing of fo vast Numbers, especially by a Hand that has fo well deferved of the Mathematical Sciences. but to fhew the exact coincidence of two fo very differing Methods to make Logarithms, which might otherwife have been questioned.

From the Logarithm given to find what ratio it expresses, is a Problem that has not been fo much confidered as the former, but which is folved with the like eafe, and demonstrated

Miscellanea Curiosa. 49 ftrated by a like Process, from the fame gene-neral Theorem of Mr. Newton: For as the Logarithm of the ratio of 1 to 1 +9 was proved to be 1+q|m-1, and that of the ratio of 1 to 1-q to be  $1-\overline{1-ql^{m}}$ : fo the Logarithm, which we will from henceforth call L, being given, 1-L, will be equal to  $\overline{1-1-g}|^{\overline{m}}$  in the one cafe; and 1-L will be equal to  $\frac{1}{1-q}$  in the other: Confequently 14-11 " will be equal to 1-19, and 1-Li to 1-9; that is, according to Mr. Newton's faid Rule, 1-1-mL-1-1m2 L2-1-1m3 L3-1 $m^{1}L^{4} + \frac{1}{12}m^{3}L^{3}$  &c. will be=: 1-q, and  $1-m^{2}L^{4} + \frac{1}{12}m^{2}L^{3} - \frac{1}{2}m^{3}L^{2} + \frac{1}{12}m^{2}L^{4} - \frac{1}{12}m^{3}L^{5}$  &c. will be equal to 1-9, m being any infinite Index what foever, which is a full and general Proposition from the Logarithm given to find the Number, be the Species of Logarithm what it will. But if Napeir's Logarithm be given, the Multiplication by m is faved (which Multiplication is indeed no other than the redu-cing the other Species to his) and the Series will! be more fimple, viz.  $I + L + \frac{1}{2}LL + \frac{1}{2}L^{3} + \frac{1}{4}L^{4}$ &c. This Series, especially in great Numbers converges fo flowly, that it were to be wifhed it could be contracted.

E

If one term of the ratio, whereof L is the Logarithm, be given, the other term will be eafily had by the fame Rule: For if L were Napeir's Logarithm of the ratio of a the leffer to b the greater term, b would be the Product of a into 1-L+ = LL+ LLL &c. ==a+aL+= aLL---aL' &c. But if were given, a would be==b_bL1-ibL1-ibL' &c. Whence, by the help of the Chiliads, the Number appertaining to any Logarithm will be exactly had to the utmost extent of the Tables. If you feek the nearest next Logarithm, whether greater or leffer, and call its Number a if leffer, or b if greater than the given L, and the difference thereof from the faid nearest Logarithm you call 1; it will follow, that the Number anfwering to the Logarithm L will be either 4 into 1-1-1+211-1-2114-1-1-1 & &c. or elfe b into  $I = l - \frac{1}{2} l l - \frac{1}{2} l l + \frac{1}{2} l^4 - \frac{1}{220} l^5 \&c.$  wherein as l is lefs, the Series will converge the fwifter. And if the first 20000 Logarithms be given to fourteen Places, there is rarely occasion for the three first steps of this Series to find the Number to as many places. But for Ulacq's great Canon of 100000 Logarithms, which is made but to ten places, there is fcarce ever need for more than the first ftep a-1- al or a-1malin one cafe, or elfe b-bl or b-mbl in the other, to have the Number true to as many Figures as those Logarithms confist of.

If future Industry shall ever produce Logarithmick Tables to many more places than now we have them; the aforefaid Theorems will be of more use to reduce the correspondent Natural Numbers to all the places thereof. In order to make the first Chiliad ferve all Uses,

Ules, I was defirous to contract this Series, wherein all the powers of *l* are prefent, into one, wherein each alternate Power might be wanting; but found it neither fo fimple or inlform as the other. Yet the first step thereof is, I conceive, most commodious for Practice, and withal exact enough for Numbers not exceeding fourteen places, such as are Mr. Brigs's large Table of Logarithms; and therefore I recommend it to common Ufe.

It is thus: 
$$a - \frac{a!}{1 - \frac{b!}{1 - \frac{b!}{1$$

will be the Number answering to the Logarithm given, differing from the Truth by but one half of the third step of the former Series. But that which renders it yet more eligible; is, that with equal facility, it serves for Brigg's or any other fort of Logarithm, with the only variation of writing  $\frac{1}{m}$  instead of  $\tau_i$ ,

$\frac{al}{\frac{1}{1}-\frac{1}{2}l}$ and	$b - \frac{bl}{\frac{1}{m} - \frac{1}{2}l_{2}}$	or	$\frac{\frac{1}{m}a + \frac{1}{2}la}{\frac{1}{m} - \frac{1}{2}l}$	and
-------------------------------------------	-------------------------------------------------	----	-------------------------------------------------------------------	-----

that is

 $\frac{\frac{1}{m}b - \frac{1}{2}lb}{\frac{1}{m} - \frac{1}{2}l}, \text{ which are eafily refolv'd into } A= \frac{1}{m} - \frac{1}{2}l, \text{ alogies, } viz.$ As 43429 &c.  $-\frac{1}{2}l$  to 43429  $+\frac{1}{5}l::$  to the Num-

or As  $43429 \& c. + \frac{1}{2}l$  to  $43429 - \frac{1}{2}l :: \int_{fought.}^{ber} fought.$ 

E 1

If

52

If more fteps of this Series be defired, it will be found as follows,  $a + \frac{al}{1 - \frac{1}{2}l} - \frac{1}{1 - l} + \frac{3}{3}\frac{al^3}{1 - 2l}$  &c. as may eafily be demonstrated by

working out the Divisions in each step, and collecting the Quotes, whose Sum will be found to agreee with our former Series.

Thus I hope, I have cleared up the Doctrine of Logarithms, and fhewn their Conftruction and Ufe independant from the Hyperbola, whofe Affections have hitherto been made ufe of for this purpofe, though this be a matter purely Arithmetical, nor properly demonstrable from the Principles of Geometry. Nor have I been obliged to have recourfe to the Method of Indivisibles, or the Arithmetick of Infinites, the whole being no other than an easie Corollary to Mr. Newton's General Theorem for forming Roots and Powers.

Λ

#### A SOLUTION,

Given by Mr. John Collins, of a Chorographical Problem, Proposed by Richard Townley, Efq;

## PROBLEM.

TheDiftances of three Objects in the fame Plain being given, as A,B, C; The Angles made at a fourth Place in the fame Plain as at S, are observed : The Diftances from the Place of Observation to the respective Objects, are required.

#### The Problem hath fix Cafes.

Cafe 1. IF the Station be taken without the Triangle made by the Objects but in one of the fides thereof produced, as at S in the 9th Figure; find the Angle ACB; then in the Triangle ACS all the Angles and the fide AC are known, whence either or both the Diftances SA or SC may be found. 54

Cafe 2. If the Station be in one of the Sides of the Triangle, as in the 10th Figure at S, then having the three fides AC, CB, BA given, find the Angle CAB; then again in the Triangle SAB, all the Angles, and the fide AB, are known; whence may be found either AS, or SB, Geometrically, if you make the Angle CAD equal to the observed Angle CSB, and draw BS parallel to DA, you determine the Point of Station S.

Cafe 3. If the three Objects lie in a right Line as A C B (suppose it done) and that a Circle passion of the station S, and the two exterior Objects A B, then is the Angle A B Dequal to the observed Angle A S C (by 21 of the 3d Book of *Euclid*) as infifting on the fame Arch A D: And the Angle B A D in like manner equal to the observed Angle C S B: By this means, the point D is determined. Join DC, and produce the fame, thn a Circle passing through Points A B D, interfects DC, produced at S, the place of Station.

#### Calculation.

In the Triangle A B D, all the Angles and the fide A B are known, whence may be found the fide A D.

Then in the Triangle  $C \land D$  the two fides  $C \land A$  and  $\land D$  are known and their contained Angle  $C \land D$  is known; whence may be found the Angles  $C D \land A$  and  $\land C D$ , the complement whereof to a Semicircle is the Angle S  $C \land A$ : in which Triangle the Angles are now all known and the fide  $\land C$ : whence may be found either of the Diftances, SC or  $S \land A$ . Cafe 4:

Cafe 4. If the Station be without the Triangle, made by the Objects, the fum of the Angles observed is less than four right Angles. The Construction is the fame as in the last Cafe, and the Calculation likewife; faving that you must make one Operation more, having the three Sides, AC, CB, BA, thereby find the Angle C A B, which add to the Angle E A D, then you have the two fides, viz. A C, being one of the Diftances, and A D, (found as in the former Cafe) with their contained Angle CAD, given to find the Angles CDA, and ACD, the Complement whereof to a Semicircle, is the Angle SC A: Now in the Triangle SCA, the Angle at C being found, and at S observed, and given by Supposition, the other at A is likewife known, as being the complementof the two former to a Semicircle, and the fide A C given ; hence the Diftance C S or AS may be found.

Cafe 5. If the place of Station be at fome Point within the plain of the Triangle, made by the three Objects, the Conftruction and Calculation is the fame as in the laft, faving only that inftead of the obferved Angle ASC, the Angle ABD is equal to the Complement thereof to a Semicircle, to wit, it is equal to the Angle ASD; both of them infifting on the fame Arch AD: And in like manner the Angle BAD is equal to the Angle DSB, which is the Complement of the obferved CSB; and in this Cafe, the fum of the three Angles obferved, is equal to four right Angles.

In

56

In these three latter Cases no use is made of the Angle observed between the two Objects, as A and B, that are made the Base-line of the Construction; yet the same is of ready use for finding the third Distance, or last fide sought, as in the fourth Scheme, in the Triangle SAB, there is given the Distance AB, its opposite Angle equal to the sum of the two observed Angles, and the Angle SAB attained, as in the fourth Case: Hence the third Side or last Distance SB may be found.

And here it may be noted, that the three Angles  $C \land S, \land S B, S B C$ , are together equal to the Angle  $\land C B$ , for the two Angles C S B, and C B S, are equal to E C B, as being the Complement of S C B to two right Angles; and the like in the Triangle on the other fide. Ergo, C c.

Cafe 6. If the three Objects be A, B, C, and the Station at S, as before, it may happen, according to the former Constructions, that the Points C and D may fall close together, and fo a right Line joining them may be produced with uncertainty; in fuch cafe the Circle may be conceived to pass through the place of Station at S, and any two of the Objects (as in the fixth Scheme) through B and C; wherein making the Angle DBC equal to the observed Angle ASC, and BCD equal to the Complement to 180 degrees of both the observed Angles in DSB thereby the Point D is determined, through which, and the points C B, the Circle is to be defcribed, and joyning DA, (produced, when need requireth) where it interfects

tersects the Circle, as at s, is the place of Station fought.

This Problem may be of good Ufe for the due Situation of Sands or Rocks, that are within fight of three Places upon Land, whole diftances are well known; or for *Chorographi*cal Ufes, & c. Efpecially now there is a Method of obferving Angles nicely accurate by aid of the *Telescope*; and was therefore thought fit to be now Publish'd though it be a competent time fince it was delivered in Writing.

The

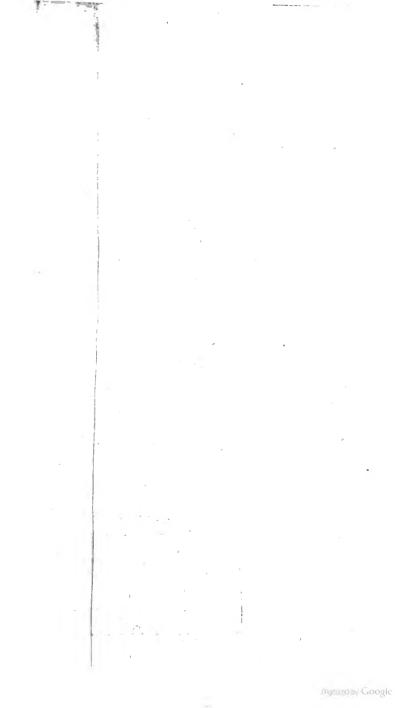
58

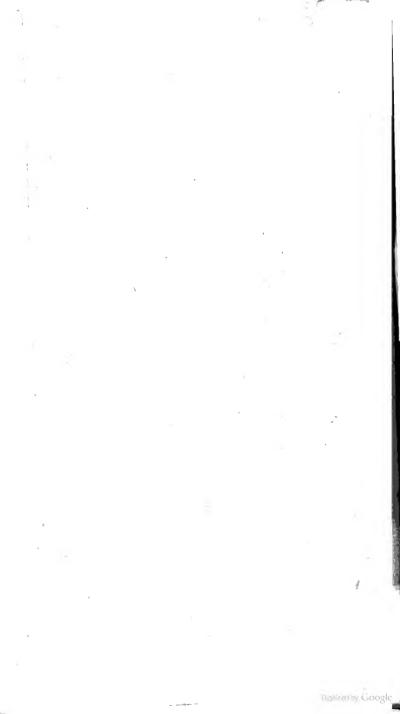
The Solutions of three Chorographic Problems, by a Member of the Philosophical Society of Oxford.

THE three following Problems may occur at Sea, in finding the diffance and polition of *Rocks*, Sands, &c. from the Sea Shoar; or in the Surveying of the Sea Coaft; When only two Objects, whole diffance from each other is known, can be feen at one Station; but efpecially they may be useful to one that would make a Map of a Country by a Series of Triangles derived from one or more measured Bases; which is the most exact way of finding the bearing and diffance of Places from each other, and thence their true Longitude and Latitude; and may confequently occur to one that would in that manner meafure a Degree on the Earth.

#### The first Problem (Fig. 3 and 4.)

There are two Objects, B and C, whofe diftance BC is known; and there are two ftations at A and E, where the Objects B Cbeing visible, and the Stations one from another, the Angles BAC, BAE, AEB, AEC, are known by Observation, (which may be made with an ordinary Serveying semicircle, or Crostaff; or if the Objects be beyond





17

beyond the view of the naked Eye, with a *Telescopick Quadrant*) to find the distances or lines *AB*, *AC*, *AE*, *EC*.

#### Construction.

In each of the Triangles BAE, CAE, two Angles at A, E, being known, the third is alfo known: then take any line  $\alpha \cdot \alpha$  pleafure, on which conflitute the Triangles  $\beta \alpha \cdot \alpha$ ,  $\alpha \cdot \gamma$  refpectively equiangular to the Triangles BAE, AEC; join  $\beta \gamma$ . Then upon BC conflitute the Triangles BCA, BCE, equiangular to the correspondent triangles  $\beta \gamma \alpha$ ,  $\beta \gamma \cdot \alpha$ , join AE, and the thing is manifestly done.

#### The Calculation.

Affuming  $\alpha \in$  of any number of parts, in triangles  $\alpha \beta \in, \alpha \gamma \in$ , the angles being given, the fides  $\alpha \beta, \alpha \gamma, \epsilon \beta, \epsilon \gamma$  may be found by Trignometry: Then in the Triangle  $\beta \alpha \gamma$ , having the angle  $\beta \alpha \gamma$ , and the legs  $\alpha \beta, \alpha \gamma$ , we may find  $\beta \gamma$ . Then  $\beta \gamma$ .  $BC:: \beta \alpha$ . BA:: $\beta \in$ .  $BE:: \gamma \alpha$ .  $CA:: \gamma \in CF$ .

### The second Problem (Fig. 5 and 6.)

Three Objects B, C, D, are given, or (which is the fame) the fides, and confequently angles of the triangle BCD are given; alfo there are two points or flations A, E, fuch, that at A may be feen the three points BCE, but not D; and at the flation E may be feen A, C, D, but not B, that is the angles BAC, BAE, AEC, AED, (and confequently EAC, AEC, are known by obfervation: to find the lines AB, AC, AE EC, ED,

#### Construction.

Take any line as at pleafure, and at its extremities make the angles say, sab, asy, es, equal to the correspondent observed an-Stes EAC, EAB, AEC, AED. Produce Ba, ", till they meet in , join , ; then upon CB defcribe (according to 33. 3. Eucl.) a fegment of a circle that may contain an angle=  $\gamma \circ \beta$ ; and upon C D defcribe a fegment of a circle capable of an angle=205; fuppofe F the common fection of these two circles; join FB, FC, FD; then from the point C, draw forth the lines CA, CE, fo that the angle FCA may be=? y a, and FCE=? y :; fo A, E, the common Sections of CA, CE, with FB, FD, will be the points required, from whence the reft is eafily deduced.

### The Calculation.

Affuming  $\alpha \in$  of any number, in the triangles  $\alpha \gamma \in$ ,  $\alpha \phi \in$ , all the angles being given, with the fide  $\alpha \in$  alfum'd, the fides  $\alpha \gamma, \epsilon \gamma, \alpha \neq$ ,  $\epsilon \phi$ , will be known; then in the triangle  $\gamma \alpha \phi$ , the angle  $\gamma \alpha \phi$ , with the legs  $\alpha \gamma, \alpha \phi$ , being known, the angles  $\alpha \phi \gamma, \alpha \gamma \phi$ , with the fide  $\phi \gamma$  will be known: then as for the reft of the work in the other figure, the triangle B C Dhaving all its fides and angles known, and the angles BFC, BFD, being equal to the found  $\beta \phi \gamma$ ,  $\beta \phi \phi$ ; how to find FB, FC, FD by Calculation (and alfo Protraction) is fhewn by Mr. Collins in the precedent Difcourfe, as to all its cafes, which may therefore fuperfede my fhewing any other way.

But here it must be noted, that if the fum of the observed angles,  $B \ A E$ , A E D, is 180 degrees: then A B and E D cannot meet, because they are parallel, and consequently the given Solution cannot take place; for which reason I here subjoin another.

#### Another Solution.

Upon BC (Fig. 7.) describe a fegment BAC of a circle, fo that the angle of the fegment may be equal to the observed L Bay, (which as above quoted is shewn 33. 3. Euclid.) and upon CD describe a segment CED of a circle capable of an angle equal to the observed CED; from c draw the diameters of these circles cGc H; then upon CG describe a segment of a circle GFC, capable of an angle equal to the observed LAEC; likewise upon CH describe a circle's fegment CFH, capable of an angle equal to the observed  $C \wedge E$ : suppose F the common Section of the two last circles HFc. GFC, join FH, cutting the circle HEC in E, join alfo FG, cutting the circle GAC in A: 1 fav that A, E, are the points required.

Demonstration.

For the  $\angle B \land C$  is  $\exists a \gamma$  by conftruction of the fegment, also the angles  $C \mathrel{E} H$ ,  $C \land G$ , are right, because each exists in a semicircle: therefore a circle being described upon  $C \mathrel{F}$  as a diameter, will pass through  $\mathrel{E}, \mathrel{A}$ ; Therefore the angle  $C \land \mathrel{E} = \mathrel{L} C \mathrel{F} \mathrel{E} = \mathrel{C} \mathrel{F} \mathrel{H} = (by$ construction) to the observed angle  $\gamma \mathrel{a} \mathrel{\varepsilon}$ . In like manner the  $\mathrel{L} C \mathrel{E} \mathrel{A} = \mathrel{C} \mathrel{F} \mathrel{A} = \mathrel{C} \mathrel{F} \mathrel{G} = ob$  $forv'd angle <math>\gamma \mathrel{\varepsilon} \mathrel{a}$ .

In the flations A, E, fall in a right line with the point C; the lines G A; HE being parallel

62

lel, cannot meet: but in this cafe the Problem is indeterminate and capable of infinite Solutions. For as before upon CG defcribe a Segment of a circle capable of the obferved  $L\gamma \in \alpha$ ; and upon CH, defcribe a Segment capable of the obferved  $\gamma \alpha \in :$  then through C, draw a line any way cutting the circles in A, E, thefe points will answer the question.

### The third Problem.

Four points B, C, D, F, (Fig. 8.) of the 4 fides of a quadrilateral, with the angles comprehended are given; also there are two ftations A and E fuch, that at A, only B C E are visible, and at E only A D F, that is, the angles B A C, B A E, A E D, D E F are given: to find the places of the two points A E, and confequently, the lengths of the lines AB, AC; AE, E D, E F;

### Construction.

Upon *BC* (by 33. 3. *Eucl.*) defcribe a fegment of a circle, that may contain an angle equal to the obferved angle *BAC*, then from *C* draw the Chord *C M*, or a line cutting the circle in *M*; fo that the Angle *BCM* may be equal to the fupplement of the obferved angle *BAE*, i. e. its refidue to 180 degrees. In like manner on *D F* defcribe a fegment of a circle, capable of an angle equal to the obferv'd *DE F*, and from *D* draw the Chord *DN*, fo that the angle *F D N* may be equal to the fupplement of the obferv'd angle *AEF*, join *MN*, cutting the two circles in *A, E*: I fay , E, are the two points requir'd.

Demonia

### Demonstration.

Join AB, AC, ED, EF, then is the  $\angle MAB$ = $\angle BCM$  (by 21.3. Eucl.)= fupplement of the obferv'd  $\angle BAE$  by confirution, therefore the confiructed  $\angle BAE$ , is equal to that which was obferved. Alfo the  $\angle BAC$  of the fegment is the confiruction of the Segment, equal to the obferv'd  $\angle BAC$ . In like manner the confiructed angles AEF, and DEF, are equal to the correspondent observed angles AEF, DEF, therefore AE are the points requir'd.

### The Calculation.

In the Triangle BC M, the L BC M (=fupplement of BAE) and LBMC (= BAC) are given, with the fide BC; thence MC may be found; in like manner DN in the  $\Delta DNF$ may be found. But the  $(MCD) = BCD_{---}$ RCM) is known, with its legs MC, CD, therefore its Bafe MD, and LMDC, may be known. Therefore the  $L M D N (= C D F_{--})$ C D M -- F D N) is known, with its legs MD. D N; thence MN with the angles DMN, DNM, will be known. Then the LCMA(=(LDMC - |-DMN)' is known, with the LMAC =MAB+BAC) and MC before found; therefore MA and AC will be known. In like manner in the triangle EDN, the angles E, N, with the fide DN being known, the fides EN, E D, will be known; therefore  $A \in (= M N - - -$ MA - EN) is known. Alfo in the triangle A BC, the LA with its fides BC, CA, being known, the fide A B, will be known, with the L B C A; so in the triangle E F D, the L E with the fides, ED, DF being known, EF will be found, with the LEDF. Laftly, in the triangle

64

gle A C D, the L A C D (= B C D - B C A) with its legs A C, C D being known, the fide A D, will be known; and in like manner E C in the triangle E D C.

Note, that in this Problem, as also in the first and second, if the two stations fall in a right line with either of the given Objects: the *locus* of A, or E, being a circle, the particular point of A, or E, cannot be determined from the things given.

As to the other cafes of this third Problem, wherein A and E, may fhift places, *i.e.* only DFE, may be visible at A, and only A, B, C, at E; or wherein B, D, E, may be visible at A, and only C, F, A, at E; or wherein A may be of one fide of the quadrilateral, and E on the other; or one of the stations within the quadrilateral, and the other without it: I shall for brevity fake omit the Figures, and diversity of the Sines + and-- in the calculation, and presume that the Surveyor will eafily direct himself in those cases, by what has been faid.

The folution of this third Problem is general, and ferves also for both the precedent. For suppose C D, the same point in the last figure, and it gives the folution of the second Problem: but if B C be supposed the same points with D, F, by proceeding as in the last, you may directly solve the first Problem.

Dig and Google

An Arithmetical Paradox, concerning the Chances of Lotteries; by the Honourable Francis Roberts, Esq; Fellow of the R. S.

A S fome Truths (like the Axioms of Geos metry and Metaphyficks) are felf-evident at the first View, fo there are others no lefs certain in their Foundation, that have a very different Aspect, and without a strict and careful Examination, rather seem repugnant.

We may find Instances of this kind in most Sciences.

In Geometry, That a Body of an infinite Length, may yet have but a finite Magnitude.

In Geography, That if Antwerp be due East to London, for that reason London cannot be West to Antwerp.

In Aftronomy, That at the Barbadoes (and other Places between the Line and Tropick) the Sun, part of the Year, comes twice in a Morning to fome Points of the Compafs.

In Hydroftaticks, That a hollow Cone (ftanding upon its Basis) being fill'd with Water, the Water shall press the bottom with three times the Weight, as if the same Water was frozen to Ice; and Figures might be contriv'd to make it press a hundred times as much.

F

Thef:

These Speculations, as they are generally pleafant, fo they may also be of good use to warn us of the Mistakes we are liable to, by careless and superficial Reasoning.

I shall add one Instance in Arithmetick, which perhaps may seem as great a Paradox as any of the former.

There are two Lotteries, at either of which a Gamester paying a Shilling for a Lot or Throw; The first Lottery upon a just Computation of the Odds, has 3 to 1 of the Gamester; the Second Lottery, but 2 to one; nevertheles, the Gamester has the very fame difadvantage (and no more) in playing at the First Lottery, as the Second.

It looks very like a Contradiction, that the Difadvantage fhould be no greater in playing against 3 to 1, than 2 to 1, but it may thus be refolv'd.  $\cong \subseteq 1/2 \ge 5 \le 3 \ge 5 \le 3 \ge 5 \le 16$  pence 2

In the first Lottery the Gamester hazards a Shilling to win a Groat, and the Chances being equal, it is evident there is 3 to one against him.

In the Second Lottery, the Gamester ventures a Shilling against a Shilling, and the Lots being 4 to 2, his Difadvantage is 2 to 1.

And a Lot at either of them being truly worth juft 8 Pence, (viz. the 6th part of 3 times 16 Pence, or twice 2 Shillings) the Difadvantage muft be the very fame in both Gafes, that is, the Gamefter pays a Shilling for a Lot that is worth but 8 Pence.

The

66

The Method of finding this Answer being fomewhat out of the common Road, I shallhere add it, and thereby infinite Solutions of the fame kind may be discovered.

### Ist. Lottery.

Let a=the number of Blanks b=the number of Prizes. r=the Value of a Prize.

### 2d. Lottery.

Let m=the number of Blanks. n=the number of Prizes. s=the value of a Prize.

### I =to what you pay for a Lot, viz. a Shilling.

So the Lottery has its Chances for 1, and the Gamester his for r-1. Now the true Odds confisting of the compounded Proportion of the Chances and the Values, viz.  $\frac{a}{b}$  and  $\frac{1}{r-1}$ , the Share of the Lottery will be a, and that of the Gamester rb-b. Therefore as the prefent cafe stands, the first Lottery mult be a=3rb-3b, and by the like reasoning, the second Lottery will be m=2 sn-2n. Now the Value of a Lot being the Sum of the Prizes divided by the number of Lots, (which must be equal in both Lotteries) it yields

 $\frac{rb}{a+b} = \frac{sn}{m-n}$ 

F 2

68 <i>b</i> <i>r</i> <i>m</i> <i>m</i> <i>s</i>	Miscellanea Curiosa. $   \begin{array}{l} 1 4 = 3rb - 3b \\ 2m = 2sn - 2n \\ 3 \hline \epsilon + b \end{array} = \frac{s}{m + n} \\   \begin{array}{l} 4(*) \\ 5(*) \\ 6(*) \end{array} $
9 =? 7*4+6 8×3 1+36 9, 10	7 Let $\frac{rb}{a+b} = q$ 8 $rb = qa + qb$ 9 $3rb = aqa + qb$ 10 $3rb = a + 3b$ 11 $3qa + 3qb = a + 3b$
Scope 11, 12 13÷3b 12, 14	12 If $a \equiv 0$ to avoid-negative Numbers. 13 $3b \equiv 34b$ 14 $q \equiv 1$ 15 $q > 1$ makes $a < 0, q < 1$ makes $a > 0$
17 - 34	16 If $b = 0$ 17 3 $q 4 = 4$ 18 $q = \frac{1}{2}$ 19 $q < \frac{1}{2}$ makes $b < 0$ $q > \frac{1}{2}$ makes $b > 0$
$3, \frac{7}{20*m+n}$ $20*m+n$ $21 \times 2$ $2+2n$ $22, 23$	$20 \frac{s n}{m + n} = 4$ $21 \frac{sn}{sn} = qm + qn$ $22 \frac{2sn}{23} \frac{sn}{sn} = m + 2n$ $24 \frac{2qm + 2qn}{sn} = m + 2n$
Scope 24, 25 26÷2n 25, 27	25 If m = 0 26 2 q n = 2n 27 q = 1 28 q > 1 makes m < 0 q < 1 makes m > 0 Scope

1.2	Miscellanea Curiosa. 69
Scope	29  If  n = 0
24, 29	$30 2 q m \equiv m$
∗ 30÷2m	
29, 31	$3^{2} q < \frac{1}{2} makes n < o q > \frac{1}{2} n > o$
15,19,28,32	33 that $abm n$ may be $> o,q$ must be $> \frac{1}{2} < 1$
33, 4 (*)	34 Let therefore $\varrho =$
7 34.	$35 \frac{rb}{r} = \frac{2}{r}$
35 *, 10	4-6 3
	13
	37   a = b
20.24	38 5 7 2
20, 34.	
38 *	$39 \frac{m+n}{3}$
39 * 2	$\begin{array}{c} 40 \\ 65n = 2m - 2m \\ 65n = 4m - 4n \end{array}$
23 * 3	$\frac{1}{1} 65n = 3m + 6n$
40, 41	$4^{2} 4^{m} + 4^{n} = 3^{m} + 6^{n}$
42	13 m = 2n
1÷37	14 = 3r - 3
44-1-3	$15 3^{r} = 4$
2 - 1, 43	16 2 = 25-2
46-1-2	$_{47}^{25} = 4$
4014	*/
5 (*)	18 Let $A = 3$
37, 48	19 B = 3
45 - 3	$\mathcal{A} = \frac{4}{3}$ , id eft, 16 Pence.
6 (*)	51 Let $M = 4$
43, 51	$\sum_{n=1}^{\infty} N = 2$
47 ÷ 2	[53] $S \equiv 2$ 2 Shillings.
1	
	F 2

F 3

4 New, Exact and Eastie Method, of finding the Roots of any Equations Generally, and that without any previous Reduction. By Edm. Halley.

HE principal use of the Analytick Art, is to bring Mathematical Problems to Equations, and to exhibit those Equations in the most simple Terms that can be. But this Art would justly feem in fome degree defective, and not sufficiently Analytical, if there were not fome Methods, by the help of which, the Roots (be they Lines or Numbers, might be gotten from the Equations that are found, and fo the Problems in that refpect be folved. The Ancients fcarce knew any thing in thefe Matters, beyond Quadratick Equations. And what they writ of the Gcometrick Construction of folid Problems, by the help of the Parabola, Ciffoid, or any other Curve, were only particular things defign'd for fome particular Cafes. But as to Numerical Extraction, there is every where a profound Silence; fo that whatever we perform now in this kind, is entirely owing to the Inventions of the Moderns. +

And first of all, that great Discoverer and Reftorer of the Modern Algebra, Francis Vieta, about 100 Years fince, shew'd a general Me-1 25

thod

thod for extracting the Roots of any Equation, which he publish'd under the Title of, A Numerical Refolution of Powers, &c. Harrior, Oughtred, and others, as well of our own Country, as Foreigners, ought to acknowledge whatfoever they have written upon this Subject, as taken from Vieta. But what the Sagacity of Mr. Newton's Genius has perform'd in this buinefs, we may rather conjecture (than be fully assured of) from that short Specimen given by Dr. Wallis in the 94th Chapter of his Algebra. And we must be forc'd to expect it, till his great Modesty shall yield to the Intreaties of his Friends, and suffer those curious Discoveries to fee the Light.

Not long fince (viz. A. D. 1690) that excellent Perfon M. Joseph Raphson, F. R. S. publish'd his Universal Analysis of Equations, and illustrated his Method by plenty of Examples; by all which he has given Indications of a Mathematical Genius, from which the greatest things may be expected.

By his Example, M. de Lagney an ingenious Professor of Mathematicks at Paris, was encourag'd to attempt the same Argument; but he being almost altogether taken up in extracting the Roots of pure Powers (especially the Cubick) adds but little about affected Equations, and that pretty much perplex'd too, and not sufficiently demonstrated. Yet he gives two very compendious Rules for the Approximation of a Cubical Root; one a Rational, and the other an Irrational one. Ex. gr. that the fide of the Cube aaa+b, is between

 $a \rightarrow \frac{ab}{3aaa-1-b}, \& \sqrt{\frac{1}{4}aa+\frac{b}{3a}} \rightarrow \frac{1}{2}a.$ 

And the root of the 5th Power  $a^{1+b}$ , he makes

 $= \frac{1}{5}a + \sqrt{\sqrt{\frac{1}{4}a^4} + \frac{b}{5a}} - \frac{1}{4}aa} \quad (\text{where Note, that})$ 

'tis 1 aa; not 1 aa, as 'tis erroneously Printed in the French Book.) These Rules were communicated to me by a Friend, I having not feen the Book ; but having by tryal found the goodnefs of them, and admiring the Compendium, I was willing to find out the Demon-Which having done, I prefently ftration. found that the fame Method might be accommodated to the Refolution of all forts of Equations. And I was the rather inclin'd to improve these Rules, because I faw that the whole thing might be Explain'd in a Synopfis; and that by this means, at every repeated ftep of the Calculus, the Figures already found in the Root, would be at least Trebled, which all other ways, are encreafed but in an equal Number with the given ones. Now, the foremention'd Rules are eafily demonstrated from the Genefis of the Cube, and the 5th Power. For, fuppoling the fide of any Cube = at e, the Cube ariling from thence, is anot 3ane + 3 are + eee. And confequently, if we fuppofe and the next lefs Cube, to any given Noncubick Number, then ece will be lefs than Unity, and the remainder b, will = the other Members of the Cube, 3ane+ 3acefeet. Whence rejecting eee upon the account of its finalincis, we have b 3 ane-1- 3 are. And fince das:

Digitarda / Google

will not much exceed e; fo that putting  $e = \frac{b}{3aa}$  then the quantity  $\frac{b}{3aa+3ae}$  (to which

e is nearly equal) will be found

$$= \frac{b}{3aa+3ab} \quad \text{or } \frac{b}{3aa+b} \quad \text{that is } \frac{ab}{3aa+b} = e$$

And fo the fide of the Cube a a a + s will be  $a + \frac{ab}{3aa+b}$ , which is the Rational Formula of M. de Lagney. But now, if and were the next greater Cubick Number to that given, the fide of the Cube ana-b, will after the fame manner be found to be  $a - \frac{ab}{3aa-b}$ . And this eafy and expeditious Approximation to the Cubick Root, is only (a very fmall matter) erroneous in point of defect, the quantity e, the remainder of the Root thus found, coming fomething lefs than really 'tis.

As for the *Irrationale Formula*, 'tis deriv'd from the fame Principle, viz. b = 3aae+3aee, or  $\frac{b}{3a} = ae+ee$ , and fo  $\sqrt{\frac{1}{4}aa+\frac{b}{3a}} = \frac{1}{3}a+e$ , and  $\sqrt{\frac{1}{4}aa+\frac{b}{3}} + \frac{1}{3}a = a+e$ , the Root fought. Alfo the fide of the Cube aaa-b, after the fame manner, will be found to be  $\frac{1}{2}a+\sqrt{\frac{1}{4}aa-\frac{b}{3a}}$ And

74

And this Farmida comes fomething nearer to the Scope, being erroneous in point of excels, as the other was in defect, and is more accommodated to the ends of Practice, fince the Reflitution of the Calculus, is nothing elfe but the continual addition or fubstraction of the according as the quantity e can Quantity 3a be known! So that we fhould rather write  $\sqrt{\frac{1}{4}a_{+}b_{-}eee}$  in the former cafe, and in the latter,  $\frac{1}{2}a + \sqrt{\frac{1}{4}aa} + \frac{eee.-b}{2}$ . But by either of the two Formula's, the Figures already - known in the Root to be extracted, are at leaft Tripled ; which I conclude will be very grateful to all the Students in Arithmetick ; and I congratulate the Inventor upon the account of his Difcovery. But that the use of these Rules may be the better perceiv'd, I think it proper to fubjoin an Example or two. Let it be propos'd to find the fide of the double Cube, or ana 1+b=2. Here a = 1, and  $\frac{0}{3a} = \frac{1}{2} \sqrt[3]{8} \text{ fo}'_{1} + \sqrt{\frac{1}{12}} \text{ or } 1, 26$ , be found to be the true fide nearly. Now, the Cube of 1, 26, is 2, 000376, and 10,0,63 1,3969-,000376 or 0, 63 +1, 39680052 3, 78 132.5 L 91005291 = 1,259921049895 +; which in 13 Figures, gives the fide of the double Gube, with very little trouble, viz. By one only division; and the extraction of the fquare Roof; when as by the common way of working, how much

75

much pains it would have coft, the Skilful very well know. This Calculus a Man may continue as far as he pleafes, by encreasing the Square by the addition of the quantity eee; which, Correction, in this cafe will give, but the encrease of Unity in the 14th Figure of the Root. · Exemp. 2d. Let it be propos'd to find the fides of a Cube equal to that English Measure commonly call'd a Gallon, which contains 231 folid Ounces. The next lefs Cube is 216, whole fide 6 = a, and the remainder 15 = b; and fo for the first Approximation, we have  $3+\sqrt{9+\frac{5}{2}} =$  the Root. And fince  $\sqrt{9,8333}$ ... is 3,1358..., 'tis plain that 6,358= 4)-e. Now, let 6,1358= a; and we shall then have for its Cube 231,000853894712, & according to the Rule, 3,0679 tv9, 41201041-,000858394712 18,4070 is most accurately equal to the fide of the given Cube, which within the space of an Hour, I determin'd by Calculation to be 6.1357924 3966195897, which is exact in the 18th Figure, defective in the 19th. And this Formula is defervedly preferable to the Rationale, upon the account of the great Divisor, which is not to be manag'd without a great deal of Labour; whereas the extraction of the fquare Root, proceeds much more eafily, as manifold Experience has taught me. But the Rule for the Root of a pure Surfohd, or the 5th Power, is of fomething a higher Enquisy, and docs much more perfectly yet, do. 7:17

76

do the business; for it does at least Quintuple the given Figures in the Root, neither is the Calculus very large or operofe. Tho' the Author no where fnews his method of Invention. or any Demonstration, altho' it feems to be very much wanting ; especially fince all things are not right in the printed Book, which may eafily deceive the Unskilful. Now the sth power of the fide a+e is compos'd of these Members,  $a^3 + 5a^{+}e^{+}10a^{3}e^{2} + 10a^{3}e^{3} + 5ae^{+}+e^{3}=a^{+}+b$ ; from whence  $b=5a^{+}e^{+}+10a^{3}e^{3}+10$  $a^{2}e^{3}+5ae^{+}$ , rejecting  $e^{5}$  because of its smalness. Whence  $\frac{b}{5a} = a^3e + 2a^2e^2 + 2ae^3 + e^4$ , and adding on both fides int, we shall have Viat 1 b = VIat +aie+2a2e2 +2ae3 +et= : aa +ae-tee. Then fubstracting tas from both fides, ta-le wifi =  $\sqrt[]{\sqrt{\frac{1}{4}a^4 + \frac{b}{5a} - \frac{1}{4}aa}}$ ; to which if  $\frac{1}{3}a$  be added, then will  $a + e = \frac{1}{2}a + \sqrt{\sqrt{\frac{1}{2}a + b}} = \frac{1}{2}aa$ = the root of the Power a' + b. But if it had a'-b (the quantity a being too great) the Rule would have been thus,  $\frac{1}{4} + \sqrt{\frac{1}{4}a^2 - b^2 - \frac{1}{4}a^2}$ And this Rule approaches wonderfully, fo that

And this Rule approaches wonderfully, to that there is hardly any need of Restitution.

But while I confidered thefe things with my felf, I light upon a General Method for the Formula's of all Powers whatfoever, and (which being handfome and concife enough) I thought I would not conceal from the Publick. Thefe Miscellanea Curiosa. 77 These Formula's, (as well the Razional, as the Irrational ones) are thus.

$$\sqrt{aa+b} = \sqrt{aa+b}, \text{ or } a + \frac{ab}{2aa+b},$$

$$\sqrt{aa+b} = \sqrt{aa+b}, \text{ or } a + \frac{ab}{3aa+b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{3a}} \text{ or } a + \frac{ab}{3aaa+b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{6aa}}, \text{ or } a + \frac{ab}{4a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{4a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{5a^{2}+2b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{6a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{6a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{6a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{6a^{2}+\frac{1}{2}b},$$

$$\sqrt{aa+b} = \frac{1}{2}a + \sqrt{\frac{1}{2}aa+\frac{b}{10a^{2}}}, \text{ or } a + \frac{ab}{7a^{2}+\frac{1}{2}b},$$

And fo also of the other higher Powers. But if a were affumed bigger, than the Root fought (which is done with fome advantage, as often as the Power to be Refolved, is much nearer, the Power of the next greater whole Number, than of the next lefs) in this cafe, Mutatis Mutandis, we shall have the same Expressions of the Roots, viz.

Vaa

Miscellanea Curiosa. 78  $\sqrt{aa-b} = \sqrt{aa-b}$ , or  $a = \frac{ab}{2aa-b}$  $\sqrt[a]{a^3-b} = \frac{1}{a} + \sqrt[r]{aa-\frac{b}{3a}}, \text{ or } a - \frac{ab}{3a^3-b}$  $\sqrt{a^{2}}-b=\frac{1}{2}a+\sqrt{\frac{1}{2}aa-\frac{b}{6}aa^{2}}$  or  $a-\frac{ab}{4a^{2}-\frac{1}{2}b^{2}}$  $a^{1}-b = \frac{3}{4}a + \sqrt{\frac{1}{16}aa - \frac{b}{10a^{3}}}$  or, a  $a^{-}b = \frac{1}{5}a + \sqrt{\frac{1}{23}aa - \frac{b}{15a^{2}}}$  or  $a - \frac{ab}{6a^{6} - \frac{1}{5}b}$ .  $b = \frac{1}{2}a + \sqrt{\frac{1}{3}}a^{44} + \frac{b}{2}a^{5}, \text{ or } a - \frac{ab}{7a^{7} - 3b}.$ 

And within these two Terms, the true Root is ever found, being fomething nearer to the Irrational than the Rational Expression. But the quantity e found by the Irrational Formula, is always too great, as the Quotient refulting from the Rational Formula, is always too little. And confequently, if we have -|-b, the Irrational Formula gives the Root fomething greater than it should be, and the Rational fomething lefs. But contrary wife if it be -b.

And

And thus much may fuffice to be faid, concerning the extraction of the Roots of pure . Powers; which notwithstanding, for common . Ufes, may be had much more eafily by the help of the Logarithms. But when a Root is to be determin'd very accurately, and the Logarithmick Tables will not reach fo far, then. we must necessarily have recourse to these, or fuch like Methods. Farther; the Invention and Contemplation of these Formulæ, leading me to a certain Universal Rule, for adfected Equations (which I hope will be of use to all the Students in Algebra and Geometry) I was willing here to give fome account of this Difcovery, which I will do with all the perspecuity I can. I had given at No. 188 of the Transactions, a very easy and general constrution of all adfected Equations, not exceeding the Biquadratick Power; from which time I had a very great defire of of doing the fame in Numbers. But quickly after, Mr. Raphfon feem'd in great measure to have fatisfy'd this Defire, till Mr. Lagney by what he had perform'd in his Book, intimated that the thing might be done more compendioully yet. Now, my Method is thus.

Let z the root of any Equation, be imagin'd to be compos'd of the parts a + or - e, of which, let a be allum'd as near z as is pollible; which is notwith flanding not neceffary, but only commodious. Then from the Quantity a+e or a-e, let there be form'd all the Powers of z, found in the Equation, and the Numerical Co-efficients be refpectively affix'd to them : Then let the Power to be refolv'd, be fubfracted from the fum of the given Parts (in the 80

the first Column where e is not found) which they call the Homogeneum Comparationis, and let the difference be  $\pm b$ . In the next place, take the fum of all the Co-efficients of e in the fecond Column, to which put  $\equiv s$ . Lastly, in the third Column let there be put down the fum of all the Co-efficients of e e, which fum call t. Then will the Root z stand thus in the Rational Formula, viz.  $z \equiv a + \frac{sb}{ss \pm tb}$ ; and thus in the Irrational Formula, viz.

 $z = a \mp \frac{1}{2} s \pm \sqrt{\frac{1}{2} s s \mp b t}$ ; which perhaps it

may be worth while to Illustrate by some Examples. And inftead of an Instrument, let this Table ferve, which shews the Genesis of the feveral Powers of ate, and if need be, may eafily be continued farther; which for its use I may rightly call a General Analytical The forementioned Powers arising, Speculum. from a continual Multiplication by a+e (=z)come out thus with their adjoyned Co-efficients. See the Table. But now, if it be  $a_e = z_i$ the Table is compos'd of the fame Members, only the odd Powers of e, as e, e3, e5, e7 are Negative, and the even Powers, as e', e', e', Affirmative. Also let the fum of the Co-efficients of the fide  $e_1$  be = s; the fum of the Co-efficients of the Square ee = t, the fum of the Co-efficint of  $e^{T} = u$ ; of  $e^{t} = w$ ; of  $e^{5} = x$ ; of  $e^6 = y$ , &c. But now, fince e is supposed only a small part of the Root that is to be enquir'd, all the Powers of e, will be much lefs than the correspondent Powers of a, and fo far the

the first Hypothesis; all the superior ones may be rejected; and forming a new Equation, by substituting a+e=z, we shall have (as was faid)  $\pm b=\pm e+e$ . The following Examples will make this more clear.

### Example I.

Let the Equation  $z^4-3z^2+75z=10000$ , be propos'd. For the first Hypothesis, let a=10, and fo we have this Equation,

 $z^{4} = [-a^{4} + a^{3}e^{+e}a^{3}e^{2} + a^{2}e^{3} + a^{2}e^{4} - dz^{2} = -da^{2}dae - de^{2} + (-dz^{2} - dz^{2}) - (-dz^{2} - dz^{2}) - (-dz^{2} - dz^{2}) - (-dz^{2}) - (-dz^{2})$ 

$$+450 4015e + 597ee - 40e^3 + e^2 = 0$$

The Signs  $-1^-$  and - with refpect to the Quantities e and  $e^3$ , are left as Doubtful, till it be known whether e be Negative or Affirmative; which thing creates fome difficulty, fince that in Equations that have feveral Roots, the Homogenea Comparationis (as they term them) are oftentimes encreafed by the minute quantity a, and on the contrary, that being increafed, they are diminified. But the Sign of e is determin'd from the Sign of the Quantity b. For taking away the Refolvend from the Homogeneal formed of a; the Sign of se (and confequently of the prevailing Parts in the composition of it) will always be contrary

trary to the Sign of the difference b. Whence 'twill be plain, whether it must be -1 e, or -e; and confequently whether a be taken greater or lefs than the *True Root*. Now the quantity e is  $=\frac{1}{2}s - \sqrt{\frac{1}{2}ss - bt}$ , when b and t have the fame Sign, but when the Signs are different, e is  $= \sqrt{\frac{1}{2}ss - bt} - \frac{1}{2}s$ . But after it is t

found that it will be -e, let the Powers  $e, e^3$ , and  $e^5$ ,  $\mathfrak{S}^{\circ}c$ . in the affirmative Members of the Equation be made Negative, and in the Negative be made Affirmative; that is, let them be written with the contrary Sign. On the other hand, if it be +e (let those foremention'd Powers) be made Affirmative in the Affirmative, and Negative in the Negative Members of the Equation.

Now we have in this Example of ours, 10450 inftead of the Refolvend 10000, or b=+450, whence it's plain, that *a* is taken greater than the Truth, and confequently, that 'tis—*e*. Hence the Equation comes to be, 10450-4015e+ 597ee-4e³+e⁴ = 10000. That is, 450-4015e +597ee = 0; and fo 450 = 4015 e -597ee,

or  $b \equiv s e = tee$ , whofe Root  $e \equiv \frac{1}{2}s = \sqrt{\frac{1}{2}ss - bt}$ 

or  $\frac{s}{2t} - \sqrt{\frac{ss}{4tt}} \frac{b}{t}$ ; that is in the prefent cafe,  $t = \frac{2007!}{5} - \sqrt{376!406!}$ , from whence we have

597 the Root fought, 9, 886, which is near the Truth. But then substituting this for a fecond

Distant by Google

cond Supposition, there comes a+e=z, most accurately 9, 8862603936495.... fcarce exceeding the Truth by 2 in the last Figure, viz. when  $\sqrt{\frac{1}{4}} sr(-bt - \frac{1}{4}s = e$ . And this (if need be) tmay be yet much farther verified, by fubstracting (if it be +e) the quantity  $\frac{1}{\sqrt{\frac{1}{4}}sr(-1)}e^{\frac{4}{2}}$ , from the Root before found; or (it it be -e) by adding  $\frac{\frac{1}{2}ue^3 - \frac{1}{2}e^{\frac{4}{2}}}{\sqrt{\frac{1}{4}}sr(-1)}$  to that Root. Which Compendium is fo much the more Valuable, in that fometimes from the first Supposition alone, but always from the fecond, a Man may continue the Calculus (keeping the fame Coefficients) as far as he pleafes. It may be noted, that the fore-mentioned Equation, has alfo a Negative Root, viz. z = 10, 26....

which any one that has a mind, may determin more accurately.

### Example II.

Suppose  $z^3 - i7z + 54z = 350$ , and let a = i0. Then according to the prefeript of the Rule,

 $-\frac{1}{2} = a^{3} - \frac{1}{3}a^{2}e + 3ae^{2} - \frac{1}{6}e^{2}$  $-dz^{2} = da^{2} - 2dae - de^{2}$  $+cz = ca - \frac{1}{6}e^{2}$ 

G 2

That

### Miscellanea Curiosa. 84 That is, +1000+300e+30e2+e3 -1700-340e-17e2 7-540-1-540 -350

Or, -510+14e-1-13ee-1-e'= 0. Now, fince we have -510, it is plain, that a is affumed less than the Truth, and confequently that e is Affirmative. And from (the Equation)  $510 = 14e + 13e^2$ , comes  $e = \sqrt{bt - \frac{1}{4}ss - \frac{1}{4}s}$ 

 $\sqrt{6679-7}$ . Whence z = 15, 7..., which is 13 too much, because of a taken wide; therefore Secondly, let a=15, and by the like way of Reafoning, we fhall find  $e = \frac{1}{2}s - \sqrt{\frac{1}{4}ss - tb}$ 

 $=\frac{109^{1}_{2}-\sqrt{11710_{4}}}{28}$  and confequently z=14,954068. If the Operation were to be repeated the third time, the Root will be found conformable to the Truth as far as the 25th Figure ; but he that is contented with fewer, by writing tb 1-te' inftead of tb, or fubitra-1 23 Ating or adding  $\sqrt{\frac{2c}{4ss-tb}}$  to the Root before found, will presently obtain his end. Note, the Equation proposed, is not explicable by any other Root, because the Refolvend 350, is greater than the Cube of  $\frac{17}{3}$ , or  $\frac{d}{3}$ 

Example

### Example III.

Let us take the Equation  $z^4 - 80 z^3 + 1998$  $z^2 - 14937 z + 5000 = 0$ , which Dr. Wallis ules Cap. 62 of his Algebra, in the Refolution of a very difficult Arithmetical Problem where by Vieta's Method he has obtain'd the root most accurately; and Mr. Raphfon brings it alfo as an Example of his Method, Page 25, 26. Now this Equation is of the form, which may have feveral Affirmative Roots, and (which increafes the difficulty) the Coefficients are very great in respect of the Resolvend given.

But that it may be the cafier manag'd, let it be divided, and according to the known Rules of Pointing, let - z4+8z3- $20 z^2$  1 15 z = 0.5 (where the quantity z is  $\frac{1}{2}$  of z in the Equation proposed) and for the first Supposition, let a = 1. Then  $+^2 - 5e -$ 2e2-1-4e3-e4-0,5=0; that is, 11= 5e-1-2ee; hence  $e = \sqrt{\frac{1}{4}ss + bt - \frac{1}{2}s}$  is  $= \sqrt{37-5}$ , and fo

z=1, 27; Whence'tis manifest that 12, 7 is near the true Root of the Equation proposed. Now Secondly, let us suppose z = 12, 7, and then according to the directions of the Table of Powers, there arifes

 $-26014, 4641-8193, 532e-967, 74e^2-50, 8e^3-e^4$ +163870, 640-38709, 60e+3048  $e^2$ -80  $e^3$ e 2 -322257, 42 -50749, 2 e-1998 - 5000. That

G 3

86

That is, +298, 6559-5296 132 e+82,  $26e^{2}$ +-29,  $2e^{3}-e^{4}=0$ ; And fo -298, 6559=-5296,  $132e^{-1}82$ , 26ee, whole Root e (accord-

ing to the Rule) =  $\frac{1}{2}s - \sqrt{\frac{1}{4}ss - bt}$  comes to

$$\frac{2648,066-\sqrt{6987686},106022}{82,26}$$

,05644080331.... = e lefs than the Truth. But that it may be corrected, 'tis to be confider'd

that  $\frac{1}{\sqrt{\frac{1}{4}ss-bt}}$ , or  $\frac{0026201....}{2643,423...}$  is,0000099,

and confequently e corrected, is = 0564470448. And if you defire yet more Figures of the Root, from the e corrected let there be made  $tue^{3}$ --re⁴ = 0, 43105602423..., and

 $\frac{1}{2}s - \sqrt{\frac{1}{4}ss - bt - tue^3} + te^4}$  or which is all one,

2648, 066--/6987685, 67496597577 ....

#### 82, 26

, 05644179448074402 = e; whence a+e=zthe Root is most accurately 12, 75644179448 074402... as Dr. Wallis found in the forementioned Place; where it may be observ'd, that the repetition of the Calculus does ever triple the true Figures in the affumed a, which the first correction, or  $\frac{1}{\sqrt{4}} \frac{ve^2 - \frac{1}{2}}{\sqrt{4}} \frac{e^4}{\sqrt{4}}$  does quintuple; which is also commodiously done by the Logarithms. But the other Correction after

aiter

after the first, does also double the number of Figures, so that it renders the assumed altogether Seven-fold; yet the first Correction is abundantly sufficient for Arithmetical uses, for the most part.

But as to what is faid concerning the number of Places rightly taken in the Root, I would have underftood fo, that when a is but ¹ part diftant from the true Root, then the first Figure is rightly assumed; if it be within  $r_{00}$  part, then the two first Figures are rightly assumed; if within  $r_{00}$ , and then the three first are fo; which confequently manag'd according to our Rule, do presently become nine Figures.

It remains now that I add fomething concerning our *Rational Formula*, viz.  $e = \frac{sb}{ss \pm tb}$ which feems expeditious enough, and is not much Inferior to the former fince it will triple the given Number of Places. Now

having formed an Equation from a le =z, as before, it will prefently appear, whether a be taken greater or lefs than the Truth ; fince se ought always to have a Sign contrary to the Sign of the difference of the Refolvend, and its Homogeneal produced from a. Then fuppofing -t-b-t-se-t-a-tee = 0, the Divisor is ss-tb, as often as t and b have the fame Signs; but it is ss+bt, when they have different ones. But it feems most commodious for Practice, to write the Theorem thus,  $e = \frac{b}{t} + \frac{tb}{t}$  fince this way the thing is done by one Multiplication and two Divisions, which otherwife would require three Multiplications, and one Divifion. A G 4 Lot

88

Let us take now one Example of this Method, from the Root (of the foremention'd Equation) 12, 7...., where

298, 6559 - 5296, 132e+82, 26ee+29, +b - s + t + u 2e³ - e⁴ = 0, and fo  $\frac{b}{s} - \frac{tb}{s} = e$ ; that is, let it be as s to t, fo b to  $\frac{tb}{s} = 5296, 132)298$ , 6559 into 82,26 (4, 63875... wherefore the Divifor is  $s - \frac{tb}{s} = 5291, 49325...)298$ , 6559 (0, 056441....=e, that is, to five true Figures, added to the Boot that was taken.

Figures, added to the Root that was taken. But this Formula cannot be corrected, as the foregoing Irrational one was; and fo if more Figures of the Root are defired, 'tis the beft to make a new Supposition, and repeat the Calculus again: And then a new Quotient, tripling the known Figures of the Root, will abundantly fatisfie even the most Scrupulous.

A Difer-

A Differtation concerning the Conftruction of Solid Problems, or Equations of the third or fourth Power, by the help of one (given) Parabola and a Circle.

By Edmund Halley.

89

H Ow all Equations (that involve the third or fourth Power of the unknown Quantity) may be conftructed by the help of any given Parabola and a Circle, the Famous M. Des Cartes has fhewn'and clearly demonstrated in the Third Book of his Geometry. But he first of all orders the second Term of the Equation (if it be there) to be thrown out, and then by the Rule there delivered, to find the Roots of the Equation fo reduced.

And fince that Operation feems too Laborious, fome thought fit to invent a like Conftruction, without any previous Reduction. Amongft whom Francis a Schooten has offer'd a Method (for conftructing cubical Equations howfoever affected) which might have been called very easie and fimple; if (by unfolding the Principle from whence he deduced his Rule) he had better confulted his Reader's. Memory, which he burdens with very many and perplex'd Cautions. But lately our Famous Countriman, Mr. Thomas Baker, in a whole Treatife written upon these Constructions.

90

Aions, has comprehended not only all Cubical, but also Biquadratical Equations of every kind, under one General Rule, which he has demonstrated, and abundantly Illustrated with Examples through all Cafes; and moreover at the Clofe, propos'd a way, by which that General Rule might be Investigated. But he does not fhew the very Method, by the help of which (as I fuspect) he obtain'd his Univer fal Geometrical Clavis, or at least might have obtain'd it with much more eafe. And fince this Rule of Baker's is no lefs perplex'd with Cautions about the Signs 1- and - than Schooten's is, fo that a Perfon can hardly perform those Constructions aright, without he has the Book by him; I thought that it wou'd not be either Unpleafant or Unprofitable to young Students, to explain the Foundations of both Rules, and by fome emendation of the Method once more, to afford as much light as I cou'd in fo difficult a Matter. Cartefius's Construction (which does very eafily discover the Roots of all Cubick or Biquadratick Equations, where the fecond Term is wanting) may be fuppos'd as known. Yet fince 'tis the main bottom, on which all that follows does depend; that this Differtation may not feem to want a principal Part, I'll here add the Rule taken out of his Geometry, altering fome few things (as I think) for the better.

The fecond Term being out of the Equation; all cubical Equations, are reduced to this Form,  $z^3 * apz. aaq = o$ ; and Biquadratical ones to this Form,  $z^4 * apz. aaqz.$  $a^2r = o$ , where a denotes the Latus Restum of any given Parabola, which is used in the Con-

91

Construction. Or else taking a for Unity, those Equations are reduced to these Forms. viz. Cub. z3. *p z. q=0, and Biguadr. z4 x.pzz. gz.r. = o. Now the Parabola FAG, Fig. 9 being given, whofe Axis is ACDKL, and Parameter = a or i; let AC be taken =  $\frac{1}{2}a$ , and be fet off always from the Vertex A, towards the inner parts of the Figure. Then take CD p, in that Line AC, continued towards C, if it be --- p in the Equation, or towards the contrary Point, if it be +p. Farther, from the point D (or from the point C, if the quantity p be not in the Equation) Let DE (crected perpendicular to the Axis) be made = 19 which is to be fet to the right hand if it be ---- , but to the other fide of the Axis if it be 1-q. And then a Circle defcribed on the Center E, which the Radius AE (if the Equation be but a Cubical one) will interfect the Parabola in as many Points (viz. F, G, G,) as the Equation has True Roots, of which the Affirmative ones, as GK, shall on the right fide of the Axis, and the Negative ones as FL, on the Left. But if the Equation be a Biquadratical one, then the Radius of the Circle AE, by adding (if it be ---- r) or Substract-Ing (if it be -)-r) from the Square of it, the Rect-angle a xr, or the content under the Parameter, and the given Quantityr; which is very eafily done Geometrically. And the Intersections of this Circle with the Parabola. will give (letting fall Perpendiculars from thence to the Axis) all the true Roots of the Biquadratical Equation; the Affirma. tive ones being on the Right fide of the Axis. and the Negative ones, on the Left. The dedemon-

92

monftration of all which I leave to Cartefius the Inventor. Let it be Noted, that I endeavour here that the Affirmative Roots, may always be had on the Right fide of the Axis, to avoid the Confusion that will neceffarily arife from a multitude of Cautions, where the reason of them is not evident.

Having premifed these things, in order to make way for the construction of these Equations, even when the second Term is found in them, we are to consider the Rule it self for taking away the second Term, and reducing the Equation to another, such as might be constructed by the foregoing Method. Now all Cubick Equations of this Classis, are reduced to this form,  $z^3 bzz$ . apz aaq = 0, or to this,  $z^3$ .  $bz^2$ . z. aaq = 0. Biquadratick ones may be reduc'd to this,  $z^4 . bz^3 . apz^2 . aaqz$ .

 $a^{3}r \equiv 0$ , or this,  $z^{4}$ , or this  $z^{4}$ .  $bz^{3}$ . *. aaqz.  $a^{3}r \equiv 0$ , or this,  $z^{+}$ .  $bz^{3}$ .  $apz^{2}$ . *.  $a^{3}r \equiv 0$ , or laftly, to this Form,  $z^{4}$ .  $bz^{3}$  *. *.  $a^{3}r \equiv 0$ . From all which there arifes a great Variety, according as the Signs  $\neg$  or - are diverfly connected together; and hence the General Rule ferving all these cases, is rendred very obscure and difficult, unless (manag'd by the help of the following Method) it be cleared up and delivered from those Intricacies.

The fecond Term in Biquadratical Equations, is taken away by putting  $x = z + \frac{1}{4}b$ , if it be  $-\frac{1}{b}$  in the Equation; or  $x = z - \frac{1}{4}b$ , if it be -b. Hence  $x - \frac{1}{4}b$  in the first case, and x- $\frac{1}{4}b$  in the fecond, is = z; and fo in any Equation proposed, substituting instead of z, its Equal, there will come forth a new Equation, wanting the fecond Term, all whose Roots x do

do exceed, or come fhort of the fought Root z, by the given difference  $\frac{1}{4}b$ . But fince in things of this kind, Examples do more than Precepts, let us propose one or two Equations to be constructed.

### Example I.

 $z^{4} - bz^{3} - apz_{7}^{2} - aaq_{7}^{2} - a^{3}r \equiv 0.$ put  $x - -\frac{1}{4}b \equiv 2^{2}$ , and then will  $x^{2} - -\frac{1}{5}bx - \frac{1}{6}bb \equiv 2^{2}.$   $x^{3} - -\frac{3}{4}x^{2}b + \frac{3}{16}xb^{2} - -\frac{1}{64}b^{3} \equiv 2^{3}.$ and  $x^{4} - -bx^{3} - \frac{1}{8}b^{2}x^{2} - -\frac{1}{16}b^{3}x - \frac{1}{256}b^{4} \equiv .2^{4}$ 

### Hence it follows, that

$$x^{4}--bx^{3}+\frac{1}{8}b^{2}x^{2}--\frac{1}{2}i_{5}b^{3}x+\frac{1}{7},\frac{1}{5}b^{5}b^{2}=\frac{1}{2}a^{2}.$$
  

$$bx^{3}--\frac{3}{7}b^{2}x^{2}+\frac{1}{3}b^{3}x--\frac{1}{6}a^{b}+\frac{1}{2}+bz^{3}$$
  

$$---apx^{2}+\frac{1}{2}apbx--\frac{1}{16}apb^{2}=--apz^{2}$$
  

$$--a^{2}qx+\frac{1}{7}a^{2}qb=-a^{2}qz$$
  

$$+a^{3}r$$

The Sum of all thefe is a new Equation wanting the fecond Term, and which confequently may be conftructed by *Cartes's* Rule, by taking inftead of  $\frac{1}{2}p$ , half the Coefficient of the third Term, divided by *a* or the Parameter, that is  $-\frac{3}{12}\frac{bb}{a}-\frac{1}{2}p$ ; and inftead of  $\frac{1}{2}q$ ; half the Coefficient of the fourth Term, divided by *aa*, that is,  $+\frac{1}{12}\frac{b^3}{a^2}+\frac{1}{4}\frac{pb}{a}\frac{1}{2}q$ . The

Mem-

92

94

Members of which that have the Sign -l- are to be fet off to the left Hand from the Axis, and those that have the Sign - to the Right; in order to find the Center of the Circle required for Construction, whose Interfections with the Parabola (letting fall perpendiculars to the Axis) may give all the true Roots x, namely, the Affirmative ones on the Right fide of the Axis, and the Negative ones on the Left. But now, when x - 1b = z, then drawing a Line Parallel to the Axis on the right fide of it, and at the distance of 1/2, the Perpendiculars terminated on this Parallel; will denote all the enquired Roots z, the Affirmative ones on the right fide, and the Negative ones on the Left. As for what relates to the Radius of the Circle, it is had, by adding the Negative, or taking away the Affirmative parts of the fifth Term divided by aa, from the Square of the Line AE, drawn from the Center E found, to A the Vortex of the Parabola; which is mostly done, by taking iustead of AE, the Line EO which is terminated at 0 the Intersection of the Para2 bola, and the fore-mentioned Parallel; for the Square of this comprehends all the parts of the fifth Term, brought into the new Equation upon the cafting out of the fecond Term; as is eafily proved : And it remains only, that the square of EO be increased, if it be ---in the Equation, or diminish'd, if it be +r, by the addition or fubstraction of the Rectangle # r; from whence the Radius of the Circle defired, is compos'd. This Method of inveftigating M. Baker's central Rule, is easie and free from all Cautions; and the difference arifes onh

Dibilized by Google

ly from hence, that I determin the center of the Circle, by the Axis, and he by a Parallel to the Axis; and that I always have four Affirmative Roots on the right fide the Axis, which he has fometimes on the right fide, and fometimes on the left.

As for cubical Equations, they are to be reduc'd to Biquadratical ones, before they can be conftructed by the fame General Rule; which is done by multiplying the Equation propos'd by its Root z, whence arifes a Biquadratick Equation, in which the laft Term or r, is wanting. Wherefore taking away the fecond Term, and finding the Center E, the line EO is the Radius of the Circle; viz. When ar is  $\equiv o$ , and the whole fifth Term in the new Equation, arifes from the taking away of the fecond Term. Let this Equation be propos'd to be conftructed.

#### Example II.

 $Z^{3} - bz^{2} + apz + aaq = o,$ which multiply'd into z, becomes  $z^{4} bz^{3} + apz^{3} + aaqz = o.$ To take away the fecond Term, put  $x - \left\{ -\frac{1}{4} b = z, \text{ and then will} \right\}$  $x^{4} + bx^{3} + \frac{2}{3} bbx^{2} + \frac{1}{12} b^{3}x + \frac{1}{12} b^{2} = -bz^{3}$  $-bx^{3} - \frac{3}{4} b^{2}x^{2} - \frac{3}{5} b^{3}x - \frac{5}{4} b^{4} = -bz^{3}$  $+apx^{2} + \frac{1}{4} abpx + \frac{1}{12} apb^{2} = -bapz^{4}$ 

 $taag x + \frac{1}{2}aagb = taagz.$ Now in this new Equation, the half Coef-

ficient (of the third Term) divided by a, viz- $\frac{3bb}{16a}$  +  $\frac{1}{2}p$ , is to be used instead of  $\frac{1}{2}p$ ; and the

96

the half Coefficient of the (fourth Term) divided by a a, the Square of the Latus Rectum,  $b^3$   $b^b$ ,  $b^3$  $\frac{b^3}{16a^2} + \frac{pb}{4a} + \frac{i}{2}q$ , is inftead of  $\frac{i}{2}q$  in Cartefius's Construction, from whence the Center E is determin'd. Then drawing a Parallel to the Axis, at the distance 1b, to the left fide (because of z = x + 4b) whose Intersection with the Parabola, let be 0; a Circle described on the Center E with the Radius EO, will cut or touch the Parabola in as many Points as the Equation has true Roots, which Roots, or z, are the Perpendiculars let fall from those Points upon the Parallel to the Axis, the Affirmative ones to the Right fide, and the Negatives to the Left. If the third or fourth Term.or both, be wanting in the Equation, there's no difference at all ( of the Method of iavestigating the Central Rule) to be observ'd. But the Quantity p or q being wanting, those parts of the Lines CD and DE (in fome manner deduced from that Quantity) will be wanting too, and we are to proceed with the other Coefficients of the third and fourth Term in the new Equation, according to the way prefcrib'd in the foregoing Examples.

Hitherto we have confider'd Mr. Baker's General Method, than which none more Easie and Expeditious is to be expected, using either a Parabola, or any other Curve for a Construction, viz. when the Equation rifes to the Biquadratick Power. For while I am writing of this, 'tis my good Luck to hit upon a certain Geometrick Effection of the central Rule, which is Expeditious beyond Hope, and will abun-

abundantly fatisfy those that are curious in these Matters.

97

(Fig. 10.) Having describ'd the Parabola NAM, whofe Vertex is A, Axis ABC, and Parameter a; let the Equation be reduced to this Form,  $z^4$ .  $bz^3$ .  $apz^3$ . aaqz.  $a^3r$ . = 0; or if it be only a Cubical one, to this,  $z^3$ .  $bz_2$ . *apz. aaq.* = 0. Then at the diftance BD =  $\frac{1}{2}b$ , let DH be drawn parallel to the Axis (to the Left Hand if it be -b, and to the Right, if it be-1-b) meeting the Parabola in the point D, from whence let fall BD perpendicular to the Axis. In the Line AB continued towards B, make BK = a, and draw the Line DK interminate on either fide. Farther, take K C=2 A B, always in the Axis produced beyond K; and if the quantity p has the Sign -, take towards the fame parts, C E=_2'p, but towards the contrary part, if it be +p. Then at the point E (but at the point C if the quantity be wanting) erect EF perpendicular to the Axis, meeting (if need be) the Line DK produced, in the point  $F_2$ , which point is the Center of the Circle required, if the quantity q be wanting: But if q be in the Equation, then we must take in the Line FE (if need be) produced the length of  $FG = \frac{1}{2}q$ , which place to the Left. Hand if it be  $\pm q$ , but to the Right if it be -q; and then the point G will be the Center of the Circle required for the Construction, and the Radius of it, will be the Line G D, if the quantity r be wanting, that is, if the Equation be only a Cubical one; the Square of which fame Line (in Biquadratick Equations) is to be encreased H bv

by the addition of the Rectangle under rand the Latus Rectum, if it be -r, or to be diminified by the fame Rectangle if it be +r. The Circle thus defcrib'd, and Perpendiculars let fall from its Interfections with the Parabola, to the Line DH, those that are at the Left Hand as NO will always be the Negative Roots of the Equation, and those at the Right, the Affirmative.

Cubick Equations are otherwife (and fomething more fimply) conftructed according to Schooten's Rule, in which alfo the Roots refpect the Axis. But becaufe the Inventor himfelf does neither explain the Inveftigation nor Demonstration, it will not be amifs to fhew the Foundation of it here, and at the fame time render the Geometrick Conftruction more Elegant, and rid it of those *Cautions* in which 'tis involv'd.

This Rule is deriv'd from hence, that every Cubick Equation may be reduced to a Biquadratick one, in which the fecond Term is wanting. Which is done, by multiplying the Equation proposed into z - b = 0, if it be -b in the Equation, or into z-b=0, if it be -b; and the new Equation thus form'd will have the fame Roots with the Cubical one, and moreover another Equal to -b, if it be -b in the Equation, or contrariwife.

Let the Equation  $z^3 - z^2 b + apz + aaq = 0$ , be proposed to be constructed; multiply this into z+b, and it makes

z* - z * b-1-apz *-1-aagz

-1-23b-bb23+abpz-faagb

Here now the fecond Term is wanting, and the

98

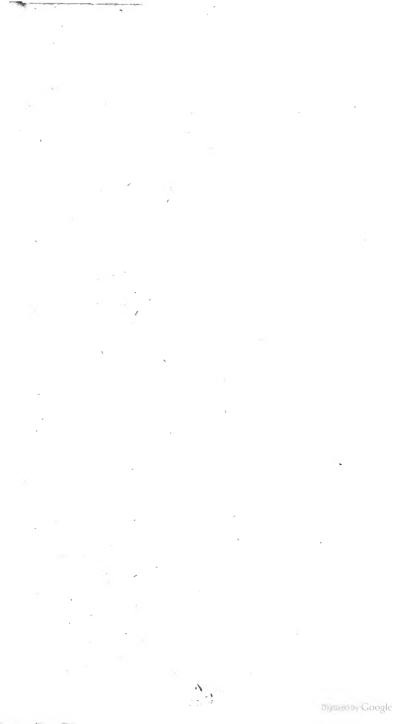
99 the Coefficient of the third Term - bb+ap, gives  $-\frac{bb}{2a} + \frac{b}{2}p$ , in the room of  $\frac{b}{2}p$  or CD in Cartefius's Construction; and from half the Coefficient of the fourth Term is made  $\frac{bp}{2}q + \frac{bp}{2}$ , inftead of  $\frac{1}{2}q$  or DE, and fo the Center of the Circle fought is determin'd. Alfo becaufe one of the Roots of the new Equation, viz.  $\pm b$  is given, a point in the Circumference will be given too, and confequently the Radius. Laftly, Having defcrib'd the Circle, Perpendiculars let fall from its Interfections with the Parabola, to the Axis, will give the Roots of the Equation, both Affirmative and Negative, in the fame manner as before.

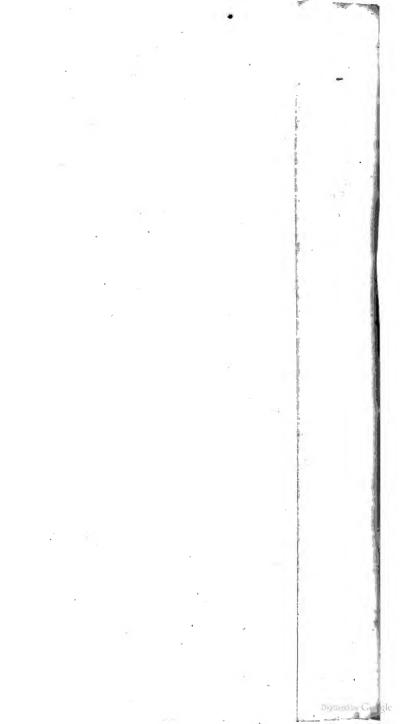
Now the Center of the Circle is found by a most easy Construction, and which is to be preferr'd to all others, in Cubick Equations.

Fig. 11. Let A be the Vertex, and AF the Axis of the defcrib'd Parabola AMD; at a distance equal to b let DK be drawn parallel to the Axis, to the Right Hand if it be-1-6 in the Equation, and to the Left, if it be -b; which Line fuppofe to meet the Para-bola in the point D. Upon the Centers D and A, and with equal Radij, defcribe on both fides two occult Arches, interfecting one another, and thro' those points of Interfection draw the interminate Line BC which cuts the imaginary Line AD in the middle and at right Angles, and meets the Hž Axis

Axis in E. From E fet off  $EF = \frac{1}{2}p$ , downwards, if it be - p in the Equation, but upwards towards A if it be +p; then at the point F( or E, if p be wanting) erect the Perpendicular FG, meeting the Line BC in G, and in G F produced take G H  $= \frac{1}{2}q$ , to the Right Hand, if it be -q in the Equation, but to the Left, if -1-q. Then will the point H be the Center, and HD the Radius of the Circle fought, which (letting fall Perpendiculars to the Axis from its Interfections with the Parabola) will fhew all the Roots (as LM) of the Equation. And how this Construction follows from what went before, is evident enough of it felf, fo that there is no need of infifting any farther upon the Demonstration of it.

A Dif.





A Discourse concerning the Number of Roots, in Solid and Biquadratical Equations, as also of the Limits of them. By E. Halley.

Aving in the precedent Difcourfe fhewn a Method, by which folid Problems however affected, might be conftructed after a moft fimple and eafy manner, by the help of one given Parabola and a Circle; towards the latter end a certain pleafant Speculation offer'd it felf, namely, that from thefe Conftructions, the Number of Roots in any Equation, with their Limits and Sines, would eafily follow and be determin'd. Upon which account, I promis'd that I would quickly write a fhort Differtation concerning this Subject, in which I was perfwaded I fhou'd perform fomething not unprofitable nor ungrateful (if not to the Geometers of the firft, yet at leaft) to thofe of the fecond Rank.

But coming to look nearer into the Bufinefs, I found I was imprudently fallen in among fome of the profound Difficulties of Geometry, and deftin'd to handle the fame Things, that formerly employ'd the Pains of two Illuftrious Men, Harriot and Cartes; H 3 102

in which they either of them (by a like Fate tho' in a different way) committed a Paralogifm, perhaps the only one in all their Geometrical Writings; as shall be afterwards prov'd. Wherefore being fenfible, as well of the Difficulty, as the Excellency of the Subject, I refolv'd to apply to it ftrenuoufly, that I might not be thought unable to perform my Promises, and that so 'noble part of Geometry, and fo little cultivated, might not lie any longer wrapt up in Darknefs, but be render'd plain and intelligible by these few Lucubrations of mine. But first the Reader must take notice, that while he fets to the Reading of this, he ought to have the foremention'd Differtation (No. 188.) at hand by him, and to understand the Conftructions there delivered very well; becaufe those things that follow do chiefly depend upon them, neither are they to be here repeated again.

It is plain from Cartefius, and what was there faid, that both in Cubick and Biquadratick Equations, the Roots may be expounded by Perpendiculars let fall, upon the Axis or given Diameter of the given Parabola, from the Interfections of that Curve with a Circle. And whereas when a Circle interfects a Parabola, it must necessarily do fo, either in four or in two points; it's manifeft; that in Biquadraticks there must always be, either two or four true Roots, Affirmative or Negative; as also if the Circle happens to touch it, in which cafe the equality of two Roots of the fame Sign, is concluded. But in Cubick Equations, because one of the Interfections

terfections is requir'd to the Conftruction, therefore either but one, or the three remaining Roots, do denote one or three; as in the cafe of Contact; whence its plain, that there are found two equal Roots, and that the Problem from whence the Equation refults, is really *Plain*.

Therefore all *Cubick* Equations however affected, are explicable by one, or by three Roots, which are always *poffible*, that is, if we admit *Negative* Roots for *true* ones. So Biquadraticks whofe last Term r is affected with the Sign -, are explicable by two or four; but if it be -1-r in the Equation, and

it be fo great that  $\sqrt{GD} q - ar$  (See Fig. 10.) be lefs than that the Circle defcrib'd with that Radius and on the Center G, can touch the Parabole in any point; the given Equation is altogether impossible, nor is it explicable by any Affirmative or Negative Root; but more of this in the following Pages.

Now fince there is fo great a difference between the Cafes of Cubick and Biquadratick Equations, that they cannot be comprehended together, we will first of all handle the Cubicks, and then the others. The Cubicks are constructed by an infinite Number of Circles in a given Parabola; but the Biquadraticks by one alone (at least by these Methods) and that because, putting z---e (or any Indeterminate) equal to nothing, the Cubick Equation is reduced to a Biqudratick having the fame Roots with the Cubick, and befides that, another Root equal to e; whence it comes to pass that the Cubick Equation  $H_4$ may may be constructed by as many different Circles, as you can imagine Quantities e, that is, an infinite Number. But among all thefe, that which I gave before, is the eafieft. Yet there is another not much inferior to this, which feems better accommodated to the defigns of determining the Number of the Roots, and their Limits; and which arifes from the taking away of the fecond Term, by putting after the common way x = z-- or - i of the Coefficient of the fecond Term. Now this way is thus : The Parabola ABY (Fig. 12.) being given, whole Vertex is A, its Axis AE, and Latus Rectum a, let the Equation be reduced to the ufual Form, viz.  $z^3$ .  $bz^2$ . apz. aag. = 0. Then at the diftance of ib let there be drawn BK (parallel to the Axis, to the Right Hand if it - - b, otherwife to the Left) which meets the Parabola in B; and let the Line DP interminate on both fides, be erected perpendicular to the fuppos'd Line AB, meeting the Axis in the point G. From the point B, let fall the Perpendicular BC, and let GE be always made equal to AC, and be fet off towards the lower parts. From E fet off EH = : p, upwards if it be -p in the Equation, but downwards if -p; and from the point H, (or E, if the quantity p be wanting) let the Perpendicular HQ be drawn out, meeting the interminate Line DP in O. Laftly, in the interminate Line HQ, make  $OR = \frac{1}{2} q$ , from 0 to the Right Hand, if it be - 9, but to the Left, if + q. Then a Circle describ'd on the Center R with the Radius RA, will cut the Parabola in as many points, as the Equation

Dialized by Google

Equation propos'd has Roots, and they will be the Perpendiculars ZY, let fall from the Interfections  $\Upsilon$ , to the Line BK parallel to the Axis; of which those that are to the *Right* Hand of the Line BK, are the Affirmative ones, and those to the Left, the Nergative.

The conveniency of this Construction, lies in this, that 'tis perform'd by a Circle paffing thro' the Vertex, in the fame manner as if the fecond Term had been wanting. And therefore to determine the Number of the Roots, 'tis fufficient to know the Properties of the Place, or that Curve Line which diftinguishes the Spaces, in which if the Center of the Circle (that passes thro' the Vertex of the Parabola) be placed, the Circumference of it shall interfect the Parabola either in one or in three other points : That is, to define the Nature of that Curve, in which, fall the Centers of all the Circles paffing thro' the Vertex, and then touching the Parabola. Now this Locus, is that very Paraboloid, which the celebrated Dr. Wallis calls the Semicubical, in which the Cubes of the Ordinates are as the Squares of the correspondent Abscisses. The Latus Restum of which, is #3 of the Latus Rectum of the given Parabola, and its Vertex the point U (Fig. 12.) the Line AU being half the Latus Rectum of the fame Parabola. That is, if we put unity for the Latus Rectum of the gi-ven Parabola, then 27 of the Cube of the ordinate applicate, will = the Square of the intercepted Diameter; or the Cube of # VH = the Square of AR, viz. if R be the Center 2

Center of the Circle that passes thro' the Vertex of the Parabola, and touches the fame afterwards.

106

This is that Curve which our Countryman Mr. Neil (the first of all Mortals) demonstrated to be equal to a given right Line, and by that means obtain'd a Reputation among the principal Geometricians. Its properties have been curioufly enquired into, by Dr. Wallis, (at the end of his Book of the Ciffoid) and Hugenius (Prop. 8 & 9. of his Tract of the Evolution of Curve Lines) and others, whofe Writings the Reader may confult. This Curve describ'd on either fide of the Axis of the Parabola (viz. VNL, VPX) comprehends a Space, in which if the Center of the Circle (which paffes thro' the Vertex A) be placed, it will cut the Parabola in three other points. But the Spaces more remote from the Axis, do afford Centers for Circles that will cut the Parabola but in one point befides the Vertex.

Thefe things well underftood, we are now prepar'd to determine the Number of the Roots. And first of all, let the fecond Term be wanting, and let the Latus Restum = 1, or  $AV = \frac{1}{2}$ . In the Construction VH is  $= \frac{1}{2}p$ , HR  $= \frac{1}{2}q$ ; and fince if it be +p in the Equation,  $\frac{1}{2}p$  is to be fet off from  $\mathcal{O}$  towards the upper parts; the Center of the Circle is always found without the Space LVX, and therefore is explicable by one Root only, which is Affirmative if it be -q, Negative if -q; and these Roots may be investigated by Cardan's Rules. But if it be -p, then  $UH = \frac{1}{2}p$ , is fet off towards the lower parts; and it is possible

107

possible that HR may fall between the Axis and the Curve UX or UL, viz. if the Cube of  $\frac{4}{3}$ UH or of p, be greater than the square of  $\frac{1}{2}q$ ; that is, if  $\frac{1}{27}p^{\frac{3}{2}}$  be greater than  $\frac{4}{4}q^2$ ; in which cafe there are three Roots, two Negative, if it be -q, and one Affirmative equal to the fum of the others; but if it be -1-q, then there are two Affirmative ones, and one Negative. But if  $\frac{1}{27}p^3$  be lefs than  $\frac{1}{4}q^2$ , then there is but one Root, Affirmative if it be -q, Negative, if -q. All which things are taught by those that have handled this part of Geometry.

Now let all the Terms be in, and first let there be proposed, as an Example, this Equation,  $z^3 - z^2 b - zp - q = o$ , to which Fig.12. ferves. In the Construction of this, we have BC =  $\frac{1}{5}b$ , UG =  $\frac{1}{5}$  AC =  $\frac{1}{5}b^2$ , UE =  $\frac{1}{16}b^2$ , UH =  $\frac{1}{5}b^2 - \frac{1}{2}p$ , GH =  $\frac{1}{5}b^2 - \frac{1}{2}p$ , or  $\frac{1}{5}p - \frac{1}{5}b^2$ . Hence HO =  $\frac{1}{27}b^3 - \frac{1}{5}bp$ , or  $\frac{1}{5}bp - \frac{1}{5}bp$ .  $\frac{1}{27}b^3$ , and HR (that is the diftance of the Center of the Circle R from the Axis) is ever the difference between & bp and 17 b3 + 1 q, which Expreffions if they are equal, then the Center falls in the Axis: If  $\frac{1}{5}$  by be greater than  $\frac{1}{27}$   $b^3$ + 1 q, then it falls to the Left Hand of the Axis, if lefs, then to the Right. If therefore the fquare Root of the Cube of  $\frac{1}{2}$  UH (that is of  $\frac{1}{2}b^2 - \frac{1}{3}p$ , or putting  $\frac{1}{2}b^2 - \frac{1}{3}p = d$ , if  $\sqrt{ddd}$  be greater than HR, that is the difference between  $\frac{1}{27}b^3 + \frac{1}{2}q$  and  $\frac{1}{2}bp$ ; the Center R will be found within the Space NPU circumscrib'd by the Paraboloids UPX, UNL, and the interminate right Line DNP; and fo the Circle will cut the Parabola in three points Y, Y, Y, polited to the Right Hand

Hand of the Line BK, and fo the Equation will have three Roots. But the Center being without this Space NUP, it is explicable but by one Affirmative Root. Here it may be noted by the by, that the Right Line DP may touch the Paraboloid UPX in the point P, EP being  $\frac{1}{2}$ ,  $b^3$ ; but will cut the other Paraboloid UNL in the point N, fo that letting fall NF perpendicular to the Axis, UF is  $\frac{1}{4}$  EU, or  $\frac{1}{24}b^2$ , and NF  $\frac{1}{108}b^3$ . But UW (which being perpendicularly applied to the Axis at the point  $\mathcal{V}$ , meets DP in W) is  $= \frac{1}{4}b^3$ , or  $\frac{1}{4}$  EP.

Hence we may fafely conclude, that if in the Equation either p be greater than  $\frac{1}{3}b^2$ , or q greater than  $\frac{1}{27}b^3$ , that there will be found but one Root, and that an Affirmative one. Carte's Rule therefore (Page 70. Edit. Amfterd. 1659.) is not true, in which he determines that there are always as many true Roots, as there are changes of the Sines + and - in the Equation: Schooten in his Commentaries vainly endeavouring the defence of this Miltake. Alfo Prop. 5. Sect. 5. of our Country-man Harriot's Ars Analytice (as also Prob. 18. of Vieta's Numer. Poteft. Retol.) is hardly found; fince from the Limitations which they have there fet down, that must agree to the whole Parallelogram PIUW, which we have prov'd does agree only to the Space NUP; that is to afford a Center to the Circle interfecting the Parabola in three other points befides the Vertex.

But the quantity g or the laft Term (band p being given, fo that p be lefs than  $\frac{1}{2}b^2$ )

. . . .

#### $\frac{1}{3}b^2$ ) is exactly limitted from the foregoing

Equation  $\sqrt{ddd} = \frac{1}{27}b^3 + \frac{1}{2}q$  or  $\frac{1}{6}bp$ ; viz. when the Circle touches the Parabola. Therefore  $\frac{1}{2}q$  ought to be lefs than  $\frac{1}{6}bp - \frac{1}{27}b^3 + \sqrt{d^3}$ ; but if p be greater than  $\frac{1}{4}b^2$ , alfo  $\frac{1}{2}q$ ought to be bigger than  $\frac{1}{6}bp - \frac{1}{27}b^3 - \sqrt{d^3}$ , that the Center may not fall in the little Space NUW. And with these Conditions the Equation will always be explicable by three Roots; otherwise but by one. But whether there be three or one, they are always Affirmative ones, because of the position of the Center R to the Right Hand of the Line DP.

And this is the most difficult Cafe; fo that those that well understand what has gone before, will without any trouble take what comes after. Now let the Equation  $z^3 - b$  $z^2 + pz + q = 0$ , be given. Here (that there may be three Roots had) the Center of the Circle ought to be found fomewhere within the Space PNA, determin'd by the right Lines PN, PA, and the Curve of the Parabola  $N_{\Delta 3}$  wherefore fince EF is  $=\frac{1}{8}bb$ , p ought to be lefs than  $\frac{1}{4}bb$ . Now for the determination of the quantity q, d being =  $\frac{1}{5}$  $b^2 - \frac{1}{3}p$  as before,  $\sqrt{d^3 + \frac{1}{2^7}b^3} - \frac{1}{5}bp$  ought always to be greater than  $\frac{1}{2}q$ , that fo the Center of the Circle may be polited in the forementioned Space PNA; which when 'tis fo, fuch an Equation has two Affirmative Roots, and one Negative. But if p be greater than  $\frac{1}{3}bb$ , or  $\frac{1}{2}q$  greater than  $\sqrt{d^3} - \frac{1}{27}b^3 - \frac{1}{27}b^3$ top; it is explicable but by one (and that a Negative) Root.

Let

Let the Equation  $z^3 = bz^2 - pz - q = o$ , be proposed in the next place. That this Equation may have three Roots, the Center of the Circle must be found somewhere in the indefinite Space between the right Line D P D and the Curve of the Paraboloid PX. The quantity p is not here liable to Limitations; but  $\frac{1}{2}$  q ought always to be lefs than  $\sqrt{d^3} - \frac{1}{27}b^3 - \frac{1}{6}bp$ , supposing d to be  $= \frac{1}{9}b^2 - \frac{1}{3}p$ . By this means, there are two Negative Roots afforded, and one Affirmative; but otherwise, if  $-\frac{1}{2}q$  be greater than  $\sqrt{d^3} - \frac{1}{27}$ ,  $b^3 - \frac{1}{6}bp$ , the Equation is explicable by one only (Affirmative) Root.

Fourthly, Let the Equation  $z^3 - bz - pz$  + q = o, be proposed, which has two Affirinative Roots, and one Negative, if the Center of the Circle be found in the indefinite Space between the right Lines P $\Delta$ , PD, and the Curve of the Paraboloid  $\Delta L$ ; that is, (putting  $d = \frac{1}{2}bb - \frac{1}{3}p$ ) if  $\frac{1}{2}q$  be lefs than  $\sqrt{d^3} - \frac{1}{27}b^3 - \frac{1}{6}bp$ ; but if  $\frac{1}{2}q$  be greater than this quantity, there is but one (Negative) Root.

But the four remaining Equations in which we have -b, do not differ from those that have been mention'd already, as to the Limitation of the Number of the Roots, if the Sign of the last Term be changed, keeping the Sign of the third Term. But then them that were the Affirmative Roots in the former, will be the Negative ones here, and contrariwife.

Thus in the Equation  $z^3 - bz^2 + pz - q$  = o, the Affirmative Roots were either one or three; but in this Equation  $z^3 + bz^2 + pz$ + q = o

IIO

ITI

 $+q \equiv o$ , there is either one or three Negative Roots, under the very fame Conditions; but no Affirmative Root at all. So alfo in the Equation  $z^3 - bz^2 - pz - q \equiv o$ , there are two Negatives and one Affirmative, if p be lefs than  $\frac{1}{3}bb$ , and  $\frac{1}{2}q$  lefs than  $\sqrt{d^3} + \frac{1}{2}$ .  $b^3 - \frac{1}{6}bp$ ; even as in the Equation  $z^3 - bz^2$  $+pz + q \equiv o$ , there were two Affirmatives and one Negative : But the quantities p and q exceeding those preferib'd Measures, there is bere only one Affirmative Root, which there was a Negative one. In like manner, in the Equation  $z^3 + bz^2 - pz + q \equiv o$ , there are either two Affirmatives and one Negative, or one Negative only

Laftly, For the fame reafons in the Equation  $z^3 + bz^2 - pz - q \equiv o$ , there are two Negatives and one Affirmative, or one Affirmative only, for which, in the Equation  $z^3 - bz^2 - pz + q \equiv o$ , there were two Affirmatives and one Negative, or one Negative alone; v/z. as  $\frac{1}{2}q$  is either greater or lefs than  $\sqrt{d^2} - \frac{1}{2} + \frac{1}{2}b^3 - \frac{1}{6}bp$ .

If the third Term (or pz) be wanting, the Center R always falls in the Line IPE_{$\Delta$}, wherefore if it be  $z^3 - bz^2 \cdot x - q$  or  $z^3 + bz^2 \cdot x - q$  or  $z^3 + bz^2 \cdot x - q$  or  $z^3 + bz^2 \cdot x - q$ , there can be but one Root, Affirmative if it be -b, Negative, if +b. But if it be  $z^3 - bz \cdot x - q$  or  $z^3 + bz^2 \cdot x - q$ , there may be two Aifirmatives and one Negative in the former, or one Affirmative and two Negatives in the latter, the Center falling in the Line  $P_{\Delta}$  between P and  $\Delta$ , that is if  $\frac{1}{4}q$  be lefs than  $2\frac{1}{7}b^3$ ; for if it be greater, there can be but one Negative in the former, or one Affirmative in the latter, Hitherto

Hitherto we have obtain'd the Number of the Roots in Cubick Equations, it remains that we add fomewhat concerning the quantity of the Roots. And here it is first of all to be noted, that every Equation having three Roots, may be expeditionally enough refolv'd by the help of the Table of Sines, that is by the Trifection of an Angle, by putting  $\sqrt{\frac{4}{5}b^2 - \frac{4}{3}p}$  or  $\sqrt{\frac{4}{4}d}$  = the Radius of the Circle, if it be -1- p in the Equation; or  $\sqrt{\frac{4}{5}b^2 + \frac{4}{3}p}$ , if - p; and the Angle to be Trifected, that which has its Sine (in the Table of Sines)  $\frac{1}{27}b^3 + \frac{1}{6}bp + \frac{1}{2}q$ . This

Angle being found, the Sine of its third part, as also the Sine of the third part of its Complement to a Semi-circle, and their Sum, will be given from the Table of Sines. Now these Sines are to be multiplied into

the Radius  $\sqrt{\frac{1}{3}b^2 + \frac{3}{3}p}$ , and thus will be obtain'd the quantities (y& y& y& in the Fig.) the Sum or Difference of which and  $\frac{1}{3}b$ , as the cafe requires, will give the true Roots of the Equation. All these things are deduced from *Cartes*'s Difcoveries. But that I may comprehend all the Cafes, with as much Brevity as is possible; I fay, that the Center  $R_5$ in the first Formula of Equations, falling in the Space UGP, the two Interfections T, T, fall between A and B, and confequently either of the leffer Roots is lefs than  $\frac{1}{3}b_5$ but the third and greater always exceeds  $\frac{1}{3}b_5$ but is exceeded by b. But if the Center falls

II2

falls in the fpace G N U, there are two greater than  $\frac{1}{3}b$ , but lefs than  $\frac{1}{3}b$ , but the third is b — the two others, and confequently lefs than  $\frac{1}{3}b$ ; but using the Limitation of the Quantity p, the Roots are included in narrower Bounds. For the greatest Root is lefs than  $\sqrt{\frac{1}{3}b^2 - \frac{1}{4}p} - \frac{1}{3}b$ , but greater than  $\sqrt{\frac{1}{4}b^2 - p} - \frac{1}{4}p + \frac{1}{3}b$ , but greater than  $\sqrt{\frac{1}{4}b^2 - p} - \frac{1}{4}p + \frac{1}{3}b$ ; but when  $\frac{1}{4}bb$  is lefs than p, that Limit becomes  $\sqrt{\frac{1}{3}b^2 - \frac{1}{3}p} - \frac{1}{3}b$ . The mean Root is always lefs than  $\sqrt{\frac{1}{4}b^2 - p}$  $+\frac{1}{3}b$ , but greater than  $\frac{1}{3}b - \sqrt{\frac{1}{9}b^2 - \frac{1}{3}p}$ ; but the least Root never exceeds this Limit, but vanishes with the Quantity q.

In the fecond Formula, according to the prefcrib'd Laws, there are two Affirmative and one Negative Root; and the Center falling in the Space G P E, one of the Affirmatives is greater, and the other lefs than  $\frac{1}{3}b$ , but the greater exceeds not b; but the Negative cannot be greater than  $\sqrt{\frac{1}{3}bb-\frac{1}{3}b}$ , and it is the difference of b and the Sum of

and it is the difference of b and the Sum of the Affirmative Roots. But the Center being polited in the Space  $E N G \Delta$ , either of the Affirmatives is greater than  $\frac{1}{3}b$ , but lefs than  $\sqrt{\frac{1}{3}bb+\frac{1}{3}b}$ ; but the Negative is ever lefs than  $\frac{1}{3}b$ . But the nearer Limits (from

the Quantity p given) are  $\sqrt{\frac{1}{4}bb-p+\frac{1}{2}b}$ , of the greatest Affirmative Root; than which it is always less, as also greater I than than  $\sqrt{\frac{1}{9}} \frac{bb}{bb} - \frac{1}{3}p - \left|-\frac{1}{3}b\right|$ ; yet the other Affirmative Root (which is diminish'd with the Quantity q) is lefs than this *Limit*. But the Negative Root is always lefs than  $\sqrt{\frac{4}{9}bb} + \frac{4}{3}p - \frac{1}{3}b}$ , and the Quantity q being wanting, vanishes.

In the third Formula; there are two Negatives and one Affirmative. In this, as in the fourth, the Roots are not limitted by the Quantity b. But the Affirmative Root is ever lefs than  $\sqrt{\frac{4}{5}bb+\frac{4}{3}p+\frac{1}{3}b}$ , yet greater than  $\sqrt{p+\frac{1}{4}bb+\frac{1}{2}b}$ ; and the greateft of the Negatives is always greater than  $\sqrt{\frac{1}{5}bb+\frac{1}{3}p-\frac{1}{3}b}$ , but lefs than  $\sqrt{p+\frac{1}{4}bb}$  $-\frac{1}{2}b$ . But the lefs of the Negatives is always leffen'd with the leffen'd Qantity q.

In the fourth Formula, the Center falling within the Space  $L \triangle PD$ ; if there be two Affirmative and one Negative Root, the greatest of the Affirmative Roots cannot be greater than  $\sqrt{p+\frac{1}{4}bb+\frac{1}{2}b}$ , nor less than  $\sqrt{\frac{1}{9}bb+\frac{4}{3}p+\frac{1}{3}b}$ . But the less Negative is less than  $\sqrt{\frac{4}{5}bb+\frac{4}{3}p-\frac{1}{3}b}$ , and greater than  $\sqrt{p+\frac{1}{4}bb-\frac{1}{2}b}$ . But 'tis to be noted here, that the Negative Roots are every where mark'd with the Affirmative Sine, because these are the Affirmative Roots of those

four Equations, in which is found -b, and q is

is affected with the contrary Sine; as I intimated above.

The Demonstration of all these things follows from hence, that where-ever the Cen-ter of the Circle R falls upon the Curve Lines UPX or UAL, the Circumference of it touches the Parabola in a Point whofe diftance from the Axis is  $\sqrt{\frac{2}{3}}$  VH, and cuts it on the other fide the Axis at the diffance of  $2\sqrt{\frac{2}{3}}$  U H; but when the Center falls on the Line D P D, one of the Roots is = o, and confequently the Cubick Equation is reduced to a Quadratick one, or to  $z^2 - bz$ -1-p = o, the Roots of which give the Limits when the Quantity q vanishes; and by how much the lefs q becomes, by fo much the nearer do the Roots approach to thefe Limits. The Equation is alfo Quadratical, when the Center falls in the Axis; that is, when  $\frac{1}{2}q = \frac{1}{6}bp - \frac{1}{27}b3$ , in the first Formula; or  $\frac{1}{2}q = \frac{1}{27}b3 - \frac{1}{6}bp$ , in the fecond; in the third 'tis impossible; but in the fourth, when  $\frac{1}{2}q = \frac{1}{27}b_3 - \left|-\frac{1}{6}b_p\right|$ ; in which cafe the lefs of the Affirmative Roots is  $\frac{1}{3}b_1$ , and the greater  $\sqrt{\frac{1}{3}bb+p+\frac{1}{3}b}$ , but the Negative  $\sqrt{\frac{1}{3}bb}$  $-p - \frac{1}{2}b$ . In the first Formula, the Roots are  $\frac{1}{3}b$ , and  $\frac{1}{3}b + \sqrt{\frac{1}{3}bb - p}$ . But in the fecond, the Affirmatives are  $\frac{1}{5}b$ , and  $\sqrt{\frac{1}{5}}bb$  $-p + \frac{1}{3}b$ , but the Negative  $\sqrt{\frac{1}{3}bb} - p - \frac{1}{3}b$ . And thefe things may feem to fuffice in Cubicks ; but because of the excellent use of the Method, by which, by the help of the Table o f I 2

of Sines, the Roots of these Equations are found; I thought convenient to add an Example or two, by which the Compendium of that Practife may be rendred manifest. Let the Equation  $z_3 - 39z^2 - 479z - 1881 = 0$ , be proposid; and the Roots z are fought. Here  $\sqrt{\frac{1}{2}bb} - \frac{1}{3}p = \sqrt{9\frac{1}{3}} = \sqrt{d}$ , whose double  $\sqrt{37\frac{1}{3}}$  is the Radius of the Circle; alfo  $\frac{1}{2}, b_3 - \frac{1}{2}q - \frac{1}{6}bp = 2197 + 940\frac{1}{2} - 3113\frac{1}{2}$ ,

$$d_3 \qquad 9\frac{1}{3}\sqrt{9}$$

or  $\frac{1}{9\frac{1}{2}}$  is the Tabular Sine of the Angle;

that is, making a Division by the help of the Logarithms, Log. 9.9251560, to which corresponds an Angle of  $57^{\circ}$ . 19'. 11'''. The third part of this is 19°. 6'. 24''. and of the Complement, is 40°. 53'. 36''. The Sines give the Logs, 9.514983 and 9.816011;

which multiplied into the Radius  $\sqrt{37\frac{1}{3}}$ , produce Y& and Y&, Log. 0.301030 = 2, and Log. 0.601059 = 4, but the third Y& is equal to the Sum of them, or 6. And therefore the Roots are 13-4=9; 13-2= 11, 13+6=19; of which feveral ones the foremention'd Equation is compos'd. Where 'tis to be noted that the two leffer Roots, do not exceed  $\frac{1}{3}b$  or 13, becaufe the Center R in the Conftruction falls on the right hand of the Axis ; that is,  $\frac{1}{6}bp$  is lefs than  $\frac{1}{27}b3+\frac{1}{2}q$ .

For

Miscellanea Curiofa. 117 For another Example, let us enquire out the Roots of the Equation  $x_3 - 15x^2 - 229x$ -525 = o. Here  $\sqrt{\frac{1}{2}bb + \frac{1}{3}p} = \sqrt{101\frac{1}{3}} =$  $\sqrt{d}$ , and the Radius of the Circle  $= \sqrt{405\frac{1}{3}}$ . Alfo  $\frac{1}{2}7b^3 + \frac{1}{6}bp + \frac{1}{2}q = 125 + 572\frac{1}{2} + 262\frac{1}{2}$ 

the Tabular Sine of an Arch, whole Log-9.9736426, and the Arch it felf  $70^{\circ}$ . 14'. 22". The third part of it, is  $23^{\circ}$ . 24'. 47^{'1}/₂, and of the Complement, is  $36^{\circ}$ . 35'.  $12''\frac{1}{2}$ , whole Log. Sines are 9.599183, and 9.775275, to

101 1 101 1

Vada

which adding the Log.  $\sqrt{405\frac{1}{2}}$ , we have the Log. 0.903089=8, and I.og. 1.079181=12, the Sum of which is equal to 20. Hence we conclude that  $20\sqrt{-\frac{1}{3}}b$  or 25, is equal to the Affirmative Root, and 8 and  $12-\frac{1}{3}b$ , that is 8 and 7 equal to the Negative Roots. But if the Equation had been  $x3\sqrt{-1}$   $15x^2-$ 229x-525=0, then 8 & 7 had been the Affirmative Roots, and 25 the Negative. As for the other Cubicks which are explicable by one only Root, they are to be refolv'd by *Cardan*'s Rules, after the fecond Term is taken away; neither do I fee how the bufinefs can be done with lefs Calculation.

But if this Root be defir'd to be expressed in the Terms of the Quantities b, p, q, Ifay that in the first Formula it is,  $\frac{1}{3}b$  + or the Sum or Difference of the Cubick Roots

of  $\sqrt{\frac{1}{4}qq - \frac{1}{108}p^2 b^2 + \frac{1}{27}b^3 q - \frac{1}{6}bpq + \frac{1}{27}p^3}$  $\pm \frac{1}{27}b^3 + \frac{1}{2}q - \frac{1}{6}bp$  (viz. -- if  $\frac{1}{27}b_3 + \frac{1}{2}q$  be I 3 greater greater than  $\frac{1}{6}bp$ , otherwife —) the Sum, when  $\frac{1}{3}bb$  is greater than p, the difference when lefs. And in the other *Formula*, the Root is always compos'd of the fame parts, only the Sines -1 and — being varied, as they will eafily perceive that are willing to make the Tryals.

But these Roots are readily enough found by the help of the Log. Table of versed Sines; viz. if the Coefficients are fund or broken Numbers, and the Roots not to be expressed in Numbers, as most commonly it happens.

Now this is the Rule. In the first and fecond Formula, if  $\frac{1}{3}$  bb be lefs than p, let  $\frac{1}{3}$  p  $-\frac{1}{9}$  bb = d, and putting the difference between  $\frac{1}{6}$  bp, and  $\frac{1}{27}$  b3  $\frac{1}{7}$   $\frac{1}{2}$  q (that is H R) in the first Formula, and the difference between  $\frac{1}{6}$  bp  $-\frac{1}{2}$  q and  $\frac{1}{27}$  b3 (in the fecond Formula) Radius, let the Angle, whole Tangent is

 $d\sqrt{d}$ , be found. Then, as the Co-fine of this Angle, to the verfed Sine of the fame, fo the Difference made Radius, to a fourth Quantity, the Cube Root of which will be had by taking the  $\frac{1}{3}$  of its Log. Then dividing  $\frac{1}{3}p - \frac{1}{2}bb$  by this Cube Root, let the Divifor be fubfracted from the Quotient, the Remainder will be the Quantity Y& at Fig. 1. The Sum of this Remainder and  $\frac{1}{3}b$  will be the Root fought, if the Center falls on the Right Hand of the Axis; otherwife their Difference will be the Root. But if  $\frac{1}{3}bb$  be greater

than p, making HR Radius, let  $d \vee d$  (or the diffance of the Paraboloid from the Axis) be the Sine of fome Arch; let the verfed

Dianced by Google

119

verfed Sine of this be multiplied into Radius or  $\frac{1}{6}bp - \frac{1}{2}$ ,  $b3 + \frac{1}{2}q$ , and taking  $\frac{1}{3}$  of the Log. of the Product, its Cubick Root will be obtained, by which let  $\frac{1}{2}bb - \frac{1}{3}p$  be divided. I fay, that the Sum of the Quotient and Divifor, after the fame manner added to or taken from 1/26, will give the Root fought. And the like for the third and fourth Formula, unless that  $\frac{1}{27}b_3 + \frac{1}{6}b_7 + \frac{1}{2}g$  is to be

taken for Radius, and 1 bb 1- 1 p into V to bb

 $\frac{1}{1}$  p, or  $d\sqrt{d}$ , for the Sine. But these Rules will be perhaps better understood by Examples.

Suppose the Cubick Equation  $z_3 - 17z_2$ -54z - 350 = 0, and let the Root z be fought. Here 3 bb is is greater than p, but q is bigger than the Cube of  $\frac{1}{3}b$ , and therefore 'tis explicable by one Affirmative Root,

only. Now  $\frac{288}{9} - \frac{1}{3} + is d$ , and  $\frac{1}{9} \sqrt{\frac{1}{9}} = is$ to be taken for the Sine, to the Radius  $\frac{4913}{27}$ 175 - 153, that is  $2\frac{53}{27}^2$ ; and the Arch agreeing thereto is  $15^\circ$ .  $30^\circ$ .  $49^\circ$ . The Log. verfed Sine of this 8.5362376, added to the Log. of the Radius 2.3095913, makes 0.8457889, the 3d part of which 0.2819276, is the Log. of the Cube Root 1.91394, by which, as a Divisor, dividing  $\frac{1}{9}^2$  or d, the Quotient is 7.37281. The Sum of the Quotient and Divifor encreafed by the addition of 1 b, is the Root fought, viz. 14.9534, &c.

Having thus difpatch'd Gubick Equations, let us proceed to Biquadratical ones. These have always either none, or 2, or 4 true Roots, 14

Roots, the determination of which depends partly on the *Coefficients*, partly on the *Sine* and *Magnitude* of the abfolute Number given. A general Conftruction for all thefe (and that eafy I conceive enough) I have delivered at N° 188, which I fuppofe the Readear to be acquainted with; but yet the Figure relating to that Matter, I think proper to bring hither, (Fig. 2.) In the Conftruction of the Equation  $z4 - bz_3 + pz_2 - qz$ +r = a, let  $BD = \frac{1}{4}b$ ,  $AB = \frac{1}{16}bb$ ,  $BK = \frac{1}{2}$ or  $\frac{1}{2}$  the Parameter,  $KC = 2AB = \frac{1}{8}bb$ ,  $KE = \frac{1}{8}bb - \frac{1}{2}p$ ,  $AE - \frac{1}{2} = \frac{1}{16}bb - \frac{1}{2}p$ ,  $FE = \frac{1}{16}b^3 - \frac{1}{4}bp$ , and  $EG = \frac{1}{16}b^3 - \frac{1}{4}bp + \frac{1}{2}q$ . Which done, a Circle on the Center G with

the Radius  $\sqrt{GD^2 - r}$ , will interfect the Parabola, either in none, or 2 or 4 Points, from whence Perpendiculars let fall on DH, will give all the Roots z. But that there may be 4, 'tis evident that the Center of the Circle ought to be found fomewhere within a space from any Point of which, three Perpendiculars may be let fall upon the Curve of the Parabola; and alfo that the Radius is lefs than the greatest of those Perpendiculars, and greater than the middle one. But that if the Center, be posited without this space, fo that there can be but one Perpendicular let fall upon the Parabola, and the Radius greater than it, or if it be lefs than the middle one of the 3 Perpendiculars, but greater than the least of them; then there can be but two Roots only. But there to Biquisicent ones. Thefe is no Root at all, when the Radius  $VGD_2^{-r}$ . 21 Get 1 2 I

120

is lefs than the least of the 3, or than the one as often as there is but one. Now it remains for us to inquire of what kind this Space is, by what Limits 'tis diffinguished, and under what Conditions the Radius of the Circle is lefs or greater than the foremention'd Perpendiculars. And first of all. we must shew how a Perpendicular is to be let fall upon the Parabola. Let (Fig. 3.) ABC be a Parabola, AE its Axis, AV the Parameter, G the point from whence the Perpendicular is to be let fall. Let GE be drawn perpendicular to the Axis, and VE be bifected in F, and erecting the Perpendicular FH on the fame fide of the Axis. let  $FH = \frac{1}{4}GE$ ; I fay that a Circle defcrib'd on the Center H, with the Radius HA, will interfect the Parabola in three points, or one, z, the right Lines GZ drawn to which, will be perpendicular to the Curve of the Parabola. But now that there may be 3 fuch Intersections, the Center H ought to be fo polited, as that it may be within the fpace included by the Paraboloids (in Fig. 1.)

that is, that FH may be lefs than  $\sqrt{\frac{5}{27}}$  FV³, or FH² lefs than the Cube of  $\frac{3}{3}$  VF; and

fo G E = 4 F H will be lefs than  $4\sqrt[4]{27} V F^3$ , that is, the fquare of GE will be lefs than  $\frac{1}{27} V E^3$ . Therefore these Limits coincide with two Paraboloids of the fame kind with those which were used in Cubical Equations, but whose Parameter is twice lefs, viz.  $\frac{1}{27}$  of the Parameter of the Parabola, that is  $\frac{1}{27}$  of A V. And therefore it is that very out of A V.

122

Curve Line, by the Evolution of which the Parabola is defcrib'd (as *Hugenius* has demonftrated) and which, the Line DF (*Fig. 2.*) which is perpendicular to the Parabola in the point D, is always a Tangent to. But the point P (that is, that in which the right Line DF touches the Paraboloid) is the Center of a Circle, which (being defcrib'd with the Radius DP) coincides with the Parabola in the point D, or has the fame *Curvature* with it, as is manifeft.

Having therefore defcib'd fuch Paraboloids UXP, VN $\triangle$  (Fig. 2.) on either fide the Axis, 'tis clear, that unlefs the Center of the Circle be placed within thefe Limits, it cannot interfect the Parabola in more than two points. From whence we may determine, under what conditions, the Coefficients of the intermediate Terms are reftrained, in Biquadratick Equations, that fo there may be four Roots. And at first fight 'tis plain that p cannot be greater than  $\frac{1}{8}$  bb, (viz. in those Forms where 'tis  $\frac{1}{7}$  p) nor q than  $\frac{1}{16}$  b³. But in General,  $\frac{1}{16}$  b³  $\frac{1}{7}$   $\frac{1}{4}$  pb  $\frac{1}{7}$  $\frac{1}{2}$  q, that is EG the diffance of the Center from the Axis, ought to be lefs than E H =

 $4\sqrt{\frac{1}{27}}$  VE³, that is (because VE= $\frac{1}{16}$  bb  $\overline{+}$ 

 $\frac{1}{2}p$ ) than  $\frac{1}{4}bb + \frac{1}{3}p \sqrt{\frac{1}{16}b^2} + \text{ or } -\frac{1}{6}p$ , the Sines + and - being left doubtful, that fo they may be varied according to the nature of any Equation; as was fhewn above in Cubicks. Neither would I be offenfively tedious to the Learned on the one hand, nor deprive Learners on the other, of the Exercife

ercife and Pleafure, of fending out thefe things by themfelves. As for the Limitation of the least Term r, it cannot be found with the fame eafinefs, and that becaufe, to let fall a Perpendicular upon the Curve of a Parabola, is a folid Probleme, and which cannot be refolv'd without the folution of a Cubick Equation. Therefore first of all let the fecond Term be wanting, or if there, let it be taken away, fo that the Equation may have this Form z4. *. pz2. qz. r. = o. And if it be -r, it is always explicable by two or four Roots; but that there may be four, the Center of the Circle ought to be polited within the foremention'd Paraboloids, or that it may be -p, and qq may be lefs than  $\frac{3}{27} p_3$  or the Cube of  $\frac{3}{3} p$ . Then let the Roots of this Equation y₃. *.  $\frac{1}{2} py$ .  $\frac{1}{4} q = 0$ , be gotten, the Quantities p and q having the fame Sines as in the Biquadratick. And these Roots are found expeditionly enough by the help of the Table of Sines. But having found those three y (which are ordinately applied to the Axis of the Parabola from the points, where the Perpendiculars to the Curve of it do fall, viz. YZ in Fig. 3.) than pyy - 3y4 of the leffer y will denote the greatest Quantity of r, if it be-r, than which if r be lefs, the Equation will have four Roots, otherwise but two. But if it be + r, it ought to be lefs than 3y4 - pyyof the middle y, for if it be greater, it can have but two Roots; at leaft, if r be lefs than 3y4 - pyy of the greatest y. But if it be greater than this, the Equation is not explicable by any true Root at all. Thefe

Digitized by Goo

122

These fame Limits, are otherwise expressed by the Quantity q, viz.  $\frac{1}{2}qy - y4$  in the first cafe,  $y4 - \frac{1}{2}qy$  in the fecond, and  $y4 + \frac{1}{2}qy$ in the third. But it may be, that the two leffer Quanties y may not be far different from one another, whence it comes to pass that both of the Perpendiculars are greater than the right Line GA, viz. when qq is greater than  $\frac{4}{27}p^3$ , but lefs than  $\frac{8}{27}p^3$ ; the Center falling within the fpace contain'd between the Paraboloids of Fig. 1. and 2. In this cafe, if it be 1 r, there can be but two Roots, y4 1 = gy of the greateft y being greater than r; otherwife none. But if  $\frac{1}{2}qy - y4$ of the least y be greater than r mark'd with the Sine -, but r be greater than  $\frac{1}{2}qy - y4$ of the mean y, then there will be four Roots; but two only, if r be found greater than the former, or lefs than the latter. But if in the Equation it be -p, or if it be -p and qq be greater than  $\frac{3}{27}p^3$ , the Equation  $y^3$ . *.  $\frac{1}{2}py$ .  $\frac{1}{4}q$ . is explicable by only one Root y; that is, there can be but one Perpendicular only let fall from the Center of the Circle. Whence it may be certainly concluded that there can be but two Roots only in the given Equation, the Sum of which, if it be -r, is increas'd with the Quantity r; but if it be +r, the Quantity y being obtain'd, that Quantity r ought to be lefs than y4-1-12 ay, for if it be greater, the Equation propos'd is abfurd and impoffible.

"Twould be both tedious and needlefs to run over all Equations of this kind, fince 'tis evident (from what has been already faid) to those that are attentive, which are Ne-

125

gative and which Affirmative, and that the Limits of these Roots are deriv'd from the found Quantities y. But for an Example (which any one may imitate in the reft of of the Cafes) let it be propos'd to difcover the Limits or Conditions, under which, there may be four Affirmative Roots in a Biquadratical Equation. Now this will be as often as the Center of the Circle G is posited in the fpace UPK (Fig. 2.) and also  $\neg r$  or the Radius of the Circle is lefs than GD. Whence 'tis plain, that the Equation here concern'd is of this Form, z4-bz3+pz2qz + r = o; and that p cannot be greater than  $\frac{1}{8}bb$ , nor  $\frac{1}{4}pb$  (in this cafe) than  $\frac{1}{16}b3$ 1- 1 q; again, 'tis neceffary that 1 bb - 1 p

in  $\sqrt{\frac{1}{6}bb - \frac{1}{6}p}$  fhould be greater than  $\frac{1}{6}b_3$  $-1-\frac{1}{2}q-\frac{1}{4}bp$ ; and from these Limits, it will be manifest that the Center is contain'd within the fpace UPK. But in order to the determination of the Quantity r, this Cubick Equation must be first folv'd, y3. *.  $-\frac{1}{16}\overline{b^2 - \frac{1}{2}py} = \frac{1}{32}b_3 - \frac{1}{4}q - \frac{1}{8}pb$ ; and fo will be obtain'd the Points upon which fall the Perpendiculars from the Center to the Curve of the Parabola. Now having found the three Values of this y; the Quantity r ought to be lefs than  $\frac{1}{256}b4 - \frac{1}{4}bq - \frac{1}{16}bbp$  $+3y4 - \frac{1}{8}b^2y^2 + pyy$  of the middle y, but greater than  $\frac{1}{256}b^4 + \frac{1}{4}bq - \frac{1}{16}bbp + 3y4 - \frac{1}{256}b^2 + \frac{1}{2}bq - \frac{1}{16}bbp + \frac{1}{2}bq - \frac{1}{2}b^2 + \frac{1}{2}bq - \frac{1}{2}b^2 + \frac{1}{2}b^2 + \frac{1}{2}bq - \frac{1}{2}b^2 + \frac{1}{2}b^2$ b2 y2 -1- pyy of the least y. But if r exceed these Limits, there can be but two Roots obtain'd. Laftly, if 236 b4 -1- 4 bp - 16 bbp +  $3y_4 - \frac{1}{8}bby + pyy$  of the greatest y, be greater than r, then the Equation proposed is im-

impossible. It happens also that there are four Affirmative Roots, when the Center G is posited in the little space UTS, viz. drawing RTS perpendicular upon the middle of the supposed Line AD. But this comes to pass when p is greater than  $\frac{1}{16}bb$ ,

and  $\frac{1}{4}bb - \frac{2}{3}p \sqrt{\frac{1}{16}bb - \frac{1}{6}p}$  greater than  $\frac{1}{8}$  $pb - \frac{1}{128}b3 - \frac{1}{2}q$ . In which cafe always two, fometimes three of the Roots are greater than  $\frac{1}{4}b$ .

But 'tis to be noted here that that Limit produced from the least y, is fometimes Ne-gative, or less than nothing; viz. as often as the greatest of the three Perpendiculars is greater than GD (Fig. 2.) If this happens, the Quantity + r may be diminish'd to nothing from the Limit prefcrib'd, by the middle y. But the defect of a Limit from the leaft y, fnews how great -r may be in the Equation, if there be three Affirmative Roots and one Negative one; which if it exceeds, there can be but two, one Affirmatsve and the other Negative. And all these things are demonstrated from hence, that the foremention'd Limits of the Quantity r, are the differences of the Squares of the Line GD, and the Perpendiculars to the Curve of the Parabola.

But becaufe of the perplexing Cautions arifing from the diverfity of Sines with thefe Equations, 'tis better always to take away the fecond Term, and then to inquire out the number of Roots and the Sines, according to the Rules already deliver'd; efpecially if thofe Quantities y are not much different from one another. But of thefe four Affirmative Roots, two

two are always lefs than  $\frac{1}{4}b$ , and two greater, viz. if DG be lefs than AG, or  $\frac{1}{4}pb$  than.  $\frac{1}{64}b3 + 9$ . But three are always lefs than  $\frac{1}{4}b$ , as often as the mean Perpendicular (or that found from the mean y) is greater than AG, or  $\frac{3}{8}bby$  greater than  $\frac{3y3}{2} - pyy$  of the fame mean y. The fourth and greateft Root is greater than the greateft  $y + \frac{1}{4}b$ ; and 'tis equal to the difference of b and the Sum of the other three Roots, and therefore is lefs than b.

But 'tis now time to have done with this Matter. Perhaps those that more perfectly understand the Nature of the Parabola, may be able to do all these things after a more compendious manner. But there is some cause to doubt, whether all these Quantities b. p, q. r. can be rightly determin'd without the Resolution of a Cubick Equation, or no. For whatsoever is done in Plain Equations in this Matter, exhibits, not the trae Limits, but fome Approximations only.

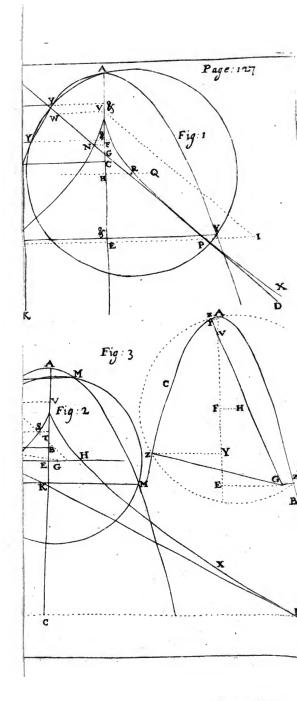
Some Illustrious Specimens of the Doctrine of Fluxions; or Examples by which is clearly shewn the Use and Excellency of that Method in solving Geometrical Problems.

By Ab. De Moivre.

1

Y OU have here alfo the Method which I promis'd, concerning the Quadratures of Curvilineal Figures, the Dimension of the Solids generated by the Relation of a Plane (and of the Surfaces) the Rectification of Curve Lines, and the Calculation of the Center of Gravity. I know these Points have been already handled by several very learned Men—But I hope this Attempt of mine will nevertheless not altogether displease, if (especially) I have had the good Luck to find a shorter and more expeditious way to these things, than what is commonly known.

But hefore I proceed farther, I would have it observ'd, that I make use here, of what the celebrated Mr. Newton has de-monstrated, Page 251, 252, 253. Princ. Phil. concerning the Momentaneous Increments or Decrements, of Quantities that Increase or Decrease





Miscellanea Curiosa. 129 crease by a continual Flux; Especially, that the Momentane of any Power, as  $A^m$  is  $A^{\overline{m}}$ . Farther; the Fluxion  $\frac{n}{m}$ being given, the flowing Quantity 4^m may be found; First, By striking a out of the Fluxion; Secondly, By encreafing the Index of the Fluxion by Unity ; Thirdly, By dividing the Fluxion by the Index thus increased by Unity. In the following Discourse, we shall express Absciffe of any Curve by x, its Fluxion by x, the ordinate by y, and its Fluxion by y. These things fuppos'd; that we may come to the Quadra-ture of Curves, First, Take the value of the ordinate applicate, by the help of the Equation expressing the Nature of the Curve. Secondly, Let this Value be multiplied by the Fluxion of the Abscisse; for the Pro-duct arising is the Fluxion of the Area. Thirdly, Having the Fluxion of the Area, let the flowing Quantity be found, and fo we shall have the Area fought. Ex. gr. Let m the Equation x = y be propos'd, which expresses the Nature of all forts of Paraboloids.

The Value of y is  $x^n$ , which multiplied by

x

x, gives  $x^{n}$  x for the Fluxion of the Area, and confequently the Area fought is  $\frac{n}{m+n}$  $\frac{m}{x^{n}}$ , or (fubflituting y inftead of  $\mathcal{X}^{n}$ )  $\frac{n}{m+n}$ 

Again, suppose a Curve, whose Equation is  $x^4 + aaxx = yy$  (which is the first of the Excellent Mr. Craig's Examples) putting y=  $x \sqrt{xx + aa}$ , the Fluxion of the Area will be xxVxx+ aa. Which Expression involving a furd Quantity, let us suppose  $\sqrt{xx + aa} = z$ , then will  $xx + aa = z^2$ , and confequently xx = zz; and fubflituting zz and z for xxand  $\sqrt{xx + aa}$ , the Fluxion thus freed from Surds, will be  $z^2 z$ ; which reduced to its Original  $\frac{1}{3}z^3$  and putting  $\sqrt{xx+aa}$  for z, we have  $\frac{1}{2}xx + aa$  for the Area fought. But to fhew more effectually how eafily these Quadratures are perform'd, I shall add one Example more. Let the Equation of the 22  $-=y^2$ , therefore y=1Curve be -2:-1-a 2-1-4 and therefore is the Fluxion of the x-1-a Area.

This and by Google

Area. Put  $\sqrt{x+a} = z$ , hence x = zz - a,

and x = 2zz. Therefore  $\frac{xx}{\sqrt{x^{1-a}}} = 2z^2z - 2az$ ,  $\sqrt{x^{1-a}}$ 

and confequently  $\frac{3}{3}z^3 - 2az$ , or  $\frac{3}{3} - \frac{4}{3}a$  $\sqrt{x+a}$  will be the Area fought.

But it often happens that we meet with fome Curves (fuch as the Circle and Hyperbola) which are of fuch a Nature, that 'tis in vain to attempt the freeing the Fluxions of them from Surds. And then reducing the Value of the ordinate into an infinite Series, and multiplying the feveral Terms of the Series into the Fluxion of the Abfciffe (as before) let the Fluent of each of thofe Terms be found, and fo there will arife a new Series, which will exhibit the Quadrature of the Curve.

This Method is with the fame eafe applied to the Menfuration of the Solids generated by the Rotation of a Plane; viz. taking for their Fluxions, the Product of the Fluxion of the Abfeiffe into the circular Bafis. Let the Proportion of a Square to the inferib'd nCircle be —. The Equation expressing the Nature of a Circle is yy = dx - xx; therefore  $4 dxx - x^2x$  is the Fluxion of a Portion

of

of the Sphere, and confequently the Portion it felf is  $4\frac{1}{2}dx^2 - \frac{1}{3}x^3$ , and the circumfcrib'd

Cylinder is  $4 \frac{dx^2 - x^3}{n}$ . Therefore the Por-

tion of the Sphere is to the circumfcrib'd Cylinder, is as  $\frac{1}{2} d \rightarrow \frac{1}{3} x$  to  $d \rightarrow x$ .

The Rectification of Curve Lines will be obtain'd, if we confider the Fluxion of the Curve as a Hypothenufe of a Rectangular Triangle, whole fides are the Fluxions of the Ordinate and Abfciffe. But in the Expression of this Hypothenuse, care must be taken that only one of the Fluxions be remaining, as also only one of the indetermitiste Quantities, viz. that whose Fluxion is retain'd. Some Examples will render this clear.

(Fig. 1.) The right Sine CB being given, to find the Arch AC. Let AB = x. CB = y. OA = r. CE the Fluxion of the Abfcille, ED the Fluxion of the Ordinate, CD the Fluxion of the Arch CA. From the Property of the Circle 2rx - xx = yy, whence

 $2rx - 2xx \equiv 2yy$ , and therefore  $x \equiv yy$ . But

 $xx = yy + y^2yy$ - 2rx + xx

therefore CD =

Miscellanea Curiosa. 133  $= \frac{1}{\sqrt{rr - yy}} \times ry = ry \times rr - yy$ And  $\sqrt{rr - yy}$ confequently if rr - yy be thrown into an infinite Series, and the feveral Members of it

be multiplied into ry, and then the flowing Quantity of each be taken, we fhall have the length of the Arch AC. After the fame manner, giving the verfed Sine, the fame Arch may be found. For refuming the Equation found above 2rx - 2xx = 2yy, we have y = rx - xx, but  $CD^q = xx + yy = xx$ 

$$rrxx - 2rxxx + x^2xx$$

$$+ \frac{1}{yy} \qquad xx + \frac{1}{yy}$$

 $rrxx - 2rxxx + x^2xx$ 

vv

to the fame Denominator, and expunging

contradictory Terms)  $\frac{rrxx}{2rx - xx}$ , whence CD = rx, and confequently the  $\sqrt{2rx - xx}$ 

length of the Arch A C may be eafily found from what is faid already. K 3. The

Digitized by Google

The Fluxion of the Curve Line is fometimes more eafily found by comparing the two fimilar Triangles CED, CBO, for this Proportion arifes, CB: CO:: CD, that is for the Circle  $\sqrt{2rx - xx}: r::x: rx}$ 

The Curve of the Cycloid may be determin'd by the fame Method too. Let (Fig. 2.) ALK be a Semicycloid, whofe generating Circle is ADL. Having any point as B in the Diameter AL, draw BI parallel to the Bafe LK meeting the Peripheus of the Circle in the point D; compleat the Rectangle AEIB, and draw FH parallel to EI and infinitely near to it, as alfo BI cutting FH in G, and the Curve AK in H. Put AL = d.

AB (= EI) = x. GH = x. It is known that the right Line BG is every where equal to the Sum of the Arch AD and the right Sine BD; whence 'tis manifeft, that the Fluxion IG is alfo the Aggregate of the Fluxions of the Arch AD and the right Sine BD. But the Fluxion of the Arch AD was found

 $\frac{1}{2} dx$  and the Fluxion of the right

 $\sqrt{dx} = xx$ , Sine BD will be found to be dx - 2xx

 $2 \sqrt{dx - xx}$ ; therefore IG = dx - xx and therefore

1/2rx - xx.

Miscellanea Curiosa. 125

 $IH^{4}(= IG^{4} + GH^{4}) = ddxx - dxxx; \text{ from}$ 

dx - xx

		$x\sqrt{dd} - dx$			12	-	12.
whence	IH =		II	$x\sqrt{d} \equiv$	d	x	x,
		$\sqrt{dx - xx}$		$\sqrt{x}$			
			1 1	-			

and confequently AI =  $2d x = 2\sqrt{dx} = 2AD$ . This Conclution may alfo very eafily be deduc'd from the known Property of the Tangent. For fince the little part of it, as IH, is always parallel to the Chord AD, the Triangles IGH, ABD are fimilar, whence AB: AD:: GH: IH, that is,  $x: \sqrt{dx} :: x:$   $\frac{1}{2} = \frac{1}{2}$ .  $x\sqrt{dx}$ , therefore IH =  $x\sqrt{dx} dx x$ . By x

the help of the Fluxion IH alfo, we may investigate the Area of the Cycloid. The Fluxion of the Area AEI, is the Rectangle EIG =  $dxx - x^2x = x\sqrt{dx - xx}$ . But the  $\sqrt{dx - xx}$ 

Fluxion of the Portion ABD is the fame; therefore the Area AEI and the correspondent Portion (of the Circle) ABD, are always equal.

K 4

Let

Let AB (Fig. 3.) be the Curve of the Parabola, whole Axis is AF, Parameter a: let AE = x, EB = y, AB = z, BD = x, DC = y, BC=z. The Equation expressing the Nature of the Parabola, being ax = yy, we have ax = 2yy, whence x = 2yy; but BC9 = BD9 -1- CD4, that is  $zz = xx - 1 - yy = \frac{4y^2 y}{-1}$ -+- yy ==  $4y^2yy + aayy$ , and therefore  $z = y \sqrt{4y^2} + aa$ aa a  $= \sqrt{y^2 + \frac{1}{4}aa}$ . If now by this Expref-1 1 fion  $y \sqrt{y^2 + \frac{1}{4}}$  as be thrown into an infinite 1 a Series, the Curve AB will eafily be known. It appears farther, that giving an Hyperbolical Space, this Curve is alfo given, and vice ver fa. For  $\frac{1}{2}az = y \sqrt{y^2 + \frac{1}{4}aa}$ , and confequently  $\frac{1}{2}$  az is the Space whole Fluxion is  $y \sqrt{y^2 + \frac{1}{4}} aa$ . But fuch a Space is no other than the Exteriour (Equilateral) Hyperbola ABEG, whole Semiaxis  $AB = \frac{1}{2}a$ , its Abfciffe AE = y, and its Ordinate EG = x.

For

137

For the Menfuration of a furface defcrib'd by the Conversion of a Curve round its Axis; we are to affume for the Fluxion of it, a Cylindrick Superficies, whofe Altitude is the Fluxion of the Curve, and whofe diffance from the Axis is the Ordinate Applicate corresponding to that Fluxion. Ex. gr. Let AC be the Arch of a Circle, which turning round the Axis AD, generates a spherical Superficies, which we would measure. Now DC the Fluxion of the Arch is already found to be

= _____, which if we multiply by the  $\sqrt{2rx - xx}$ 

rx

Periphery belonging to the Radius BC, that is, by  $\frac{c}{r}\sqrt{2rx-xx}$  (putting  $\frac{c}{r}$  the Ratio of the Circumference to the Radius) we shall

have cx for the Fluxion of the fpherical Superficies, and confequently that Superficies it felf, is cx.

As for Centers of Gravity; having gotten the Fluxion of the Solid or Surface, and multiplied the fame into its diffance from the Vertex, the flowing Quantity must be found, which divided by the Solid or Surface it felf, the Quotient will shew the distance of the Center of Gravity from the Vertex. Thus to find the Center of Gravity of all the Paraboloids; their Fluxion is thus generally expressed

Distantly Google

expressed  $x^{\frac{n}{n}}$ , which multiplied by x, makes m+1 x"x, the flowing Quantity of which, m-1-2 n divided by the Area of viz. m-1-2n m-1-2# the Paraboloid  $-x^n$ , gives  $-x^n$ , m +- 11 m+2n the distance of the Center of Gravity from the Vertex. The Center of Gravity of a Portion of a Sphere, is found after the fame manner. For its Fluxion  $4 dxx - x^2x$  multiplied into x makes  $4 dx^{2}x - x^{3}x$ , whole flowing Quantity  $4\frac{1}{3}dx^3 - \frac{1}{4}x^4$ , divided by the folid n Content of the Portion, viz.  $4\frac{1}{2}dxx - \frac{1}{3}x^3$ , gives  $\frac{\frac{1}{3}d - \frac{1}{4}x}{\frac{1}{2}d - \frac{1}{3}x}$ , or  $\frac{4d - 3x}{6d - 4x}$ , the distance of the Center of Gravity from the Vertex. My

Dig under Google

Miscellanea Curiosa. 139 My defign here was not to be large, and purfue all the Difficulties that may occur. 'Tis fufficient to have made a beginning, and led the Way to those greater Things.

A Me-

A Method of Squaring Some Sorts of Curves, or Reducing them to more fimple Curves. By A. De Moivre, R.S.S. E T A be the Area of a Curve, whofe Abfciffe is x, and Ordinate Applicate  $x \sqrt{dx - xx}$ . Let B be the Area of a Curve, whofe Abfciffe is the fame with the former, and its Ordinate  $x \sqrt[m-n]{dx - xx}$ . Put  $\sqrt{dx - xx} = y$ . Then will A = dB into  $\frac{2m-1}{2m-4}$  into  $\frac{2m-1}{2m-2}$  into  $\frac{2m-3}{2m}$  into  $\frac{2m-5}{2m-2}$  & c.=P  $-\frac{1}{m+2}x^{m-1}y^3 = -Q$  $-\frac{d}{m+1} \text{ into } \frac{2m+1}{2m+4} x^{m-2} y^{3} = -R$  $-\frac{d^2}{m} \text{ into } \frac{2m-1}{2m-1-4} \text{ into } \frac{2m-1}{2m-2} x^{m-3} y^3 = -S$  $-\frac{d^3}{m-1}$  into  $\frac{2m-1}{2m-4}$  into  $\frac{2m-1}{2m-2}$  into  $\frac{2m-3}{2m}$  $x^{m-4}y^3 = -T$ , &c.

Where we are to Note, 1. That n is fuppos'd to be an Integer and Affirmative Number. 2. That the Quantity  $d^nB$ , in the Series

141

Series expressed by P, is to be multiplied into as many Terms as there are Unities in *n*. 3. That as many of the following Series expressed by  $-Q_{,-}R_{,-}S_{,-}T_{,}$  &c. are to be taken, as there are Unities in *n*. Which to illustrate by an Example or two: If n = 1, the I fay that  $A = d^{n}B$  into  $\frac{2m+1}{2m+4} - \frac{1}{m+2}x^{m-1}y^{3}$ ; and if n = 2, then  $A = d^{n}B$  into  $\frac{2m+1}{2m+4}$  into  $\frac{2m-2}{2m+1} - \frac{1}{m+2}x^{m-1}y^{3} - \frac{d}{m+1}$  into  $\frac{2m+1}{2m+4}$  $x^{m-2}y^{3}$ . 4. That if y be put  $= \sqrt{dx - xx}$ , then A will = Q - R + S - T, &c.  $\pm P$ .

#### COROL. I.

If m be put = to any Term of the following Series  $-\frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{1}{2}, \frac{2}{2}, \frac{2}{2}, & \&c. then the$ Quadrature of the Curve whofe Ordinate is 272  $x \sqrt{dx - xx}$ , or  $x \sqrt{dx + xx}$ , will be expreffed in finite Terms, and be found by our Series. To illustrate which by an Example or two; Let it be required to find the Area of a Curve whofe Ordinate is  $\sqrt{dx - xx}$ . Let us imagine this Curve x to be compar'd with another Curve whofe Ordinate is  $x = \sqrt{dx - xx}$ . Now because in this cafe n = 1, therefore will  $A = d^n B$  into  $\frac{2m+1}{2m-1-4} - \frac{1}{m+2} x^{m-1} y^3$ ; but  $m = -\frac{1}{2}$ , therefore

fore  $2m + 1 \equiv 0$ , and therefore  $A \equiv \frac{1}{m + 2}$  $x^{m-1}y^{3} \equiv -\frac{2y^{3}}{3\sqrt{x^{3}}}$ .

It is here to be observed, that the Area thus found, is sometimes deficient from the true Area, by a given Quantity, or exceeds it by that fame given Quantity. And in order to find that Defect or Excess, let the Area found be supposed to be encreased or diminiss x = 0, let the Area increased or diminiss x = 0, let the Area increased or diminiss of the fupposed x = 0. Thus in the prefent case, we shall find  $q = \frac{2}{3} d\sqrt{d}$ , and confequently  $A = \frac{2}{3} d\sqrt{d} - 2y^3$ 

31/23.

#### COROL. II.

If *n* be put equal to any Term of the following Series, 3, 4, 5, 6, 7, &c. then the Quadrature of the Curve whole Ordinate is  $\sqrt{x} \sqrt{dx - xx}$ , or  $x \sqrt{dx - xx}$ , is exprefied in finite Terms, and is found by our Series.

Let the Area of the Curve be to be found, whofe Ordinate is  $x \sqrt[3]{dx - xx}$ . Suppofe it to be compared with the Area of a Circle, which call A. Then will  $m \equiv o$ ,  $n \equiv 3$ , and fo  $A \equiv P - Q - R - S$ . But funce, in the Denominator of the third Term by which  $d^{m}B$  is multiplied, there is found 2m,

2*m*, a Quantity infinitely fmall, or rather nothing; the Quantity express'd by P is Infinite; and for the fame reason the Quantity express'd by -S is Infinite, and fo the Quantities A, -Q, -R, do vanish. Therefore P=S, and dividing the Equation by  $\frac{2m-1}{2m-4}$  into  $\frac{2m-1}{2m-2}$ , we have  $d^{n}B$  into  $\frac{2m-3}{2m}$  $= \frac{dd}{m}x^{m-3}y^{3}$ , or dB into  $\frac{2m-3}{2} = ddx^{m-3}y^{3}$ ; and putting o and 3 for m and n, there will be dB into  $-\frac{3}{2} = \frac{y^{3}}{x^{3}}$ , or  $B = -\frac{2y^{3}}{3x^{3}}$ .

#### COROL. III.

If *m* be put equal to any Term of the following Series, -2, -1, 0, 1, 2, 3, 4, 5, &c. the Quadrature of the Curve, whole Ordimate is  $x \sqrt{dx - xx}$ , depends upon the Quadrature of the Circle. But the Area of the Curve, whole Ordinate is  $x \sqrt{dx - xx}$ , depends upon the Quadrature of the Hyperbola; and the relation of that Curve to the Circle of the Hyperbola, is exhibited by our Series in finite Terms.

#### COROL.

143

Distanced by Google

#### COROL. IV.

If *m* be expounded by any other Termi, different from them already mention'd; the Curve whofe Ordinate is  $x \sqrt[m]{dx - xx}$ , or  $x \sqrt[m]{dx + xx}$ , is neither exactly fquar'd, nor does it depend upon the Circle or Hyperbola, but is reduced to a more fimple Curve by our Series.

#### THEOR. II.

Let A be the Area of a Curve whole Ab- $x^m$ 

fciffe is x, and Ordinate  $\sqrt{dx - xx}$ . Let B be the Area of a Curve whole Absciffe is the fame with the former, but the Ordinate  $x^{m-n}$ 

 $\sqrt{dx - xx}$ . Let  $\sqrt{dx - xx} = y$ . Then will

A =

Miscellanea Curiosa. 145 A=dⁿB into  $\frac{2m-1}{2m}$  into  $\frac{2m-3}{2m-2}$  into  $\frac{2m-5}{2m-4}$  into 2m-7 &c. = P.  $-\frac{1}{m}x^{m-1}y = -Q.$  $-\frac{d}{m-1}$  into  $\frac{2m-1}{2m} x^{m-2} y = -R$ .  $-\frac{dd}{m-2} \operatorname{into} \frac{2m-1}{2m} \operatorname{into} \frac{2m-3}{2m-2} x^{m-3} y = -S.$  $-\frac{d^3}{m-3}$  into  $\frac{2m-1}{2m}$  into  $\frac{2m-3}{2m-2}$  into  $\frac{2m-5}{2m-4}$  $x^{m-4}y = -T$ . &c.

The Observations to the first Theorem, take place here also, as in what follows.

#### COROL. I.

If m be put equal to any Term of the following Series,  $\frac{1}{2}$ ,  $\frac{1}{2}$ ,  $\frac{1}{2}$ ,  $\frac{1}{2}$ ,  $\frac{2}{3}$ , &c. the Quadrature of the Curve whole Ordinate is  $\frac{m}{x}$ , or  $\frac{m}{\sqrt{dx-xx}}$ , is expressed in  $\sqrt{dx-xx}$ ,  $\sqrt{dx-xx}$ finite Terms, and exhibited by this Series.

L

÷.

COROL

#### COROL. II.

If *n* be put equal to any Term of the following Series, 1, 2, 3, 4, 5, 6, 7, &c. Every Curve whofe Ordinate is  $\frac{x^{-n}}{\sqrt{dx - xx}}$ , or  $\sqrt{\frac{x^{-n}}{\sqrt{dx - xx}}}$ , is fquared by this Series in finite

 $\sqrt{dx + xx}$ Terms.

#### COROL. III.

If *m* be expounded by any Term of the following Series, 0, 1, 2, 3, 4, 5, 6, 7, &c. *m* the Curve whofe Ordinate is  $\frac{x}{\sqrt{dx-xx}}$ , depends upon the Quadrature of the Circle. But the Curve whofe Ordinate is  $\frac{x^m}{\sqrt{dx-xx}}$ depends upon the Quadrature of the Hyperbola. For if on the Center C, (Fig.20.) the Diameter AB = d, the Circle AEB be defcribed, and

AD be taken = x, also erecting the Perpendi-

cular DE, the Line CE be drawn. Then the Se-

Miscellanea Curiofa. 147 Sector AEC divided by  $\frac{1}{6} dd$  is equal to the Area of the Curve whose Ordinate is  $x^{0}$ 

After the fame manner, if on  $\sqrt{dx-xx}$ 

the Center C, and the Transverse Axis AB=d, the Equilateral Hyperbola AE be described, and taking AD=x, and erecting DE at right Angles, and joining CE, the Sector ACE divided by  $\frac{1}{8} dd$  is equal to the Area of the Curve whose Ordinate is  $x^{0}$ 

Vdx-1-xx

### COROL. IV.

If *m* be equal to any Term, that does not fall into the foregoing Limitations, then the Curve whofe Ordinate is  $\frac{x^m}{\sqrt{dx+xx}}$ , is nei- $\sqrt{dx+xx}$ ther exactly fquar'd, nor does it depend upon the Circle or Hyperbola, but is reduced

to a more fimple Curve.

#### L₂ THEO-

Dention by Google

#### THEOR. III.

Let A be the Area of a Curve whole Abfciffe is x, and its Ordinate Applicate  $x \sqrt{rr-xx}$ ; let B be the Area of a Curve whole Abfciffe is also x, and its Ordinate  $x \sqrt{rr-xx}$ .

Let  $\sqrt{rr-xx} = y$ . Then will  $A = r^{n_2}B$  into  $\frac{m-1}{m+1-1}$  into  $\frac{m-3}{m}$  into  $\frac{m-5}{m-2}$  into  $\frac{m-7}{m-4}$  &c.=P.  $-\frac{1}{m+2}x^m - \frac{1}{y^3} = -Q$ .  $-\frac{rr}{m}$  into  $\frac{m-1}{m+2}x^{m-3}y^3 = -R$ .  $-\frac{r^4}{m-2}$  into  $\frac{m-2}{m+2}$  into  $\frac{m-3}{m}x^{m-5}y^3 = -S$ . &c.

#### COROL. I.

If *m* be expounded by any Term of the following Series, 1, 3, 5, 7, 9, &c. the Quadrature of the Curve whofe Ordinate is,  ${}^{m}\sqrt{rr-xx}$ , or  ${}^{m}\sqrt{rr-xx}$ , is had in finite Terms, and that by the help of this Theorem.

### COROL

148

#### COROL. II.

If *n* be expounded by any Term of the following Series, 2, 3, 4, 5, 6, &c. then the Curve whole Ordinate is  $x = \sqrt{rr-xx}$ , or  $x = \sqrt{rr-|-xx}$ , is exactly fquar'd by this Theorem.

#### COROL. III.

If *m* be expounded by any Term of the following Series, -2, 0, 2, 4, 6, 8, &c. then the Quadrature of the Curve whole Ordinate is  $x \sqrt{rr-xx}$ , depends upon the Circle, but the Quadrature of the Curve whole Ordinate is  $x \sqrt{rr-xx}$ , depends upon the Hyperbola.

#### COROL. IV.

If *m* be expounded by any Term different from those already taken notice of; then the Curve whose Ordinate is  $x \sqrt{rr-xx}$ , or  $x \sqrt{rr-xx}$ , depends neither upon the Circle nor the Hyperbola, but is reduced to a more fimple Curve.

L 3

THEO-

#### THEOR. IV.

Let A be the Area of a Curve whole Abfciffe is x, and whofe Ordinate is let B be the Area of a Curve whofe Abscille  $\sqrt{rr-xx}$ is alfo x, and its Ordinate Then will  $A = r^{2n}B$  into  $\frac{m-1}{m}$  into  $\frac{m-3}{m-4}$  into  $\frac{m-5}{m-4}$ into  $\frac{m-7}{m-6}$  &c. = P.  $-\frac{1}{x}x^{m-1}y=-Q.$  $-\frac{rr}{m-2} \operatorname{into} \frac{m-1}{m} x^{m-3} y = -R.$  $-\frac{r^4}{m-4} \operatorname{into} \frac{m-1}{m} \operatorname{into} \frac{m-3}{m-2} x^{m-5} y = -S.$  $-\frac{r^6}{m-6}$  into  $\frac{m-1}{m}$  into  $\frac{m-3}{m-2}$  into  $\frac{m-5}{m-4} x^{m-7} y$ =- T. &c. the second states of the

· …

COROL.

COROL. I.

151

If *m* be expounded by any Term of the following Series, 1, 3, 5, 7, 9, &c. the Quadrature of the Curve, whole Ordinate is  $m_{x}^{m}$ 

, is obtained in finite Terms by this  $\sqrt{rr}$ -J-xx Theorem.

#### COROL. II.

If *n* be expounded by any Term of the following Series, 1, 2, 3, 4, 5, 6, &c. the  $x^{-2n}$ Curve, whole Ordinate is  $\sqrt{rr-xx}$  or  $\sqrt{rr-xx}$ 

, is fquar'd exactly by this Theo- $\sqrt{rr+xx}$  rem.

#### COROL. III.

If *m* be expounded by any Term of the following Series, 0, 2, 4, 6, 8, 10, &c. the Quadrature of the Curve, whofe Ordinate L 4 is

is  $\frac{x}{\sqrt{rr-xx}}$ , depends upon the Quadrature  $\sqrt{rr-xx}$ 

of the Circle. For if on the Center C, and the Radius CA=r, the Circle AEG be defcrib'd, and taking CD=x, DE be erected perpendicular to CD, and CE be drawn; then the Sector CAE divided by  $\frac{1}{2}rr$ , is equal to the Area of the Curve, whofe Or-

dinate is  $\frac{\pi}{\sqrt{\pi r}}$ . In like manner, (Fig.21.)

if on the Center C, and the Semitransverse Axis CA=r, the Equilateral Hyperbola EAM be deforib'd, then drawing CF perpendicular to CA, equal to x, FE parallel to the Axis till it meets the Hyperbola in E, and join CE; then the Hyperbolical Sector ACE divided by  $\frac{1}{2}rr$ , is equal to the Area of the

Curve, whofe Ordinate is  $\sqrt{rr-r}$ 

#### COROL. IV.

If *m* be expounded by any Term different from the foregoing, then the Curve, *x*^{*m*} whofe Ordinate is  $-\frac{x}{\sqrt{rx-t-xx}}$ , is neither ex-

actly

atly squar'd, nor does it depend upon the Circle or the Hpperbola, but is reduced to a more simple Curve.

#### THEOR. V.

Let A be the Area of a Curve, whole Abfciffe is x, and its Ordinate ----; let B be d-r the Area of a Curve, whofe Absciffe also is xm-n x, and its Ordinate -----. Then will A= dan ddx.m-2  $dx^{m-1}$ ----- &c.  $d^{n}B$ m-1 m m-2 2.m Let the Ordinate be -----, and then the , d--x  $dx^{m-1}$ 2m  $ddx^{m-2}$ ----- &c. --Area A will = --mil 11:---2 m  $d^m \mathbf{B}$ .

### COROL.

If m be expounded by any Term of the following Series, 1, 2, 3, 4, 5, 6, &c. the Quadrature of the Curve, whose Ordinate  $x^m$  is  $\frac{x^m}{d-x}$  or  $\frac{x^m}{d-x}$ , depends upon the Quadra-

ture

Distand by Google

ture of the Hyperbola. For (See Fig. 22.) drawing DE, EF at right Angles, take EG =d, and draw GH at right Angles to EF, and equal to it. Within the Afymptotes DE, EF, let an Hyperbola be defcrib'd, paffing through the point H; which done, take GK=x, towards E in the first Cafe, and towards F in the freend; and draw the Ordinate KL. Then the Area HGKL divided by dd, is equal to the Area of the Curve,  $x^0 \qquad x^0$ 

whose Ordinate is  $\frac{1}{d-x}$  or  $\frac{1}{d-x}$ . Hence the

Solid generated by a Portion of the Ciffoid, while it turns about the Diameter of the Generating Circle, is exhibited in finite Terms, fuppofing the Quadrature of the Hyperbola.

#### THEOR. VI.

Let A be the Area of a Curve, whole Abfciffe is x, and Ordinate -; let B be rr- -xx the Area of a Curve, whole Absciffe is also 232-22 x, and Ordinate . Then will the rr---xx  $\frac{x^{m-1}}{m-1} \frac{rrx^{m-3}}{m-3}$ r4xm-5 rrxm-3 Area A = - ----- &c. m-5 -+ r2"B.

#### COROL.

### COROL.

If *m* be expounded by any Term of the following Series, 0, 2, 4, 6, 8, &c. the Quadrature of the Curve, whose Ordinate is  $x^{m}$ 

 $\frac{1}{rr-1-xx}$ , depends upon the Rectification of

the Arch of a Circle. For defcribing the Circle AEG on the Center C, with the Radius CA=r, draw the Tangent AK=x, and join CK meeting the Periphery in E; then the Arch AE divided by rr is equal to the  $x^{0}$ 

Area of the Curve, whole Ordinate is  $\frac{1}{rr-1-xx}$ .

#### General Corollaries to these six Theorems.

E Very Mechanick Curve (whole Quadrature depends upon any one of that Infinite Number of Curves, whole Ordinates have the following Forms,  $x \sqrt[m]{dx + xx}$ ,  $x^m$ ,  $\sqrt[m]{x^m}$ ,  $x^m \sqrt{rr + xx}$ ,  $x^m \sqrt{rr + xx}$ ,  $\frac{d+x}{d+x}$ ,  $\sqrt[m]{dx + xx}$ ,  $x^m \sqrt{rr + xx}$ ,  $\frac{d+x}{d+x}$ ,  $x^m$ ) may be fquar'd by these Series. It may fuffice to illustrate this by an Example.

Sup-

Dattion by Google

Supposing the Cube of the Arch of a Circle (corresponding to the versed Sine) to be the Ordinate of a Curve, whose Abscille is the fame versed Sine; let it be required to find the Area of this Curve.

Let the Abfciffe be x, the Circular Arch v; then the Fluxion of the Area is  $v^3x$ . Let the Area be  $v^3x - q$ . Therefore  $v^3x + 3v^2$  $vx - q = v^3x$ , whence  $q = 3v^2vx$ . But  $u = \frac{dx}{2\sqrt{dx - xx}}$ , therefore  $q = \frac{3dv^2xx}{2\sqrt{dx - xx}}$ . But

(by Theor. 2.)  $\frac{xx}{\sqrt{dx-xx}} = \frac{dx}{2\sqrt{dx-xx}}$ 

y = v - y; and confequently  $q = \frac{1}{2}dv^2v - \frac{1}{2}$  $dv^2y$ , therefore  $q = \frac{1}{2}dv^3 - s \cdot dv^2y$ . Therefore now we are come to this, that the Fluent, of the Expression  $\frac{1}{2}dv^2y$ , is to be found.

Let this Fluent be  $\frac{3}{2}dv^2y$ —r.

Therefore  $\frac{1}{2}dv^2y - |-3dvvy - r = \frac{3}{2}dv^2y$ .

And confequently  $r=3dvvy=\frac{3}{2}ddvx$ . Let  $r=\frac{3}{2}ddvx$ .

Therefore  $\frac{1}{2}ddvx = \frac{1}{2}ddvx + \frac{1}{2}ddxv - s$ .

And

And confequently  $s = \frac{3}{2} ddxv = \frac{3d^3xx}{4\sqrt{dx-xx}}$ =  $\frac{3}{4}d^3v = \frac{3}{4}d^3y$  (by Theor. 2.) Therefore now

 $= \frac{1}{4}d^3v - \frac{1}{4}d^3y.$  And confequently the Area fought, is  $= v^3x - \frac{1}{2}dv^3 - \frac{1}{2}dv^2y - \frac{1}{2}ddvx - \frac{1}{4}d^3$  $v - \frac{3}{4}d^3y.$ 

Since the Solids and Surfaces generated by the Rotations of Curve Lines, as alfo the Lengths of Curves, and the Centers of Gravity of all thefe, do depend upon the Quadratures of Curves; 'tis plain, that thefe are eafily obtain'd too, if they depend upon the foremention'd Curves.

After I had compos'd these Theorems, and shewn them to the Celebrated Mr. Nemton, (as the supream Judge in all Matters of this Nature) he was pleas'd to give me a sight of some Papers of his, by which I find that he has a long time been Master of a Method, by which any Trinomial Equation (expressing the Nature of a Curve) being given, that Curve is either squar'd, or reduced to a more simple one.

And 'twere to be wish'd, that he thought fit to communicate to the Publick, not only those Things which he has relating to these Matters, but others also of his Noble Inventions, which are not a small Number neither. And I believe this is not my Wish alone, but that of the whole Learned World besides.

I make no question but those Learned Perfons (whose Writings in the Asta Eruditorum and

and otherwhere, have tended fo much to the Advancement of Mathematicks) have Methods not unlike to this of mine; and therefore I afcribe no more to my felf in this Matter, than only that I found out thefe Theorems, not knowing whether any Body elfe had done fo before or no; and reduced them into fo eafie a Form, that the whole Calculus relating to them, might be taken in, as it were, at one View.

But before I make an end of Writing, I think it improper, if (having not had an Opportunity fooner) I make fome little reply to the Famous Mr. Leibnitz's Animadverfions upon my Series for finding the Root of an Infinite Equation.

That Excellent Perfon thinks this Series not to be General enough, as not reaching the Cafes where z and y are multiplied into one another; upon which account he fubfitutes another Series in the room of it, which he afferts is infinitely more General. But that which led him into this finall Miftake, I guess to be this, that he took the Quantities a, b, c, d; &c. for given Quantities, whereas they were to be us'd indifferently, either for given or indeterminate ones.

But I shall add one Example to shew that my Series extends to all Cafes. Let Equation be  $yz - z^3 = y^3$ .

În our T	1  let  a = ny; b = 0.
b=0, i=1.	her let $g = yy$ ,
1.000	24

Then in  $(12y^{12})$ 

nio

fe will z=

Two Problems; viz. concerning the Solid of Least Resistance, and the Curve of Swiftest Descent.

Solv'd by J. Craig.

#### LEMMA.

TO find the Proportion between the Refiftance made to the Rectangular Triangle AIG, and that made to the circumfcrib'd Rectangle AIGg, while each moves in a Fluid, in the direction of the Line IA, from I towards X.

From any point B let there be drawn BC perpendicular to AG, Bb parallel to AI, and BM perpendicular to AI. Then in bB take  $bH = CM^2$  and bE = BC; and thro' the BC

points H and E, draw the Lines HA, EA, which being produced cut Gg in K and F. I fay the Refiftance of the Triangle AIG is to the Refiftance of the Reftangle AIGg, as the Area of the Triangle AKG, to the Area of the Triangle AFg. And alfo, that the Refiftance upon any part of the Line AG, is to the Refiftance upon the correspondent part of the Line Ag

Ag (ex. gr. upon AB and Ab) as the Area AHB to the Area AEB. The Demonfitration of which depends upon a General Theorem, which I deduced very eafily from *Prop.* 35. Newt. p. 324.

#### COROL. I.

Let BG, bg, be infinitely fmall parts of the Lines AG, Ag, and let bB be produced to L; I fay, that the Refiftance upon BG (which call e) is to the Refiftance upon bg (which call E) as  $GL^2: GB^2$ .

For e: E:: KHgb: FEgb; that is, e: E:: bg x bH: bg x bE (by the foregoing *Lemma*) therefore e: E:: bH: bE; that is, e: E::  $CM^2$ : BC ::  $CM^2$ : BC². But  $CM^2$ : BC²:: BC

GL²: GB² (because of the similar Triangles BMC, GLB.) Therefore c: E:: GL²: GB². Q: E: D.

#### COROL. II.

The Refiftance upon the infinitely finall part GB, is =  $GL^3$ . For if all the infinite- $\overline{GB^2}$ 

ly finall parts in the Line Ag (as bg) be fuppos'd equal, then the Refiftance upon bg, may be express'd by bg, that is E= bg, and fo E=GL. Therefore (by Cor. 1.) e:GL:: GL²:GB², whence e= GL³. Q: E: D.

GB²

#### COROL

#### COROL. III.

Let r be the Radius, and c the Circumference of any Circle. I fay, that the Refistance upon the Conick Surface generated by the Rotation of the Lineola GB about AI, is equal to the Product of c x BM into r

GL³. For the Reliftance upon that Conick GB²

Surface, is equal to all the Reliftances upon the Lincola GB, that is, to all the e; that is, equal to the Circumference of the Circle (whofe Radius is BM) multiplied into e; that is, the Refiftance upon that Conick c x BM

Surface, is equal to _____ x e, and confer

c x BM GL³

quently (by Corol. 2.) equal to ------ × --GB² r

Q: E: D.

#### PROB. I.

To find a Curve Line, by the Rotation of which a Round Solid shall be gene-rated, that, while 'tis moved in a Fluid Medium, in the Direction of its Axis, (ball meet with the least Resistance.

#### (Fig. 24.)

Let OG, GB, be two infinitely fmall Particles in the Curve fought, which rould about its Axis, will produce the Solid of leaft M

least Refistance. Draw BM, GP, perpendicular to AQ, also BL, GN, parallel to AQ, and ON, parallel to BM.

#### $c \times BM \times GL^3$

Now - , is the Refiftance upr x GB²

on the Surface generated by the Rotation of  $c \times GP \times ON^3$ 

the Lineola GB about AQ, and  $r \times GO^2$ ,

is the Reliftance upon the Surface generated OG, in like manner (by Cor. 3.) And the Sum of both these Resistances must be a Minimum, viz.

 $\frac{c \times BM \times GL^3}{r \times GB^2} + \frac{c \times GP \times ON^3}{r \times OG^2} = a Mini-$ 

mum.

And confequently in the Line RS (fo parallel to AQ that ON = GL) the point G is to be fought, fuch, that this may happen; which, fuppofing the points O and B to be fix'd, will be eafily found by the common Method, de Maximis & Minimis. And profecuting the Calculus, we fhall come at laft to this Equation,  $\frac{BM \times BL}{BG^4} = \frac{GP \times NG}{OG^4}$ ; BM x BL

whence 'tis plain that  $\frac{BG^4}{BG^4} = a \text{ con-}$ 

ftant Quantity. So that if the Abfcille  $AM \equiv x$ , and the Ordinate  $BM \equiv y$ , then will  $BL \equiv dx$ ,  $LG \equiv dy$  (which I have fupposed

pos'd conftant every where in this Calculus) and confequently  $BG^2 = dx^2 + dy^2$ , whence ydx

-= a constant Quantity. Let a dxdx-1-dydy2

be any conftant Quantity, and confequently (to observe the Law of Homogeneals) we

have  $\frac{ydx}{dxdx+dydy^2} = \frac{a}{dy^3}$ , as has been found

by the Illustrious L' Hospitall, and the celebrated Jo. Bernoulli.

#### PROB. II.

#### To find the Line of Swiftest Descent.

(Fig. 25.)

Let BC, CD, be two infinitely fmall Particles in the Curve fought. Now this Curve ought to be of fuch a Nature, that, fuppofing a Body to have fallen from the Horizontal Line AQ, it may pass from B to D in the *fborteft Time*. Therefore we are to find out the Point C (in the Line RS drawn in fuch a manner parallel to AQ, that the differences of the Ordinates GC, DE, may be equal) fuch that this may come to país.

Now the Velocity in C is  $\sqrt{LC}$ , and that BC in D is  $\sqrt{QD}$ ; therefore — is the Time VLC. of

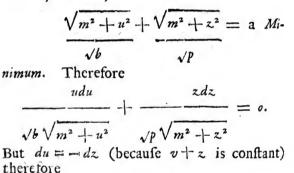
M 2

of Defcent thro' BC,  $\frac{CD}{\sqrt{QD}}$  is the Time of

Defcent thro' CD (by Prop. 54. pag. 158. Newt.) Therefore the point C ought to be

fuch that  $\frac{BC}{\sqrt{LC}} + \frac{CD}{\sqrt{QD}}$  may be a Minimum.

Supposing the points B and D to be fix'd, let the conftant Quantities GC = DE = m, LC = b, QD = p; the *Indeterminate* Quantities BG = u, CE = z; whence



 $\frac{u}{\sqrt{b}\sqrt{m^2 + u^2}} = \frac{z}{\sqrt{p}\sqrt{m^2 + z^2}}; \text{ whence 'tis}$ 

manifest that ----= = a constant  $\sqrt{b}\sqrt{m^2 + u^2}$ 

Quantity. Now let the Abfciffe AL = x, the Ordinate LC = y, and to BG = dx, GG

Distance Google

GC = dy,  $BC = \sqrt{dx^2 + dy^2}$ , and let *a* be any conftant Quantity. Then fhall  $\frac{dx}{dx} = \frac{1}{\sqrt{a}}$ , whence  $dx \sqrt{a} = \sqrt{y} \times \sqrt{y}\sqrt{dx^2 + dy^2}$ . But now in all Curves, 'tis  $dx:: \sqrt{dx^2 + dy^2}$ . But now in all Curves, 'tis  $dx:: \sqrt{dx^2 + dy^2}$ . But now in all Curves, 'tis to the Tangent. Therefore the Nature of the Curve fought is fuch, that its Subtangent, is to the Tangent, as  $\sqrt{a}: \sqrt{y}$ , which that it is a Property of the Cycloid, is known to all,

that know that the Tangent of the Cycloid, is parallel to the Chord of the Conterminal Arch, in the Generating Circle, whofe Diameter is *a*, and whofe Vertex is downwards.

And with the like eafe, I can find the Curve of the Swiftest Defcent, in any other Hypothesis of Gravity.

M 3

The

## The Quadrature of the Logarithmical Curve.

## (Fig. 26.)

By J. Craig.

ET ONF be the Logarithmical Curve, whofe Afymptote is AR, in which let fuch a point A be taken, as that the first Ordinate AO may be equal to the Subtangent or Unity. 'Tis requir'd to find the Area of the Curvilineal Space AONM comprehended under the two Ordinates AO, MN, the Abfciffe AM, and the Curve ON. From O draw OE parallel to AM and cutting MN in E; I fay, that the Rectangle under the Segments ME, EN, is equal to the Space fought. Demonstration. Let the Ordinate MN = Z, Subtangent AO or ME =s; and to the Axis AR let another Curve HGE be constructed, whose Equation shall be  $2sz = x^2$ , its Ordinate GM being = x. I fay, that this Curve is the Quadratis of the Logarithmical Curve (according to the Principles of my Method) viz. its Subnormal is respectively equal to the Ordinate of this, as is plain from the Calculus of that Method. Therefore (according to what I have shewn in another place) if to the point G we draw GC perpendicular and equal to GM, as alfo HD parallel to GC, and meeting the Lines GM, CM, in B and D; then will the

Dailtream Google

the Trapezium GBDC = AONM. But GBDC = GMC - BMD =  $\frac{1}{2}x^2 - \frac{1}{2}BM^q = SZ - \frac{1}{2}HA^q$ ; but HA =  $\sqrt{2}$  AO^q from the Nature of the Curve HGQ, therefore GBDC = SZ - AO^q = AO × MN - AO^q = AO × MN - AO = ME × MN - ME = ME × EN. Therefore alfo AONM = ME × EN. Q: E.D.

When I applied my Method to thefe fort of Figures, I found that a Miftake had fome way or other crept into M. Bernoulli's Calculus. For in his most excellent Tract of the Principles of the Differential Calculus, he affigns to the Figure whose Equation is 2yyly - yy

 $a^{2} = y^{3}$ , this for its Quadrature, viz.  $\frac{1}{la}$ ;

2yyly - yy

Theo-

Dignation Google

whereas the Area of that Figure is  $\frac{4}{4a}$ ;

where y denotes the Abscisse and z the Ordinate.

1 23.4

at felt in

17 22761

this :

H.

1. 15%3 JAY

A Theorem concerning the Proportion of the Time that a heavy Body spends in descending thro' a right Line joining two given Points, to the (fhortest) Time, in which it passes from the one to the other of these Points, by the Arch of a Cycloid.

#### THEOREM.

(Fig. 27.)

IN the Cycloid AVD, whofe Bafis AD is parallel to the Horizon, and the Vertex V turn'd downwards, if from A be drawn the right Line AB meeting the Cycloid in any point as B, from whence is drawn BC perpendicular to the Curve of the Cycloid in B, and AC be let fall perpendicular to BC from the point A: Then the Time that a Body at reft in A, fpends in defcending thro' AB (by the force of its Gravity) is to the Time that it fpends in falling thro' the Curve AVB, as AB to AC.

Thro' B draw BL parallel to the Axis of the Cycloid VE, and BK parallel to the Bafis AD, meeting the Axis in G, and the Circle (whofe Diameter is EV) in F and H, and the Cycloid it felf in K. Draw the right

Inhilized by Goo

right Line EF, which from the Nature of the Cycloid is parallel to BC; whence BM is = EF, and EM = BF = the Arch VF from the Nature of the Cycloid; and confequently AM is = the Arch EHVF.

By Proposition 25. Part II. Horolog. Ofcillat. Hugen. the Time in which a Body at reft in A defcribes the Cycloidal Arch AV, is to the Time of Defcent thro' EV, as the half Circumference to the Diameter.

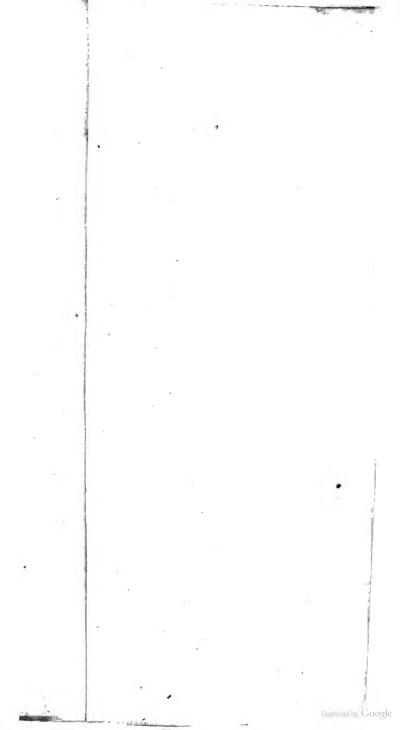
And (by the last Proposition of the foremention'd Part) the Time of Descent thro' VB, after the Descent thro' AV (which is equal to the Time of Descent thro' KV, after the Defcent thro' AK) is to the Time of Descent thro' AV, as the Arch VF, to the Semicircumference; and confequently to the Time of Defcent thro' EV, as the Arch FV, to the Diameter. Wherefore the Time of defcribing the Curve AVB, is to the Time of Descent thro' EV, as the Arch EHVF, to the Diameter EV. But the Time of Defcent thro' EV, is to the Time of Defcent thro' LB or EG, as EV to EF. Therefore (by Equality) the Time of defcribing AVB, is to the Time of Defcent thro' LB, as the Arch EHVF, to the Subtenfe EF, that is, as AM to MB. Again, the Time of Descent thro' LB, is to the Time of Descent thro' AB, as LB to AB. Therefore the Time of defcribing AVB, is to the Time of Descent thro' AB, in the Ratio compounded of AM to BM, and LB to BA, and confequently is equal to the Ratio of AM x LB to MB x BA.

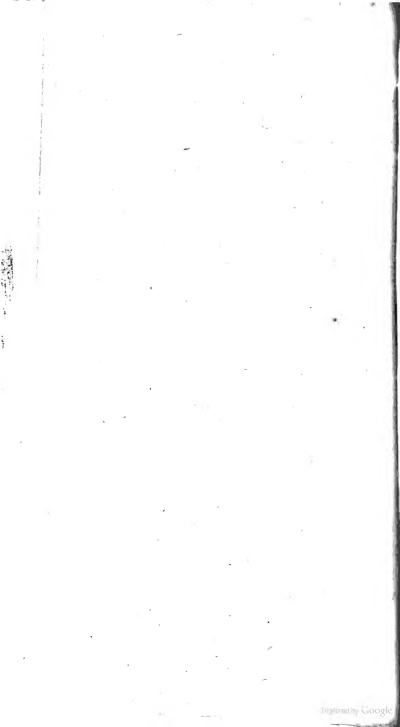
But

But  $AM \times LB = MB \times AC$ ; and therefore the Time in which a Body at reft in A, fhall defcribe the Cycloidal Arch AVB, is to the Time of defcribing the right Line AB, as MB  $\times$  AC to MB  $\times$  BA; that is, as AC to AB. Q: E: D.

And the Demonstration will proceed in like manner, if the point B be between A and V.

4n





An Extract of a Letter from the Reverend Dr. John Wallis, to Richard Waller, Esq; Secretary to the Royal Society, concerning the Spaces in the Cycloid, which are perfectly Quadrable.

Oxford, August 22. 1695.

SIR,

I Find it is thought by most, that there is no other part of the Semicycloid Figure (adjacent to the Curve) that is capable of being Geometrically Squared, but these two, viz.

1. The Segment AbV; (Fig. 28.) taking  $AV = \frac{1}{4}A\alpha$ , (which was first observed by Sir Christopher Wren, and after him by Hugenius and others) and it is  $=\frac{1}{2}sR = \frac{1}{4}R^2$ .

2. The Trilinear A d D (taking d D, in the Parallel d D C, paffing through the Center  $C_{,}$ ) which is  $= R^{2}$ .

But it is otherwife (as I have fhewed in my Treatife, *De Cycloide*, and that, *De Moru*; the Figures of which latter I retain here, fo far as they concern this Occafion;) there being

being other Portions of it, equally capable of Quadrature.

In order to which, I there fhew (De Motu, Cap. 5. Prop. 20. A. p.802,803,804.) that not only the Cycloid is Triple to the Circle Generant, (which was known before) but that the refpettive Parts of that are Triple to those of this: Which is the Foundation on which I build my whole Process concerning the Cycloid in both Treatifes, (and which is not pretended, that I know of, to have been observ'd or known by any Body before me:) That is,  $b \beta \alpha A$  (Fig. 28.) Triple to the Sector  $B \alpha A$ (taking  $b\beta$  parallel to  $B \alpha$ ) where-ever, in the Curve  $A \tau$ , we take the point b.

I then fhew, that the Cycloid is a Figure compounded of thefe two; the Semicircle  $A D \alpha$ , and the Trilinear  $A D \alpha \tau b A$ , lying between the two Curves  $A D \alpha$  and  $A d \tau$ , (and therefore, to Square any part of thefe, is the fame as to Square the refpective part of the Cycloid.

I fnew farther (*Ibidem*, pag. 804.) that this Trilinear is but a difforted Figure (by reafon of the Semicircle thruft in between it and its Axis) which being reftored to its due Pofition (by taking out the Semicircle into a different Figure, (as *Fig.* 29.) and thrufting the Lines *b B* home to the Axis, fo as that BV be the fame point) is the fame with  $A\tau\alpha$ , (*Fig.* 30.) (the Parallelograms  $b\beta\alpha B$ being fet upright, which in the Cycloid ftand floping; and the Circular Arches  $b\beta$ , (*Fig.* 28.) becoming ftreight Lines (in *Fig.* 30.) and the Lines *b B* being, in both, equal to the refpective

dive Arches *B* A, every where;) which therefore I call *Trilineum Reftitutum* (the Trilinear reftored to its due Position, which Figure I do not find that any before me has confider'd:) So that to Square any part of this, is the fame as to Square the respective part of the *Cycloid*, (or of the Trilinear in the *Cycloid*:) That which in the *Cycloid* lies between two Arches of the Circle Generant in different Positions, answering to that which, in the restored Figure, lies between the respective streight Lines.

And therefore A d D A,  $= \tau d \delta \tau$ , (Fig.28.)  $= A d D A = \tau d \delta \tau$ , (Fig. 30.)  $= R^2$ . And A b k d A,  $\tau b k \delta \tau$ , (Fig. 28.) = A b k d A,  $\tau b k \delta \tau$ , (Fig.30.) = s R. And b k d (Fig.28.) = b k d, (Fig. 30.)  $= R^2 - s R$ , Ibid. Cap. 17. B. pag. 756. Where, if b be taken above d k D C, (paffing through the Center C) thefe Figures are within the Cycloid, and within the reftored Figure; but without them, if b be taken below that Line, and adjacent to the Curve  $A b \tau$ , in both Cafes.

By R, I understand the Radius of the Circle Generant; and by s, the Right Sine of the Arch B A, whose versed Sine is V A.

And, where-ever in my whole Difcourfe of the Cycloid, or the reftored Trilinear (which is a Figure of Arches, and a Figure of verfed Sines) the Arch *a* is no Ingredient in the defignation; fuch part or portion of them is capable of being Geometrically fquared. But when I exclude *a*, I do therein exclude *P* (for that is an Arch alfo) and f = a - |-s, and e = a - s, becaufe *a* is therein included.

Mr.

173

Mr. Cafwell, (not being aware that I had fquared thefe Figures) had done the fame by a Method of his own, (which he fhewed me lately) which I would have inferted here, but that he thought it not neceffary; and inftead thereof, hath given me the Quadrature of a Portion of the Epicycloid (which you will receive with this) and, I think, it is purely new.

The

## The Quadrature of a Portion of the Epicycloid.

# By Mr. Cafwell.

175

(Fig. 31.) SUppose DPV to be half of an exterior Epicycloid, VB its Axis, V the Vertex, VLB half of the Generant Circle, E its Center; DB the Base, C its Center: Bifect the Arc of the Semicircle V B in L, and on the Center C through L draw a Circle cutting the Epicycloid in P: Then I fay the Curvilinear Triangle VLP will be = BEqin  $\frac{CE}{CB}$ ; that is, the Square of the Semidiameter of the Generant Circle will be to the Curvilinear Triangle VLF, as CB the Semidiameter of the Bafe, to CE; which CE in the exterior Epicycloid is the Sum of the Semidiameters of the Base and Generant, but in the interior Epicycloid D p u, 'tis the difference of the faid Semidiameters.

#### COROLLARY.

In the interior Epicycloid, if CE is  $\frac{1}{2}CB$ . the Epicycloid then degenerating into a right Line, the Quadrature of the Triangle lup will be in effect the fame with the Quadrature of Hippocrates Chins.

COROL.

176

#### COROL. II.

If the Semidiameter of the Bafe is fuppofed infinite, the Epicycloid then being the common Cycloid, the Area of the faid Triangle will be equal to the Square of the Radius of the Generant, and fo it falls in with that Theorem which Lalovera found, and calls Mirabile.

Though I do not think the abovefaid Quadrature can eafly be deduced from what has been yet publified of the Epicycloid, I have not added the Demonstration; but think it enough to name a general Proposition from whence I deduced it, viz. The Segments of the Generant Circle are to the Correspondent Segments of the Epicycloid, as C B to 2 C E + C B. For Example, fuppofe FmG the Position of part of the Generant when the point F of the exterior Cycloid was defign'd, then the Segment FmGn is to the Segment DFnG: as CB to 2 C E + CB.

And confequently the whole Epicycloid to the whole Generant in the fame Proportion: Which is the only Cafe demonstrated by Moufieur De la Hire.

It follows also that in the Vulgar Cycloid, its Segments are triple of the Correspondent Sectors of the Generant, which was first shewn by Dr. Wallis.

A General Proposition shewing the Dimension of the Areas in all those kinds of Curves which are "describ'd by the Equable Revolution of a Circle upon any Basis, either a Rectilineal or a Circular one.

#### By Edm. Halley.

IS known that the Primary Cycloid, as also the Prolate and the Cycloid, one (which they call Trochoids) have been largely handled by the Celebrated Dr. Wallis and others, and their Properties are now common enough ; fo that there's fcarce any thing new left to be difcover'd concerning them. But the famous M. De Lattire in a late Treatife, having shewn some of the Properties of the Primary Epicycloid, the most Ingenious Mr. Caswell did upon that occasion not only demonstrate that the Mensuration of the whole Epicycloidal Space, obtain'd allo in the parts of the fame, but alfo gave a perfect Quadrature of the Curvilineal Space UPL. But while I was enquiring after the Demonstration of this Quadrature, which is not very obvious, nor as yet given by the In-ventour, I light upon the following general N ProProposition, by the help of which all forts of Curvilineal Spaces, as well of the Cycloidal as Epicycloidal kind, as well the whole Spaces as the parts, are measur'd. And further, not only the Spaces VPL, but also innumerable others, QTP and VQTL, are demonstrated to be capable of an exact Quadrature; and this not only in the Primary Epicycloids, but but also in the Prolate and contracted ones.

The Proposition is as follows.

#### Proposition.

The Area of any Cycloid or Epicycloid, either Primary, Prolate, or Contracted, is to the Area of the Generating Circle, and alfo the Areas of the generated parts in those Curves, are to the Areas of the Analogous Segments of the Circle; as the Sum of twice the Velocity of the Center, and the Velocity of the Circular Motion, to the Velocity of the Circular Motion.

#### Demonstration.

#### (Fig. 32.)

Let YPQRSUB be any Epicycloid defcrib'd by the Revolution of the Circle ULB, upon the Circular Bafis YMNB. Let the Center of the Generating Circle be in c, and drawing cMK, let the Circle ftand upon the Bafis in the point M, and let the defcribing point be S. Now diffinguifhing the Motions, let the point S first of all be carried by the Circular Motion into R, fo that the Arch

SM

179

SM is increased by the indivisible Particle RS. Next fuppofe the Center c to be tranfferr'd to C: by which Motion the Segment RSM being brought into the Polition QTN, the point Q will touch the Curve. 'Tis plain that the Triangle RSM is the Momentum or Fluxion of the Segment of the Circle, and that the Trapezium QSMN is the Fluxion of the Curvilineal Space generated in the fame time. And fince SM, RM, QM, are suppos'd to differ but by a point from one another, let the little Area QSMN be conceiv'd to confift of the three Sectors RMS, RMQ, MQN; and fo the little Area RMS to be to the little Area QSMN, as the Angle RSM to the Sum of the three Angles RMS--RMQ--MQN. But the Angles RMQ-|-MQN, are equal to the Angles MCN -- MKN, or to the Angle cMC ; becaufe of the Lines RM, QN, inclin'd to one another in an Angle equal to MKN, and becaufe of the Angle MQN equal to 1 MCN (by Eucl. 3. 20.) confequently the Angle RMS is to the Angles RMS + cMC, that is (by the fame Proposition mention'd) the Arch  $\frac{1}{2}$  RS to the two Arches Cc  $+\frac{1}{2}$  RS, or RS to 2Cc - RS, as the little Area RSM, to the little Area QSMN, or as the Momentum of the Circular Segment QTN, to the Momentum of the Epicycloidal Segment QSYMN generated in the fame time. And fince these Momenta are ever in that same Ratio, where-ever the point Q be taken, 'tis manifest that the Areas QTN, QSMYN themfelves, generated from thefe Momenta, have alfo the fame conftant Ratio, viz. of N 2 the

Division of Google

the Velocity of the Circular Motion RS, to double the Velocity of the Center, adding the Circular Motion, or 2Cc -1-RS: As alfo the Area UBZ to the Area UBN, and confequently the Semicircle ULB to the Curvilined Space UQYNB. Wherefore the *Propolition* is manifeft.

And there is no other difference in the manner of demonstrating, if the generating Circle moves upon the Concave fide of the Arch, except only that the Angle cMC, in this cafe, is the difference of the Angles MCN, MKN. But if the Basis were a right Line, then MKN vanishing, and RM, QN, being parallel, the Construction will be eafier. I forbear drawing Corollaries from this Proposition, fince they are obvious. But now in all these Curves, the Portions that are Analogous to those Portions which Doetor Wallis has found capable of a perfect Quadrature in the Primary Cycloid, are here alfo equally fquarable; which eafily follows from what has been faid.

Upon the Center K, thro' the point Q, draw the Circular Arch QZ, and draw ZB cutting off the Segment ZLB = the Segment QTN. Then bifect the Semicircle UB in L, and thro' the point L and on the Center K, defcribe the Arch PL cutting the Epicycloid in P, the generating Circle in T, and the Chords QN, ZB, in y and X. Let the Arch VZ = a, its Sine = s, the Radius of the generating Circle = r, the Radius of the Bafe = R, and the Arch CE or the Motion of the Center = m. It is plain that the Sector CKE, is to the Space XyNB, as the Square

Diguered & Google

Square of KE, to the difference of the Squares of KL and KB, or as RR+2Rr+rr, to 2Rr 1- 2rr, that is, as R 1- r, to 2r, or KE to B And confequently the Rectangle BE x CE or rm is equal to the Space XyNB. But the Space VZB is co to the Rectangle  $\frac{1}{2}$  ar  $+\frac{1}{2}$  sr, and fo according to our *Proposition* it will be as a to 2m, fo  $\frac{1}{2}$  ar + 1 sr, to mar + msr equal to the Curvilineal a Space QUZLBNQ. From hence fubftract the Space XyNB = rm, and there remains the Space  $QUZXy = \frac{mrs}{a}$ . And fince the Spaces ZXL, QyT, are equal, the Space QULTQ fhall also be equal to  $\frac{\text{mirs}}{2}$ . Therefore when a to m, or the Circular Motion is to the Progressive Motion of the Center, in a given Ratio, there will be a perfect Quadrature of the Curvilineal Spaces QULTQ. And the whole Space UPL, will be to the Square of the Radius BE, in the fame Ratio (m to a) of the Motions, that is in every Primary Epicycloid, in the Pro-portion of the Radii, KE, KB, which is Mr. Cafwell's Proposition.

But the leffer Spaces QULTQ will be to one another, as the Sines of the Arches UZ; and the Triangular Spaces QTP, by the fame Argument, will be as the verfed Sines of the Arches QT or ZL, and confequently are alfo fquar'd. After the fame manner it will be prov'd, that the Spaces PAT, pLu, par, are ever to the Square of the Radius BE N 3 (in

(in all thefe Figures) in the aforefaid Ratio of m to a; and their Portions pqt, as the verfed Sines of the intercepted Arches qt; but the remaining Segments as qtra, qtra, &c. will be as the right Sines of the Compliments of the fame Arches qt.

But the Ratio of the Velocities, m to a, is compounded of the Ratio of the Radii KB, BE, and the Ratio of the Angles CKE, VEZ, equably defcrib'd together; and iconfequently giving that Ratio of the Angles, all the foremention'd Epicycloidal Spaces will be fquar'd alfo.

I can eafily draw Tangents to all thefe Curves, as alfo I feem to my felf to have gotten their Rectifications, from fome Areas Analogous to them; which may give occafion to a more particular handling of this Family of Curves another time. A Method of Raising an infinite Multinomial to any given Power, or Extracting any given Root of the same.

#### By Mr. A. De Moivre.

"Is about two Years fince, that confidering Mr. Newton's Theorem for Railing a Binomial to any given Power, or Extracting any Root of the fame; I enquir'd, whether what he had done for a Binomial, could not be done for an infinite Multinomial. I foon found the thing was poffible, and effected it, as you may fee in the following Paper; I defign in a little time to fhew the Ufes it may be applied to: In the mean while, thofe that are already vers'd in the Doctrine of Infinite Series, and have feen what Applications Mr. Newton has made of his Theorem, may of themfelves derive feveral Ufes from this.

I fuppofe that the Infinite Number Multinomial is  $az + bzz + cz^3 + dz^4 + ez^5$ , &c. m is the Index of the Power, to which this Multinomial ought to be Rais'd, or if you will, 'tis the Index of the Root which is to be Extracted: I fay that this Power or Root of the Multinomial, is fuch a Series as I have exprest.

N 4

For

For the understanding of it, it is only neceffary to confider all the Terms by which the fame Power of z is multiplied; in order thereto I diffinguish two things in each of these Terms; First, The Product of certain Powers of the Quantities, a, b, c, d, &c. Secondly, The Uncia (as Oughtred calls 'em) prefixt to these Products. To find all the Products belonging to the fame Power of z, to that Product, for instance, whose Index is m-|r (where r may denote any integer Number) I divide these Products into feveral Clasfes; those which immediately after some certain Power of a (by which all these Products begin) have b, I call Products of the first Class; For Example, a^{m-4} b³e is a Product of the first Class, because & immediately follows  $a^{m-4}$ ; those which immediately after some Power of a have c, I call Products of the fecond Cla/fis, fo  $a^{m-3}$  ccd is a Product of the fecond Class; those which immediately after some Power of a have d, I call Products of the third Class, and fo of the reft.

This being done, I multiply all the Products belonging to  $z^{m+r-1}$  (which precedes immediately  $z^{m+r}$ ) by b and divide 'em all by a; Secondly, I multiply by c and divide by a, all the Products belonging to  $z^{m+r-2}$ , except those of the first Class; Thirdly, I multiply by d and divide by a all the Products belonging to  $z^{m+r-3}$ , except those of the first and fecond Class; Fourthly, I multiply by e and divide by a all the Terms belonging

longing to  $z^{m-|-r-4}$ , except those of the first, fecond, and third *Claffus*, and fo on, till I meet twice with the fame Term. Laftly, I add to all these Terms the Product of  $a^{m-1}$ into the Letter whole Exponent is r-|-1.

Here I must take notice that by the Exponent of a Letter, I mean the Number which expresses what Place the Letter has in the Alphabet, fo three is the Exponent of the Letter c because the Letter c is the third in the Alphabet.

It is evident that by this Rule, you may cafily find all the Products belonging to the feveral Powers of z, if you have but the Product belonging to  $z^m$ , viz.  $a^m$ .

To find the Uncia which ought to be prefixt to every Product, I confider the Sum of Units contain'd in the Indices of the Letters which compose it (the Index of a excepted) I write as many Terms of the Series  $m \times m - 1 \times m - 2 \times m - 3$ , &c. as there are Units in the Sum of these Indices, this Series is to be the Numerator of a Fraction, whofe Denominator is the Product of the feveral Scries 1 x 2 x 3 x 4 x 5, &c. 1 x 2 x 3 x 4 x 5, &c. 1x2x3x4x5x6, &c. the first of which contains as many Terms as there are Units in the Index of  $b_1$ , the fecond as many as there are Units in the Index of c, the third as mainy as there are Units in the Index of d, the fourth as many as there are Units in the Index ofe, &c.

De-

2

#### Demonstration.

To raife the Series  $az + bzz + cz^3 + dz^4$ . &c. to any Power whatfoever, write fo many Series equal to it as there are Units in the Index of the Power demanded. Now it is evident that when these Series are fo multiplied, there are feveral Products in which there is the fame Power of z, thus if the Series  $az + bzz + cz^3 + dz^4$ , &c. is rais'd to its Cube, you have the Products b3z6, abcz6, aadz⁶, in which you find the fame Power  $z^6$ Therefore let us confider what is the Condition that can make fome Products to contain the fame Power of z, the first thing that will appear in relation to it, is that in any Product what foever, the Index of z is the Sum of the particular Indices of z in the multiplying Terms (this follows from the Nature of Indices) thus  $b^3 z^6$  is the Product of  $bz^2$ ,  $bz^2$ ,  $bz^2$ , and the Sum of the Indices in the multiplying Terms, is 2 + 2 + 2 = 6;  $abcz^6$ in the Product of az, bzz,  $cz^3$ , and the Sum of them Indices of z in the multiplying Terms is 1 + 2 + 3 = 6 and 26 is the Product of az, az, dz⁴, and the Sum of the Indices of z in the multiplying Terms is 1+1+4=6; the next thing that appears is, that the Index of z in the multiplying Terms is the fame with the Exponent of the Letter to to which z is join'd, from which two Confiderations it follows, that, To have all the Products belonging to a certain Power of z, you must find all the Products where the Sum of the Exponents

186

187

ponents of the Letters which compose 'em shall always be the fame with the Index of that Power. Now this is the Method I use to find easily all the Products belonging to the fame Power of z, Let m - |-r| be the Index of that Power, I confider that the Sum of the Exponents of the Letters which compose these Products must exceed by one those which belong to  $m^{m-1-r-1}$ , now because the Excess of the Exponent of the Letter b above the Exponent of the Letter a, is one, it follows that if each of the Products belonging to  $z^{m+r-1}$  is multiplied by b, and divided by a, you will have Products the Sum of whofe Exponents will be m - [-r; Likewife the Sum of the Exponents of the Letters which compofe the Products belonging to  $z^{m-|-r}$  exceeds by two the Sum of the Exponents of the Letters which compose the Products belonging to  $z^{m+r-2}$ ; Now because the Exponent of the Letter a is less by two than the Exponent of the Letter c, it follows, that if each Product belonging to  $z^{m+r-2}$  is multiplied by c and divided by a, you will have other Products, the Sum of whofe Exponents is still m-1-r; Now if all the Products belonging to  $z^{m+r-2}$  were multiplied by c and divided by a, you would have fome Products that would be the fame as fome of those found before, therefore you must except out of 'em those that I have call'd Products of the first Class; what I have faid shows why all the Products belonging to  $z^{m+r-3}$ , except those of the first and second Cla/fis must be

be multiplied by d and divided by a: Laftly you fee the Reafon why to all thefe Products is added the Product of  $a^{m-1}$  by the Letter whofe Exponent is r - 1 = 1; 'Tis becaufe the Sum of the Exponents is ftill m - 1 = r.

As for what relates to the Uncia; observe that when you multiply az-1-bzz-1-cz³-1-dz⁴, &c. by az 1-bzz 1-cz³-1-dz⁴, &c. each Letter a, b, c, d, &c. of the fecond Series is multiplied by each of the Letters a, b, c, d, &c. of the first Series; Thus the Letter a of the fecond Series is multiplied by the Letter b of the first, and the Letter b of the fecond Series is multiplied by the Letter a of the first; therefore you have the two Plancs. ab, ab or 2ab; for the fame reason you have 2ac, 2ad, &c. Therefore you must prefix to each Plane of those that compose the Square of the infinite Series az -1- bzz -1- cz3, &c. the Number which expresses how many ways the Letters of each Plane may be changed : likewife if you multiply the Product of the two preceeding Series by az-1-bzz-1-cz3, &c. each Letter a, b, c, d, of the third Series is multiplied by each of the Planes form'd by the Product of the first and second Series : Thus the Letter a is multiplied by the Planes be and cb; the Letter b is multiplied by ac and ca; the Letter c is multiplied by ab and ba; therefore you have the fix Solids, abc, acb, bac, bca, cab, cba, or 6abc; Therefore you must prefix to each Solid whereof the Cube of the infinite Series is compos'd, the Number which expresses how many ways the Letters of each Solid may be changed. And gene-

189

generally, You must prefix to any Product whereof any Power of the infinite Series az - bzz cz³, &c. is compos'd the Number which expreffes how many ways the Letters of each Product may be changed.

Now to find how many ways the Letters of any Product, for inftances am-n bb cp dr may be changed; this is the Rule which is commonly given : Write as many Terms of the Series 1x2x3x4x5, &c. as there are Units in the Sum of the Indices, viz. m - n-1-b+p+r, let this Series be the Numerator of a Fraction whofe Denominator shall be the Product of the Series  $1 \times 2 \times 3 \times 4 \times 5$ , &c. 1x2x3x4x5, &c. 1x2x3x4x5x6, &c. 1 x 2 x 3 x 4 x 5, &c. whereof the first is to contain as many Terms, as there are Units in the first Index m - n; the fecond as many as there are Units in the fecond Index h; the third as many as there are Units in the third Index p; the fourth as many as there are Units in the fourth Indexs r. But the Numerator and Denominator of this Fraction have a common Divifor, viz. the Series 1 x 2 x 3 x 4 x 5, &c. continued to fo many Terms as there are Units in the first Index m - n; therefore let both this Numerator and Denominator be divided by this common Divifor, then this new Numerator will begin with m - n + 1, whereas t'other began with 1, and will contain fo many Terms as there are Units in h+p+r, that is, fo many as there are Units in the Sum of all the Indices, excepting the first; as for the new Denominator, it will be the Product of three

three Series only, that is, of fo many as their Indices, excepting the first. But if it happens withal, that " be equal to h+p+r as it always happens in our Theorem, then the Numerator beginning by m - n + 1, and being continued to fo many Terms as there are Units in b - p - r or *n*, the laft Term will be m neceffarily, fo if you invert the Series and make that the first Term which was the last, the Numerator will be  $m \times m - 1 \times m - 2$ x m - 3, &c. continued to fo many Terms as there are Units in the Sum of the Indices of each Product, exceping the first Index. There remains but one thing to demonstrate, which is, that, what I have faid of Powers whofe Index is an Integer, may be adapted to Roots, or Powers whole Index is a Fraction; but it appears at first fight why it should be so: For, the fime Reason which makes me confider Roots under the Notion of Powers, will make me conclude, that whatever is faid of one may be faid of t'other; however, I think fometime to give a more formal Demonstration of it.

A Method of Extracting the Root  
of an Infinite Equation.  
By A. De Moivre, F.R.S.  
THEOREM.  
$$IF az + bzz + cz^{3} + dz^{4} + cz^{5} + fz^{6}, &c. = gy$$
$$+ byy + iy^{3} + ky^{4} + ly^{5} + my^{6}, &c. then will$$
$$z be = \frac{g}{a}y + \frac{b - bAA}{a}y^{2}$$
$$\frac{i - 2bAB - cA^{3}}{a}y^{2}$$
$$\frac{i - 2bAB - cA^{3}}{a}y^{2}$$
$$\frac{k - bBB - 2bAC - 3cAAB - dA^{4}}{a}t^{4}$$
$$+ \frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{bBB - 2bAC - 3cAAB - dA^{4}}{a}t^{4}$$
$$+ \frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{bBB - 2bAC - 3cABB - 3cAAC - 4dA^{3}B}{a}$$
$$+ \frac{a}{a}y^{3}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{a}{a}y^{2}$$
$$\frac{bBD - bCC - 2bAE - cB^{3} - 6cABC}{a}$$
$$- 3cAAD - 6dAABB - 4A^{3}C - 5eA^{4}B}$$
$$- - fA^{6}$$
$$+ \frac{bBB - bCC - 2bAE - 4A^{3}C - 5eA^{4}B}{a}$$

For the understanding of this Series, and in order to continue it as far as we please; it is to be observ'd, 1. That every Capital Letter is equal to the Coefficient of each pre-

preceding Term; thus the Letter B is equal to the Coefficient  $\frac{b-bAA}{a}$ . 2. That the De-

nominator of each Coefficient is always a. 3. That the first Member of each Numerator, is always a Coefficient of the Series  $gy - \frac{1}{y^3}$ , &c. viz. the first Numerator begins with the first Coefficient g, the fecond Numerator with the fecond Coefficient h, and fo on. 4. That in every Member after the first, the Sum of the Exponents of the Capital Letters, is always equal to the Index of the Power to which this Member belongs : Thus confidering the Coefficient

longs to the Power  $y^4$ , we fhall fee that in every Member bBB, 2bAC, 3CAAB, dA⁺, the Sum of the Exponents of the Capital Letters is 4, (where I must take notice, that by the Exponent of a Letter, I mean the Number which expresses what Place it has in the Alphabet; thus 4 is the Exponent of the Letter D) hence I derive this Rule for finding the Capital Letters of all the Members that belong to any Power; Combine the Capital Letters as often as you can make the Sum of their Exponents equal to the Index of the Power to which they belong. -- 5. That the Exponents of the fmall Letters, which are written before the Capitals, express how many Capitals there is in each Member. 6. That the Numerical Figures or Uncie that occur in these. Members, express the Number of Permutations

 $[\]frac{k-bBB-2bAC-3cAAB-dA^{4}}{n}, \text{ which be-}$ 

tions which the Capital Letters of every Member are capable of.

For the Demonstration of this; suppose  $z=Ay-|-Byy-|-Cy^3-|-Dy^4$ , &c. Substitute this Series in the room of z, and the Powers of this Series, in the room of the Powers of z; there will arife a new Series; then take the Coefficients which belong to the feveral Powers of y, in this new Series, and make them equal to the corresponding Coefficients of the Series  $gy-|-byy-|-iy^3$ , &c. and the Coefficients A, B, C, D, &c. will be found such as I have determin'd them.

But if any one defires to be fatisfied, that the Law by which the Coefficients are form'd, will always hold, I'll defire 'em to have recourse to the Theorem I have given for raifing an infinite Series to any Power, or extracting any Root of the fame; for if they make use of it, for taking fucceffively the Powers of Ay-|-Byy-|-Cy3, &c. they will fee that it must of necessity be fo. I might have made the Theorem I give here, much more General than it is; for I might have fuppos'd,  $az^{m}-bz^{m+1}-cz^{m+2} \& c. = gy^{m}+by^{m+1}-j-iy^{m-1}-2$ &c. then all the Powers of the Series Ay-|-Byy-|-Cy3,&c. defign'd by the universal Indices, must have been taken successively; but those who will please to try this, may easily do it, by means of the Theorem for raising an infinite Series to any Power, &c.

This Theorem may be applied to what is called the Reversion of Series, such as finding the Number from its Logarithm given; the Sine from the Arc; the Ordinate of an O Ellipse 194

Ellipfe from an Area given to be cut from any Point in the Axis: But to make a particular Application of it, I'll fuppofe we have this Problem to folve; viz. The Chord of an Arc being given, to find the Chord of another Arc, that fhall be to the first as *n* to 1. Let *y* be the Chord given, *z* the Chord required; now the Arc belonging to the Chord *y* is,  $y + \frac{y^3}{6dd} + \frac{3y^5}{40d^4} + \frac{5y^7}{112d^6}$ &c. and the Arc belonging to the Chord *z* is  $z + \frac{z^3}{6dd} + \frac{3z^5}{40d^4} + \frac{5z^7}{112d^6}$  &c. the first of thefe Arcs is to the fecond as 1 to *n*; therefore multiplying the Extreams and Means together, we fhall have this Equation:

 $z - \left| -\frac{z^{3}}{6dd} - \left| -\frac{3z^{5}}{40d^{4}} + \frac{5z^{7}}{112d^{6}} \right| \&c. = ny + \frac{ny^{3}}{6dd} \\ - \left| -\frac{3ny^{5}}{40d^{4}} + \frac{5ny^{7}}{112d^{6}} \right| \&c.$ 

Compare thefe two Series with the two Series of the Theorem, and you will find  $a=1, b=0, c=\frac{1}{6dd}, d=0, e=\frac{3}{40d^4}, f=0,$ &c.  $g=n, b=0, i=\frac{n}{6dd}, k=0, l=\frac{3n}{40d^4},$ m=0, &c. hence z will be  $=ny - \left| -\frac{n-n^3}{6dd} y^3 \right|$ &c. or  $ny + \frac{1-nn}{2 \times 3dd} yyA$ , &c. Supposing A

to denote the whole preceding Term, which will be the fame Scries as Mr. Newton has . first found.

By the fame Method, this general Problem may be folv'd; the Abfciffe corresponding to a certain Area in any Curve being given, to find the Abiciffe, whose corresponding Area shall be to the first in a given Ratio.

The Logarithmick Series might also be found without borrowing any other Idea, than that Logarithms are the Indices of Powers: Let the Number, whose Logarithm we inquire, be 1-z, fuppose its Log. to be  $az -bzz - cz^3$ , &c. Let there be another Number 1 - y; thereof its Logarithm will be  $ay - byy - cy^3$ , &c. Now if 1 - z =

 $\overline{1-y|^n}$ , it follows, that  $az + bzz + cz^3$ , &c.  $ay + byy + cy^3$ , &c. :: n, 1. that is,  $az + bzz + cz^3$ , &c.  $= nay + nbyy - ncy^3$ , &c. Therefore we may find a Value of z express by the Powers of y; again, fince 1-z =

 $1 - |-y|^n$ , therefore  $z = 1 - |-y|^n - 1$ , that is,  $z = ny - |-\frac{n}{1} \times \frac{n-1}{2} yy - |-\frac{n}{1} \times \frac{n-1}{2} \times \frac{n-2}{3}$   $y^3$ , &c. Therefore z is doubly express by the Powers of y. Compare these two Values together, and the Coefficients a, b, c, &c. will be determin'd, except the first a which may be taken at pleasure, and gives accordingly, all the different Species of Logarithms.

0 2

An

An Experiment of the Refraction of the Air made at the Command of the Royal Society, March 28. 1699.

By J. Lowthorp, A. M.

W E took a Cylinder of Caft-Brass (Fig. 33.) ABCD, and cut one end of it CD perpendicular to the Axis ax, the other end AB enclin'd to it at an Angle of about 27°. 30'. and therefore the Perpendicular to this enclining plain, pc, and the Axis of the Cylinder ax comprehended an Angle pca of about 620. 30d. These ends were groun'd very true upon a Glafs-Grinder's Brafs-Tool, and each of them was compast about with a narrow Ferule of thin Brass bbbb. Into the upper fide of the Cylinder at E was folder'd the Brafs-Pipe EF, and into the under fide at 6 the other Brass-Pipe GH; the former of these Pipes being about 3 Inches long, and the latter 6 Inches. Upon the Plate ddd were fixt to two other Plates LL perpendicular to it and parallel to each other. Each of these two Plates had an Arch of a Circle (equal to the Circumference of the Cylinder) cut out of its upper Edge, fo that when the Pipe GH was let through a hole near the middle of the Plate ddd, the Cylinder fell into the Arches; and being fasten'd there with Soder, the Axis ax laid parallel to the Plate ddd and about an Inch and half

Dig and Google

half above it. The perpendicular End of the Cylinder DC was clos'd with an Object Glass of a feventy fixth Foot Telescope oo; and the other end AB, with a well polifh'd flat Glafs ff; which was carefully chosen to transmit the Object diftinct enough notwithfanding its Obliquity to the Vifual Rays. The Ferules were well fill'd with Cement round about the Edges of the Glafs, and they laid flat and every where touch'd the finooth Ends of the Cylinder, that they might firmly refift the preffure of the excluded Air.

Instead of a Cistern (as in the Torricellian Experiment) we made use of the Inverted Siphon of Brafs (Fig. 34.) MNO, foder'd to the Plate ggg. One of the fides MN flood perpendicular to the Plate, and the other fide NO enclin'd to it, and was fupported near the upper end O with a little prop kk.

We then plac'd the Cylinder (as in Fig.35.) upon a Table which was well fasten'd to a firm Floor; the Pipe GH was let through a hole, and the Axis laid almost parallel to the fides of the Table, and the Plate ddd was nail'd down to it. The Tube of the Telefcope //, with the Eye-glafs, was apply'd to the Object Glafs, and a Hair fix'd within it at the common Focus of both Glasses in the Axis of the Cylinder continu'd, x. Upon the Floor (under the Cylinder) we nail'd the Plate ggg with the Inverted Siphon upon it, and join'd M to H by the Infection of the Glafs Tube T. The Joints were very carefully clos'd with Cement: And then they were cover'd over with pieces of a Bladder and wrapt hard with ftrong Thread. There was

03

was also a Bladder ty'd below each Joint at m, and when it was fill'd with Water it was ty'd above it at n; fo that no Air could come to the Cement, or infinuate it felf through its Pores or Siffures if any happen'd to be left unclos'd.

It is not (I think) an unneceffary trouble, that in this account of the Apparatus I have mention'd fo many minute Circumftances, for we found it difficult enough to exclude the Air, and almost impossible to difcover the very little holes through which fo fubtil a Fluid would freely enter and posses the Spaces deferted by the fubliding Mercury. But with all this Precaution the Experiment fucceeded at last, as I wish'd, after this manner.

We plac'd the Object a (which was a black Thread fliding in a little Frame over a piece of white Paper) in the Axis of the Cylinder ex continu'd to it, we filled the Pipes and Cylinder with Mercury; and having ftopt the uppermost Pipe at F with the little Iron-stopple K and clos'd it at the other Joints, we let the Mercury run out gently at O into the Bladder v, till it remain'd fospended at the usual height (as in the Barometre) leaving the fpace above it between the Glasses oo and ff void of Air. We then found the Object, which before appear'd in the Axis at x, rais'd confiderably above it; and we reduc'd it to appear at x by removing it from a to x. The Axis therefore, of the vifual Ray x4, (which was also the Axis of the Cylinder) xa, falling perpendicularly on the void space in the Cylinder past through it without any Refraction :

199

fraction: But emerging obliquely into the Air, it was refracted towards the Perpendicular pc, and there receiv'd a new Direction to x. And therefore the fpace ax fubftended the Angle of Refraction acx; which we meafur'd and found as follows.

The height of the Object Inches Depths above the Axis of vi- fual-Ray ax the unre- fracted
The Diftance of the Ob- ject from the Refract- ing Plain, &c. about 51 Feet or
Therefore the Angle of $\left\{ \begin{array}{ccc} \circ & 2 \end{array} \right\} \circ \left\{ \begin{array}{ccc} \circ & 2 \end{array} \right\}$ The Angle of Emerfion-
pca (by the Conftructi- on of the Cylinder) 62. 30. was
Therefore the Angle of $1$ Incidence $pcx = (=pca) = 62$ . 27. 37. $-1 = acx^2 = -2$
And therefore univerfally (according to the known Laws of Refraction)

The Sines of the Angles of Incidence being — The Sines of the Angles Emerfion are — And the Refractive Power of the Denfe Air — 36

By the Refractive Power of a pellucid Body, I mean that property in it whereby the Oblique Rays of Light are diverted from their direct Courfe; and which is meafur'd O 4 by

Dig under Google

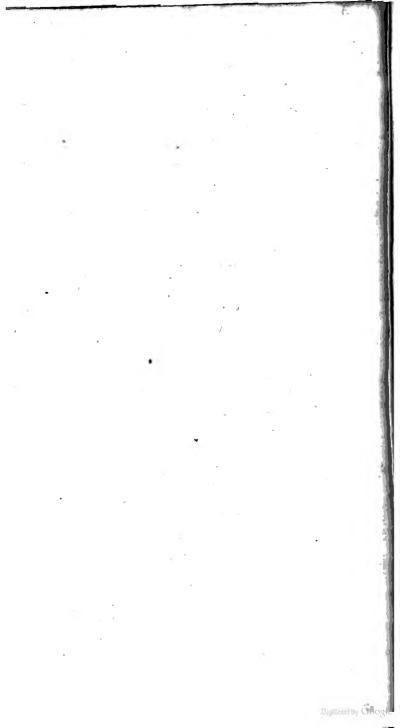
by the proportional Differences always, obferv'd between the Sines of the Angles of Incidence and Emerion.

This Property is not always proportional to the Density (at least not to the Gravity) of the Refracting Medium. For the Refractive Power of Glass to that of Water, is as 55 to 34, whereas its Gravity, is as 87 to 34; that is, the Squares of their Refractive Powers are (very near) as their respective Gravities. And there are fome Fluids which though lighter than Water yet have a greater power of Refraction; thus the Refractive Power of Spirit of Wine (according to Dr. Hook's Experiments, Microg. p. 220.) is to that of Water, as 36 to 33, and its Gravity reciprocally, as 33 to 36, or 361. But the Refractive Powers of Air and Water feem to obferve the fimple Proportion of their Gravities, directly; as I have compar'd them in the following Table. The Numbers there expressing the Refraction of Water are taken from the Mean of * Nine Obfervations at fo many feveral Angles of Incidence, made Jan. 25. 1647. by Mr. Gascoigne the ingeniuos First Inventor of the Micrometer, and the ways of measuring Angles by Telescopes, and those of Air are produc'd by the Experiment above related.

* I am indebted for them to Mr. Flamsteed, who had cover'd them with his Observations, and several Passages relating to them, from his Letters to Mr. Crabtree, which were happily preserv'd in the Time of our Civil War by Sir Jonas Moor, and Mr. Christopher Towneley; and are now in the Hands of Mr. Richard Towneley of Towneley in Lancashire, by whom they were imparted to bim. The

200

19 22 Dislaged by Google



From hence it feems very probable that their Respective Densities and Refractive Powers are in a just Simple Proportion : And if this should be confirm'd by fucceeding Experiments, made at different Angles of Incidence, and with Cylinders continuing exhaufted through feveral Changes of the Air, it would be more than probable that the Refractive Powers of the Atmosphere are every where, at all heights above the Earth, in proportion to its Densities and Expansions. And here it would be no difficult matter to trace the Light through it, thereby to terminate the Shadow of the Earth; and (together with proper Expedients for measuring the Quantity of Light illuminating an Opaque Body) to examine at what diffances the Moon must be from the Earth to fuffer Eclipfes of the obferv'd Duration. This Limitation is confiderable enough in Aftronomy, abundantly to recompense the Trouble of profecuting fuch a new Experiment.

1

----

A Discourse concerning a Method of Discovering the true Moment of the Sun's Ingress into the Tropical Sines.

# By E. Halley.

T may perhaps pass for a Paradox, if not feem extravagant, if I should affert that it is an eafier matter to be affur'd of the Moments of the Tropicks, or of the Times of the Sun's Entrance into Cancer and Capricorn, than it is to observe the true Times of the Equinoctials or Ingress into Aries and Libra. I know the Opinion both of Ancient and Modern Aftronomers to the contrary; Ptolemy fays exprelly, Ta's Two Teonwo Inphoess Suodianeirus Evan; And Ricciolus begins his Chapter of the Solftitial Obfervations with thefe words, Merito Snellius, in notis ad obfervationes Hassiacas, pronunciavit, Herculei esfe laboris vitare in Solftitiis observandis errorem quadrantis diei, and this because of the exceeding flownefs of the change of the Sun's Declination on the day of the Tropick, being not a quarter of a Minute in twenty four Hours. This indeed would make it very difficult, nor would any Inftruments fuffice to do it, were the Moment of the Tropick to be determin'd from one fingle Obfervation. But by three fubsequent Observations made

202

made near the Tropick, at proper Intervals of Time, I hereby defign to fhew a Method to find the Moment of the Tropicks capable of all the Exactness the most Accurate can defire; and that without any confideration of the Parallax of the Sun, of the Refractions of the Air, of the greatest Obliquity of the Ecliptick, or Latitude of the Place: All which are requir'd to ascertain the Times of the Equinoctials from Observation, and which being faultily affum'd, have occafion'd an Error of near three Hours in the Times of the Equinoctials deduced from the Tables of the Noble Tycho Brahe and Kepler, the Vernal being fo much later, and the Autumnial fo much earlier than by the Calculus of those Famous Authors.

Now before we proceed, it will be neceffary to premise the following Lemmata, ferving to demonstrate this Method, viz.

1. That the Motion of the Sun in the Ecliptick, about the Time of the Tropicks, is fo nearly equable, that the difference from Equality is not fenfible, from five days before the Tropick, to five days after: And the difference arifing from the little Inequality that there is, never amounts to above  $\frac{1}{4}$  of a fingle Second in the Declination, and this by reafon of the nearnefs of the Apogaon of the Sun to the Tropick of Cancer.

2. That for five Degrees before and after the Tropicks, the differences whereby the Sun falls flort of the Tropicks, are as the verfed Sines of the Sun's diffance in Longitude from the Tropicks, which verfed Sines in Arches under five Degrees, are beyond the utmoft nicety nicety of Senfe, as the Squares of those Arches. From these two follow a third :

3. That for five days before and after the Tropicks, the Declination of the Sun falls fhort of the utmost Tropical Declination, by Spaces which are in duplicate Proportion, or as the Squares of the Times by which the Sun is wanting of or past the Moment of the Tropick.

Hence it is evident that if the Shadows of the Sun, either in the Meridian or any other Azimuth, be carefully observ'd about the Time of the Tropicks, the Spaces whereby the Tropical shade falls short of, or exceeds those at other Times, are always proportionable to the Squares of the Intervals of Time between those Observations and the true Time of the Tropick, and confequently. if the Line, on which the Limits of the shade is taken, be made the Axis, and the correfoondent Times from the Tropick expounded by Lines, be erected on their refpective Points in the Axis as Ordinates, the Extremities of those Lines shall touch the Curve of a Parabola; as may be feen in the Figure: Where a, b, c, e, being fuppofed Points observed, the Lines a B, b C, c A, eF, are respectively proportional to the Times of each Obfervation before or after the Tropical Moment in Cancer.

This premifed, we fhall be able to bring the Problem of finding the true Time of the Tropick by three Obfervations, to this Geometrical one; having three Points in a Parabola A, B, C, or A, F, C given, together with the direction of the Axis, to find the Di-

Digitized by Google

205

Diftance of those Points from the Axis. Of this there are two Cafes, the one when the Time of the fecond Observation B is precifely in the middle Time between A and C: In this Cafe putting t for the whole Time between A and C, we shall have A c the Interval of the remotest Observation A from the Tropick by the following Analogy.

As 2 a c—bc to 2 a c— $\frac{1}{2} b c$ :: So is  $\frac{1}{2} t$ or AE: to Ac the Time of the remotest Observation A from the Tropick.

But the other Cafe when the middle Obfervation is not exactly in the middle between the other two Times, as at F, is fomething more operofe, and the whole Time from A to C being put =t, and from A to F = s, c e = c, and b c = b, the Theorem t t c - b s swill ftand thus - - = A c at the Time

fought.

To illustrate this Method of Calculation it may perhaps be requisite to give an Example or two for the fake of those Astronomers that are less instructed in the Geometrical part of their Art.

2 t c-2 b s

Anno 1500. Bernard Walther, in the Month of June, at Nuremburg, observ'd the Chord of the distance of the Sun from the Zenith by a large Parallactick Instrument of Ptolemy, as follows:

June	2.	45467.	June ?	8.	44975.
		44934.			44883.
June	16.	44990.	June	16.	44990.

In

Dhi and by Google

In both which Cafes the middle Time is exactly in the middle between the Extreams. and therefore in the former three, a c = 533, b = 477 and t, the Time between being 14 days, by the first Rule, the Time of the Tropick will be found by this Proportion, as  $(89 to 827 \frac{1}{2}:: So \frac{1}{2} t \text{ or } 7 \text{ days to } 9 \text{ days})$ 20h. 2'. whence the Tropick, Anno 1500. is concluded to have fallen June 11d. 20h. 2'. In the latter three, a c is = 107, and bc = 15, and the whole Interval of Time is 8 days = to t; whence as 199 : to  $206\frac{1}{2}$ :: fo is 4 days to 4d. 3h. 37'. which taken from the 16th day at Noon, leaves 11d. 20h. 23'. for the Time of the Tropick, agreeing with the former to the third part of an hour.

Again, Anno 1636. Gallendus at Marfeilles, observ'd the Summer Solftice by a Gnomon of 55 Foot high, in order to determine the Proportion of the Gnomon to the Solftitial fhade, and he hath left us thefe Observations, which may ferve as an Example for the fecond Rule.

June 19. St. N. shadow 31766 parts, whereof the Gnomon was 89.128.

June 20.	31753	
June 21.	. 31751	
June 22.	31759	

These being divided into two Sets of three Observations each, viz. the 19th. 20th. and 22th. and the 19th. 21th. and 22th. we fhall have in the first three c = 13 and b = 7, t = 3 days, s = 1, and in the fecond c = 15and

Initized by Google

207

and b = 7, t = 3 and s = 2. Whence, according to the Rule, the 19th day at Noon the Sun wanted of the Tropick a Time proportionate to one day, as ttc - ssb to 2tc - 2bs, that is, as 110 to 64 in the first Set, or 107 to 62 in the fecond Set; that is,  $1d \cdot 17^{h} \cdot 15'$ . in the first, or  $1d \cdot 17^{h} \cdot 25'$ . in the fecond Set: So that we may conclude the Moment of the Tropick to have been *June*  $10d \cdot 17^{h} \cdot 20'$ . in the Meridian of *Marfelles*.

Now that these two Tropical Times thus obtain'd, will be found to confirm each others Exactness from their near Agreement, appears by the Interval of Time between them; viz. 1^{d.} 2^{h.} 30'. less than 136 Julian Years: whereof 1^{d.} 1^{h.} 8'. arises from the defect of the length of the Tropical Year from the Julian, and the rest from the Progression of the Sun's Apogeon in that Time; so that no two Observations made by the same Obferver in the same Place, can better answer each other, and that without any the least Artifice or Force in the management of them.

What were the Methods ufed by the Ancients to conclude the hour of the Tropicks, *Ptolemy* has no where delivered; but it were to have been wifhed that they had been aware of this, that fo we might have been more certain of the Moments of the Tropicks we have receiv'd from them, which would have been of fingular ufe to determine the Queftion, Whether the Sun's Apogaon be fixt in the Starry Heaven; or if it move, What is the true Motion thereof? It is certain, that if we take the Account of

of Ptolemy, the Tropick faid to be obferv'd by Euctemon and Meton, Junii 27. mane, Anno 432. ante Christum, can no ways be re-concil'd without supposing the Observation made the next day, or June 28th in the Morning. And Ptolemy's own Tropick obferv'd in the third Year of Antoninus, Anno Christi 140. was certainly on the 23th and not the 24th day of June; as will appear to those that shall duly confider and compare them with the length of the Year deduced from the diligent and concordant Obfervations of those two great Astronomical Genii, Hipparchus and Albatani; establish'd and confirm'd by the Concurrence of all the Modern Accuracy. For these Observations give the length of the Tropical Year, fuch as to anticipate the Julian Account only one day in 300 Years; but we are now fecure that the faid Period of the Sun's Revolution does anticipate very nearly three days in 400 Years; fo that the Tables of Ptolemy founded on that Supposition, do err about a whole day in the Sun's Place, for every 240 Years. Which principal Error in fo Fundamental a Point, does vitiate the whole Superstructure of the Almagest, and serves to convict its Author of want of Diligence, or Fidelity, or both.

But to return to our Method, the great Advantage we have hereby, is, that any very high Building ferves for an Inftrument, or the Top of any high Tower or Steeple, or even any high Wall whatfoever, that may be fufficient to intercept the Sun, and caft a true fhade: Nor is the Position of the Plane

209

on which you take the fhade, or that of the Line therein, on which you measure the Recefs of the Sun from the Tropick, very material; but in what way foever you difcover it, the faid Recefs will be always in the fame Proportion, by reafon of the finalnefs of the Angle, which is not fix Minutes in the first five days: Nor need you enquire the height or diftance of your Building, provided it be very great, fo as to make the Spaces you measure large and fair. But it is convenient that the Plane on which you take the shade be not far from Perpendicular to the Sun, at least not very Oblique, and that the Wall which cafts the fhade, be ftraight and fmooth at Top, and its Direction nearly East and Weft, for Reafons that will be well underftood by a Reader skilful in the Doctrine of the Sphere. And it will be requisite to take the Extream greatest or least Deviation of the shadow of the Wall, because the shade continues for a good Time at a ftand, without alteration, which will give the Obferver leifure to be affur'd of what he does, and not be furpriz'd by the quick transient Motion of the fhade of a fingle Point at fuch a diftance. The principal Objection is, that the Penumbra or Partile shade of the Sun, is in its Extreams very difficult to diffinguish from the true shade, which will render this Observation hard to determine nicely. But if the Sun be transmitted through a Telescope, after the manner us'd to take his Species in a Solar Eclipse, and the upper half of the Objectglafs be cut off by a Paper pasted thereon, and the exact upper Limb of the Sun be feen P juft

just Emerging out of, or rather continging the Species of the Wall, (the Polition of the Telescope being regulated by a fine Hair extended in the Focus of the Eye-glass) I am affur'd that the Limit of the shade may be obtain'd to the utmost Exactness: And of this I defign to give a Specimen by an Obfervation to be made in June next, by the help of the high Wall of St. Paul's Church, London, of which fome following Transaction may give an Account. In the mean time what I have premis'd may fuffice to fet others at work, where fuch or higher Buildings are to be met with. I shall only Advertise, that the Winter-Tropick by this Method may be more certainly obtain'd than the Summer's, by reason that the fame Gnomon does afford much larger Radius for this manner of a Obfervation-

Thilzed by Google

# A Scale of the Degrees of H E A T.

# The Signs and Descriptions of the several Degrees of Heat.

THE Warmth of the Winter Air when Water begins to freeze. This is known accurately by placing a Thermometer in Snow prefs'd clofe together at the Time of a Thaw.

D. I. 2. 2. 3. 4. The Warmths of the Winter Air. The Warmths of the Air in Spring and Autumn.

4.5.6. The Heat of the Air in Summer. 6. The Heat of the Air at Noon in the Month of July.

12. I. The greatest Heat that a Thermometer acquires, by the contact of a Humane Body; which is much the fame with that of a Bird brooding upon its Eggs.

14.¹.1.¹. The nearly greateft Heat of a Bath, that a Perfon holding his Hand fleady and immoveable in the fame, can endure for fome time.

17.1 $\frac{1}{2}$ . The greatest Heat of a Bath, that a Person holding his Hand steady and immoveable in the P 2 fame,

- 201².1³. The Heat of a Bath by which melted Wax fwimming upon it, begins to grow ftiff, and lofe its Transparency.
  - 24. 2. The Heat of a Bath by which Wax fwimming upon it, is melted and preferv'd in a State of Fluidity, without Ebullution.
  - $28\frac{4}{11}\cdot 2\frac{1}{4}$ . The middle Degree of Heat, between that by which Wax is melted, and that which makes Water boil.
    - 34.2¹/₂. The Heat by which Water is made to boil vehemently; and a Mixture of 2 parts of *Lead*, 3 of *Tin*, and 5 of *Bifmuth*, cooling, begins to harden.

Water begins to boil with a Heat of 33 parts, and by boiling, hardly conceives a greater Heat than that of 34 parts.

But Iron growing cool, when it has a Heat of 35 or 36 parts, ceafes to make any Ebullition when warm Water falls drop by drop upon it; as it does also with a Heat of 37 parts, when cold Water falls on it in the like manner.

40⁴¹. 2³/₄. The leaft Heat, by which a Mixture of 1 part of *Lead*, 4 of *Tin*, and 5 of *Bifmuth* is liquefied and preferv'd in a State of Fluidity.

> 3. The leaft Heat, by which a Mixture of equal parts of *Tin* and *Bifmuth* is liquefied. This Mixture grow-

48.

212

Distant by Googl

213

growing cool, when it has a Heat of 47 parts, is coagulated.

- 57.3¹/₄. The Heat, by which a Mixture of 2 parts of *Tin* and 1 of *Bifmuth* is liquefied; as alfo of 3 parts of *Tin* and 2 of *Lead*. But a Mixture of 5 parts of *Tin* and 2 of *Bifmuth* (cooling) does with this Degree of Heat become hard: And the fame come to pafs in a Mixture of equal parts of *Lead* and *Bifmuth*.
- $68.3\frac{1}{2}$ . The leaft Heat by which a Mixture of 1 part of *Bifmuth* and 8 of *Tin* is liquefied. *Tin* by it felf is fus'd with a Heat of 72 parts, and growing cold, hardens with a Heat of 70 parts.
- 81.3¹/₄. The Heat by which *Bifmuth* is fus'd, as alfo a Mixture of 4 parts of *Lead* and 1 of *Tin*. But a Mixture of 5 parts of *Lead* and 1 of *Tin* when fus'd, and growing cold, hardens with this Degree of Heat.
  - The leaft Heat by which Lead is melted. Lead melts with a Heat of 96 or 97 parts, and growing cold, hardens with a Heat of 95 parts.

114.44

The Heat with which Fiery Bodies (growing cool) wholly ceafe fhining in the Night; as alfo, that Heat with which (growing warm) they first begin to fhine in the the Darknefs of the Night, but with a faint and feeble Light, fuch as can fcarce be difcern'd. P 3 This

This Heat liquefies a Mixture of equal parts of *Tin* and *Regulus Martis*; and a Mixture of 7 parts of *Bi/muth* and 4 of the fame *Regulus* (growing cool) hardens with the fame Degree of Heat.

- 136.4 $\frac{1}{2}$ . The Heat-by which Fiery Bodies do in the dark Night appear bright and fhining, but not in the Twilight. A Mixture of 2 parts of *Regulus Martis* and 1 of *Bifmub*, as alfo of 5 parts of *Regulus Martis* and 1 of *Tin*, growing cool, will at this Degree of Heat become hard. The *Regulus* by it felf, hardens with a Heat of 146 parts.
- 161.44. The Heat by which Fiery Bodies, in the Twilight, a little before the Sun's rifing or after his fetting, do fhine difcernably; but not at all in the clear Day-light, or at leaft very obfcurely.
- 5. The Heat of a small Culinary 192. Fire made of Sea-Coal, burning freely by it felf without the help The fame is the Heat Bellows. of Iron, as Red-hot as it can be made in fuch a Fire. The Heat of fmall Culinary Fire made of a Wood, is fome little matter greater, viz. about 200 or 210 parts. And the Heat of a large Fire is Itill greater, especially if it be blown up by the Bellows.

In

Google Google

In the first Column of this Table are the feveral Degrees of Heat, going on in an Arithmetical Progression, beginning with that Degree of Heat, which there is in the Air in Frosty Weather, when Water makes the first Advances towards Freezing; beginning the Account from this, as the lowest Degree of Heat, or common Terminus of Heat and Cold) and supposing the external Heat of a Humane Body to be rated at 12 parts. In the fecond Column are the Degrees of Heat in a Geometrical Proportion, fo that the fecond Degree is double the first, the third double the fecond, and fo on; the first Degree being that external Heat of a Humane Body, proportion'd to the Senfe. But now 'tis manifest from this Table that the Heat of Boiling Water is almost 3 times greater than that of a Humane Body; and that the Heat of melted Tin is 6 times, of melted Lead 8 times, of melted Regulus 12 times, and of ordinary Culinary Fire 16 or 17 times greater than the foremention'd Heat of a Humane Body.

This Table was made by the help of a Thermometer and Red-hot Iron. By the Thermometer I found the Meafure of all the Degrees of Heat as far as that by which Tin is melted; and by the hot Iron I found the Meafure of the reft. For the Heat which hot Iron does communicate to cold Bodies contiguous to it in a given time, (that is the Heat which the Iron it felf lofes) is as the whole Heat of the Iron. And therefore if the Times of Refrigeration are P 4 taken 216

taken equal, the Degrees of Heat shall be in a Geometrical Proportion, and confequently may cally be found by a Table of Logarithms. First of all therefore I found by a Thermometer made of Linfeed Oil, that if when the Instrument was placed in melting Snow, the Oil occupied a Space of 10000 parts, the fame Oil rarified by a Heat of the first Degree (that is by that of a Humane Body) would extend to 10256 parts; and by the Heat of Water beginning to boil, to 10705 parts; and by the Heat of Water boiling vehemently, to 10725 parts; and by the Heat of melted Tin (cooling, and beginning to be of the Confiltence of an Amalgama) to 11516 parts; and by the Heat of the fame Tin when 'tis quite harden'd, to 11496 parts. Therefore the Oil was rarified in the proportion of 40 to 39, by the Heat of a Humane Body; and in the proportion of 15 to 14, by the Heat of boiling Water; and in the proportion of 15 to 13, by the Heat of the melted Tin, beginning to come to the Confiftence of an Amalgama; and in the proportion of 23 to 20, by the Heat of the fame Tin quite hardned.

The Rarefaction of Air with an equal Degree of Heat, was 10 times greater than that of Oil; and the Rarefaction of Oil nearly 15 times greater than that of Spirit of Wine. Now these things thus found, supposing the Degrees of Heat in the Oil to be proportional to its Rarefaction, and the Heat of a Humane Body to be 12 parts; from hence the Heat of Water when it begins to boil, comes

217

comes to be 33 parts, and when it boils venemently, 34 parts; and the Heat of melted Tin beginning to come to the Confiftence of an Amalgama, 72 parts; and the Heat of the fame, when in cooling 'tis come to downright Hardnefs, 70 parts. And having determin'd these Things, in order to find out the reft, I heated a piece of Iron 'till it was Red hot enough, and taking it out of the Fire with a pair of Tongs that were alfo Red-hot, and I it in a cool place, where the Wind blew conftantly. Then putting upon it little pieces of Metals and various other liquable Bodies, I observ'd the times of Refrigeration, 'till all those melted parts having quite loft their Fluidity, became hal'd and folid again, and the Heat of the Iron was equal to that of a Humane Body. Then supposing the Excesses, of the Heats of the Iron and the liquefied Particles approaching to Induration, above the Heat of the Atmosphere founded by the Thermometer, to be in a Geometrick Progreffion, when the Times are in an Arithmetick one; by this means all the Degrees of Heat were discover'd. But 'tis to be obferv'd that I plac'd the Iron not in a ferene and quiet Air, but in a Wind blowing uniformly, fo that the Air which was heated by the Iron might always be carried away by the Wind, and a cold Air with an uniform Motion might fucceed in the place of it. For thus, equal parts of the Air were heated in equal times, and acquired a Heat proportional to that of the Iron. But the Degrees

Degrees of Heat found by this Method had the fame Proportion among themfelves, that those had which were found by the Thermometer; and therefore the Assumption, that the Rarefactions of the Oil were proportional to the Degrees of Heat, was a just and true one.

The

Dation by Google

# The Properties of the Catenaria.

By David Gregory, M. D. Savilian Professor of Astronomy, and F. R. S.

## PROP. I. PROBLEM.

To find the Relation of the Fluxion of the Axis, to the Fluxion of the Ordidinate in the Catenaria.

**L** E T FAD be a Catena hanging on the Extremites F and D, the loweft point of which (or the Vertex of the Curve) is A, the Axis AD perpendicular to the Horizon, and the Ordinate BD parallel to the fame! We are to find the Relation between Bb or  $D\mathcal{A}$ , and  $d\mathcal{A}$ ; fuppofing the point b infinitely near to B, and bd parallel to BD, as alfo  $D\mathcal{A}$  to BA.

From the Principles of Mechanicks, 'tis plain that three Powers which are in Equilibrio, are in proportion to one another, as three right Lines parallel to their respective Directions (or inclin'd in any given Angle to them) and terminated at their mutual Intersection.

And

220

## Miscellanea Curiosa.

And confequently if Dd expounds the abfolute Gravity of the Particle Dd (as it will be in a Catena equally thick) then  $d\mathcal{A}$  will reprefent that part of the Gravity which acts perpendicularly upon Dd, and by which it comes to pass that dD (being by the flexibility of the Chain moveable about d) endeavours to bring it felf into a Vertical Position. And therefore if  $\mathcal{A}d$  (or the Fluxion of the Ordinate BD) be Constant, the Action of the Gravity exerted perpendicularly upon the correspondent parts of the Catena Dd, will also be constant, or every where the fame. Let this Action or Force be expounded by a.

Farther; From the above cited Proposition in Mechanicks, Ds or the Fluxion of the Axis AB, will expound the Force to be exerted in the direction dD, which is equivalent to the former Endeavour of Dd (by which it tends to bring it felf into a Vertical Position) and is fufficient to hinder it.

But this force arifes from the Linea Gravis DA pulling with the direction dD, and is confequently (all the reft continuing as before) proportional to that Line DA. Therefore  $\mathcal{A}d$ , the Fluxion of the Ordinate, is to  $\mathcal{A}D$ , the Fluxion of the Absciffe, as the constant right Line *a*, to the Curve DA. Q: E: F.

#### COROL.

If the right Line DT touches the Catenaria, and meets the Axis AB produc'd in T, then will DB : BT ::  $(d^{\mathcal{I}} : \mathcal{A}D : :)$  a : DA Curve.

PROP.

### PROP.II. THEOREM.

(Fig.34.) If upon the Perpendicular AB as an Axis, and the Vertex A, an Equilateral Hyperbola AH be describ'd, whose Semiaxis AC = a; as also upon the same Axis and Vertex, a Parabola AP whose Parameter is quadruple the Axis of the Hyperbola, and the Ordinate of the Hyperbola HB be always produc'd till HF be equal to the Curve AP: I say then, that (making BD and BF, equal) the Curve FAD, in which the Points F, D, are posited, is the Catenaria.

Put AB = x; then Bb = x, and  $BH = \sqrt{2ax - xx}$ ; whence (from the Method of Fluxions) the Fluxion of BH, that is mb = x

ax+xxAgain, fince the Parabola AP  $\sqrt{2ax+xx}$ has for its Parameter 8a, BP fhall =  $\sqrt{8ax}$ . Whence the Fluxion of BP, that is np =

 $\frac{2ax}{\sqrt{2ax}}$ . Wherefore the Fluxion of the Curve  $\sqrt{2ax}$ .

AP

222 Miscellanea Curiosa. AP (= Pp =  $\sqrt{vp^{q} - Pn^{q}}$ ) =  $\sqrt{\frac{4a^{2}x^{2}}{a^{2}x^{2}}}$ 2ax-1-xx  $=\sqrt{2ax^2+xx^2}$ ; which is equal to -Vzax-l-xx x as appears by multiplying both Numerator and Denominator into  $\sqrt{2a+x}$ . And fince HF is every where = AP, the Fluxion of HF that is mb - sf, fhall = But ax--xx we have hitherto found mb = Vzax+xx Therefore sf (the Fluxion of BF the Ordinate in the Catenaria) = -----: and 2ax-1-xx confequently the Fluxion of the Curve AF (that is,  $Ff = \sqrt{sf^{9} - |-Fs^{9}|} = \sqrt{\frac{a^{2}x^{2}}{2ar - |-rr|}} + \dot{x}^{2}$ ) is =  $\frac{ax + xx}{\sqrt{2ax + xx}}$ , the Flowing Quantity of

which

Digitized by Google

which was fhewn but now to be  $\sqrt{2ax - xx}$ . And therefore  $AF = \sqrt{2ax - xx}$ . And 'tis plain, that the Fluxion of the Ordinate BF,

or ----, is to x the Fluxion of the  $\sqrt{2ax-xx}$ 

Abfcisse AB, as the constant Quantity *a*, to the Curve AF; which was the Property of the *Catenaria* found above. Therefore the points of the *Catenaria* are rightly determin'd by the foregoing Construction. Q: E: D.

## COROL. I.

It is manifest from the Construction, that BF the Ordinate in the *Catenaria*, is equal to the Parabolick Curve AP, taking away BH, the correspondent Ordinate, of the conterminal Hyperbola AH.

#### COROL. II.

'Tis plain from the Demonstration, that the Curve of the Catenaria AF, is equal to BH the correspondent Ordinate of the conterminal Equilateral Hyperbola. For fince the Fluxions of these Lines are equal, and the Lines themselves do arise together, it is manifest that they are always, and every where equal. Whence, giving the Catena, AC or a will be given also, as being equal to the Semiaxis

miaxis of the Equilateral Hyperbola, whofe Vertex is A, and whofe Ordinate belonging the Abfciffe AB, is equal to the *Catena* AD.

#### COROL. III.

All the Catenaria are fimilar to one another; as being generated from the fimilar Conftruction of Similar, and fimilarly posited Figures. From whence it follows, that two right Lines fimilarly inclin'd to the Horizon, carried thro' the Vertices of the Catenaria, will cut off fimilar Figures, and proportional to the Lines cutting off the Portions of the Catenaria.

## COROL. IV.

If the Catena QAD be fuspended on the points Q and D, which are unequally high, the part FAD of the Curve remains the fame as if it were fuspended by the points F and D, which are equally high. For it is no matter, whether the point be fix'd to the Vertical Plane or not.

## COROL. V.

If the force of the Catena drawing in the Direction dD, be divided (as is commonly known) into the force as dA acting with an Horizontal Direction, and the force as ADwith a perpendicular Direction: Then it follows, that the force (in the end of the Catena)

225

rena) of approaching directly to the Axis, is to the force of defeending perpendicularly in the fame (or that part of the fulfaining force that acts in the direction BD, is to that part that acts in the direction DF) as the Semiaxis of the Conterminal Hyperbola AH, to DA the length of the Catena to the Vertex. Whence, the Catena being given, this Ratio is alfo given. And in the fame Catena, fulpended with different degrees of Laxity, that Horizontal force, is as the Axis of the Conterminal Hyperbola; fince DA remains the fame, if the Extremities be equally high.

## COROL. VI.

The Catena placed in an Inverted Polition in a Vertical Plane, maintains its Figure and does not fall down; and fo makes a fine Arch or Fornix. That is, very fmall hard flippery Spheres, difpos'd in the Inverted Catenaria, will form an Arch, no part of which will be thrust inwards or outwards by the reft, but (the lowest Points continuing unmov'd) it is preferv'd by vertue of its Figure. For fince the Position of the Points of the Catenaria, and the Inclination of the parts to the Horizon, is the fame, whether it be in the Polition FAD, or in an Inverted Polition, provided the Curve be in a Plane that is perpendicular to the Horizon, it is evident that it preferves its Figure unchanged, equally in one Polition as the other. And on the other hand, the Catenaria are the 0

the only Genuine Arches or Fornixes. And an Arch of any other Figure, is for this reafon only, fuftain'd, becaufe a Catenaria is included in the thicknefs of it. For if it were very thin, and confifted of parts that were flippery, it would not be fuftain'd. From the foregoing fifth Corol. it may be gather'd with what force an Arch thrufts the Walls outwards, that it ftands upon; for this is the very fame with that part of the force (fuftaining the Catena) which draws with the Horizontal direction. All other Matters requir'd in the ftrength and firmnefs of Walls, that have Arches fet upon them, are Geometrically determin'd from this Theory; which are the principal Things in Building.

#### COROL. VII.

If inftead of Gravity, any other force were fuppos'd acting in like manner upon a flexible Line, the fame Curve would be produced. *Ex. gr.* Suppofe a Wind blowing equably, and in directions parallel to a given right Line, the Line thus inflated by the Wind, would be the fame with the *Catenaria*. For fince all things that were confider'd in Gravity, obtain in this other force, 'tis plain that the fame Curve will be produced.

#### PROP.

### PROP. III. THEOREM.

(Fig. 35.) The Hyperbola AH continuing as before, if through A be drawn the right Line GAL perpendicular to the Axis AB, and the Curve KR be deforib'd of fuch a Nature, that BK be a third proportional to BH and AC, and to the right Line AC be applied the Rectangle AV equal to the Interminate Space ABKRLA; then fball the Point F (the Concourfe of the right Lines HB, VG) be in the Catenaria.

					2	
For by	Conft	ruction	BK :			
~				$\sqrt{2}$	ax-l-x	.2
wherefore	the	Fluxion	of	f the	Spa	ce
Intras				R2	• x	
ABKRLA (	=BKk	b = BK	(Bb)	=		-
				$\sqrt{2}$	axx	2
		Space A	BKR	LA		
And fince E	$F = \cdot$				and A	C
	,	AC	2			
is given,	the 1	Fluxion	of	BF 1	hall =	=
the Flu	ixion o	f the Sp	ace A	BKR	LA	
	- internet		-		2	-
		AC				
		Q 2			4	x .

But in the Conftruction of the  $\sqrt{2ax+x^2}$ 

foregoing Proposition, the Fluxion of the Or-

dinate BF, was =  $\frac{1}{\sqrt{2ax-|-x^2}}$ . Therefore

this Conftruction amounts to the fame with that, and confequently the point F is in the Catenaria. Q: E: D.

### COROL.

As in the foregoing Proposition, the Catenaria is defcrib'd from the length of the Parabolical Curve given; fo in this, the defcription of it depends upon the Quadrature of the Space in which  $x^2y^2 = a^4 - 2axy^2$ . For

BK or y = -----.

Vzar

PROP.

#### PROP. IV. THEOREM.

(Fig. 36.) The Space AGF contain'd under the Catenaria AF, and the right Lines FG, AG, parallel to AB, BF, is equal to the Rectangle under the Semiaxis AC, and DH the difference of the Ordinates in the Hyperbola and Catenaria.

For DH (= BH = BD = by Propofition II.  $\frac{ax - -xx}{\sqrt{2ax - x^2}} - \frac{ax}{\sqrt{2ax - x^2}} = )$ 

Wherefore the Fluxion of the  $\sqrt{2ax-|-x^2}$ Rectangle under the given Line AC and DH axx.

xx

 $(= \frac{1}{\sqrt{2ax-1-x^2}} = x \times \frac{1}{\sqrt{2ax-1-x^2}} = f_s \times FG)$ 

= the Fluxion of the Space AFG. And fince these Figures do arise both together, it follows that the Rectangle under AC and DH is equal to the Space AGF. Q: E: D.

QB COROL

### COROL.

Hence it follows that the Space FAD comprehended under the Catenaria and Horizontal Line FD, is equal to the Rectangle under FD and BA, lefs the Rectangle under either Axis of the Hyperbola AH, and DH the excefs of the right Line BH or the Curve AD, above the Ordinate BD.

#### PROP. V. THEOREM.

(Fig. 36.) If to the right Line AL be applied the Rectangle LE, equal to the Hyperbolical Space ALH, then E will be the Center of Gravity of the Catenaria AFD.

Let the Curve FA be conceiv'd to be librated upon the Axis GL. Then (from the Doctrine of Centers of Gravity) it is manifeft that the *Momentum* of the ponderating Curve FA is expounded by the Superficies of an upright Cylinder crected upon FA, and cut off by a Plane, paffing through GL, and making an Angle of 45° with the Plane of the Curve. And the Fluxion of this Superficies or FA x FG, is equal to the Fluxion of the Space ALH or BH x HL; becaufe FA, BH, as alfo FG and HL, are equal And

And confequently (fince they arife together) the faid Superficies of the upright Cylinder is equal to the Hyperbolical Space ALH. Which therefore divided by the Pondus it felf AF, or its equal the right Line AL, gives the right Line AE, for the diftance of the Center of Gravity from the Axis of Libration GL. So that the point E is the Center of Gravity of the Curve FAD, lying equally on both fides the Axis. Q: E: D.

## COROL. I.

The Spaces ABHL, BAH, and AFG, are in Arithmetick Proportion. For the Fluxion

of the Space ALH is (= -xx + xx) = x = x = x

 $\frac{\sqrt{2ax+x^2}}{\sqrt{2ax+x^2}} = \frac{\sqrt{2ax+x^2}}{\sqrt{2ax+x^2}} = \frac{\sqrt{2ax+x^2}}{\sqrt{2ax+x^2}}$ 

 $x \sqrt{2ax + xx} - \frac{axx}{\sqrt{2ax + x^2}} = \text{to the Fluxion}$   $\sqrt{2ax + x^2}$ 

of the Space BAH lefs the Fluxion of the Space AGF, by Proposition IV. And fince these three Figures do arise together, it follows that BAH — AGF = (ALH =) BL — BAH. Wherefore 2BAH = BL 4- AGF. Q.4 Whence

Whence 'tis plain that the Spaces BL, BAH, and AGF, are in Arithmetical Proportion.

## COROL. II.

The Center of Gravity of the Catenaria, is the Lowest of all those Lines that have the fame Termini, and are of the fame length. For a heavy Body will defeend as far as it can. And fince the Figure it felf defeends as much as its Center of Gravity defeends; 'tis manifest that a flexible heavy Line, will dispose it felf in such a manner, as that its Center of Gravity may be lower, than if it assumed any other Figure. And from this one Property of such a Line, all the rest may easily be deduced.

## COROL. III,

If there be upright Cylinders erected upon any fort of Curves, that are of the fame length, and have the fame Termini D and F, with the Catenaria FAD; and thefe Cylinders be cut by a Plane paffing through DF; then the greatest of all these Superficies shall be that which stands upon the Catenaria. For these Superficies (if the-Angle contain'd under the Planes be half a right one) divided by the Curves (which in the prefent Cafe are all of the fame length) give the distances of the Centers of Gravity from the right Line DF. And fince this distance is greatest in the Catenaria (because of the greatest Descent of the Center of Gravity) therefore . 1 and and

therefore the Cylindrick Superficies shall there also be greatest. Lastly, Because the fame is to be faid of Cylindrick Superficies cut off by a Plane that make any Angle with the Plane of the Bass, as is when the faid Angle is half a right one; the Truth of what was afferted is evident universally.

#### LEMMA.

(Fig. 37.) Any Curve as AFQ, defcrib'd by the Evolution of another Curve KU, if upon any Ordinate, as FB (at right Angles to the Axis AB) be let fall perpendicularly UR, from the correspondent Point U in the Curve KU; then (the Fluxion of the Axis AB continuing the fame) shall the Fluxion of the Fluxion of the Ordinate BF, the Fluxion of the Curve AF, and the right Line FR be continual Proportionals.

Let the Lineola Ff be produc'd 'till it meets the next Ordinate  $w_{\phi}$  in o. And becaufe by the Hypothefis Fs = fw, alfo fhall of = Ff, and confequently  $o_{\phi}$  fhall be the Fluxion of fs, that is the *Fluxion* of the Fluxion of the Ordinate. Farther, the Triangles  $o_{\phi} f$ , fFR, are Equiangular, becaufe  $o_{\phi} f = its$  Alternate fFR, and  $fo_{\phi} = (Ffr =) FfR$ , becaufe their difference R fr is as nothing in refpect of either of them, fince R r is nothing in refpect of fr. And therefore  $o_{\phi}:$  $o_{f}f:: fF: FR;$  but  $\phi f = fF$ , fince they differ

fer but by the Fluxion of either. Therefore also  $o_{\varphi}$ : fF:: fF: FR. Q: E: D.

## PROP. VI. PROBLEM.

(Fig. 37.) To find the Curve KV by the Evolution of which the Catenaria is deforib'd.

Let (as before) AB = x, BF = y. Then by

Proposition II.  $y = \frac{ax}{\sqrt{2ax+xx}}$ , or  $2axy^2 + \sqrt{2ax+xx}$ 

 $xxy^2 = a^2x^2$ . Wherefore (by the Newtonian Method which now generally obtains)  $2axy^2$  $+ 4axyy + 2xxy^2 + 2xyy (= 2axx, which$ becaufe of <math>x = o, fince the conftant x has no Fluxion, is) = o. Therefore y = $(\frac{-axy - xxy}{2ax + xx} =) \frac{a + x}{2ax + xx} \times \frac{ax^2}{\sqrt{2ax + xx}}$ 

putting inftead of y its Value  $\sqrt{2ax - xx}$ 

(For the Sign - before the Quantity y, denotes

235

--- 10- 1 ---

notes only the place of the point R, with respect to F, to be opposite to the place of the point F, with respect to B, when the Curve AFQ is concave towards the Axis a--xxx AB.) And Ff (by Prop. II.) = -V 2ax-1-xx Wherefore (by the foregoing Lemma) FR = $=\frac{\overline{a-x} \times x^2}{2ax+xx} \times \frac{2ax-xx}{\overline{a+x} \times ax^2} \times \frac{2ax-xx}{\overline{a+x} \times ax^2}$ =)  $a + x \times \sqrt{2ax + xx}$ . Again, because of the Rectangular Triangles Fsf, FRU, having the Angles fFs, UFR, equal to one another (because UFs is the Complement of either to a right Angle) we have Fs: sf::  $a - |-x \times \sqrt{2ax} - |-xx$ ax FR:UR, or x :-Vzax--xx UR, which therefore is = a + x. Therefore the Nature of the Curve KU is fuch, that if  $a + x \times \sqrt{2ax} + xx$ AB = x, FR fhall = --, and  $UR = a_{-} x. Q: E: I.$ 

3 - 19 M

Digitized & Google

COROL.

## COROL. I.

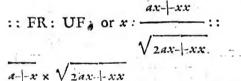
AC: CB:: BH: BH: FR. For this is the Property of the right Line FR, that was found just now.

#### COROL. II.

The right Line CB is = the right Line BI or UR. For each is = a-|-x.

#### COROL. III.

The Evoluent Line UF is a third Proportional to AC and CB. For because of the fimilar Triangles fF s, UFR, it is sF: Ff



. UF, which is therefore

 $= \frac{a+x}{a}$ . Whence a:a+x::a+x: UF,

which is the Radius of the Circle that has the fame Curvature with the Catenaria at the point F.

## COROL.

' Digituday Google

## COROL. IV.

When the point F is in A, or when the Vertex is defcrib'd by the Evolution, that is, when x = o, then the Value of the Evoluent Line (or the Radius of the Curvature) UF (which in this Cafe coincides with KA) viz.  $\frac{a-1-x}{x}$ , becomes only a. That is, the point K where the Curve UK meets the Axis, is as much above the Vertex of the Catena A, as C is below it. Whence the Diameter of a Circle that has the fame Curvature with the Catena at the Vertex, is equal to the Axis of the Conterminal Hyperbola AH. And confequently the Catena AD and the Hyperbola AH, have the fame Curvature in the Vertex A. For it is known that the foremention'd Circle has the fame Curvature with the Equilateral Hyperbola AH, in the Vertex A. But this follows also from the Property of the Catenaria, demonstrated at Proposition II. For the Nascent FH or (AP. = the Nascent BP =)  $\sqrt{8ax}$ , is double the Nascent BH or (Vzax- xx, that is, xx vanifhing, when x is very fmall)  $\sqrt{2ax}$ . And therefore the fame point is as well in the Nafcent Hyperbola, as in the Nafcent Catenaria; that is, the one is coincident with the ! other at their first arising, and confequently

thefe Curves have the fame Curvature at the

Veitex A.

COROL.

227

2012

## COROL. V.

## COROL. VI.

The right Line KI is double of AB. For fince BI = (BC =) CA - AB, alfo AI fhall = CA - 2AB. But AK = AC (by Cor. IV. of this Proposition.) Therefore KI = 2AB.

#### COROL. VII.

The Rectangle AC×BR is = to double the Hyperbolical Space BAH. For FR×AC =  $(a+x\times\sqrt{2ax+xx} + a = a+x\times\sqrt{2ax+xx} = AB \times BH$  $x\times\sqrt{2ax+xx} + a\times\sqrt{2ax+xx} = AB \times BH$ +AC

Dipleted of Google

Miscellanea Curiosa. 239 ---AC×BH=) AB×BH---AC×BD+AC×DH. Wherefore FR×AC-BD×AC (that is, BR× AC) = AB×BH---AC×DH. But (by Propofition IV.) AC×DH=Space AGF. Therefore BR×AC= (ABHL---AGF= by Cor. I. Propofition V.) 2BAH.

## PROP. VII. THEOREM.

(Fig. 37.) If in the Logarithmical Curve LAG (whofe Subtangent HS, given, is equal to the Line a, determin'd as at Cor. II. Prop. II.) be taken the point A, whofe diftance AC from the Afymptote HP, is equal to the Subtangent HS; and from the points H, and P (taken at Liberty in the Affymptote, and equally diftant from the point C) be erected the Lines HL, PG, Ordinates to the Logarithmical Curve, the half Sum of which is equal to HD or PF: Then the points D and F, fhall be pofited in the Curve of the Catenaria, correfponding to the right Line AC.

Let AB be put = x, and confequently CB or DH the half Sum of the Ordinates HL, PG, will = a - |-x|. Let the half difference of them be put = y; whence HL = a - |-x| - y, and PG = a - |-x| - y. And fince from the Nature of the Logarithmical Curve, CA is a mean Proportional between them, as fhall =

Ex.14

aa + 2ax + xx - yy, whence  $y = \sqrt{2ax + xx}$ . Confequently HL =  $a + x + \sqrt{2ax + xx}$ , and PG =  $a + x + \sqrt{2ax + xx}$ . Wherefore the Fluxion of HL, or lm, is ax + xx + xx + xx + xx And because of the

 $\frac{\alpha_{x} - -xx - -x \sqrt{2\alpha_{x}} - xx}{\sqrt{2\alpha_{x}} - xx}$ . And because of the

fimilar Triangles lmL, LHS, 'tis LH: HS:: lm: mL; whence mL or dA the Fluxion of

BD, is = _____. That is, the Curve  $\sqrt{2ax - |-xx|}$ 

AD, generated after the foregoing manner, from the Logarithmical Curve, is of fuch a Nature, that if the Axis be x, and its Fluxion x, the Fluxion of the Ordinate BD

is _____. But this is the Property of  $\sqrt{2ax-1-xx}$ 

the Catenaria corresponding to the right Line a, as was demonstrated at Prop. I. Therefore the Curve FAD defcrib'd as above, is this very Catenaria it felf. Q: E: D.

C 0.

# COROLLARIES.

## COROL. I.

As the Catenaria is defcrib'd by the help of the Logarithms, fo on the other hand, by the help of the Catenaria (a Curve produced by Nature it felf) the Logarithm of any given Number, or rather of any given Ratio, may be found. As if, putting CA=1, whole Log.=0; the Log. of the Number CQ, or of the Ratio between CA and CQ, were fought. Let CV be a third Proportional to CQ and CA, and CB the half Sum of CQ and CV; then an Ordinate to the Catenaria from the point B, viz. BD, will be the Log. fought.

#### COROL. II.

Vice verfa, if giving the Log. CH or CP, the correspondent Number HL or PG, or the Ratio of HL to CA, or PG to CA, be fought. From H or D erect a Perpendicular meeting the Catena in D or F, and in the Horizontal Line AR, take CR=HD or PF, or CB. And then will AR be the half difference of the fought Lines LH, GP, as HD or CR, is (from the above demonstrated Property of the Catenaria) their half Sum. For in three Quantities Geometrically Proportional, fuch as are HL, CA, PG, the Square of the half Sum of the Extreams leffen'd by R

Digitized by Google

the Square of the middle Term, is equal to the Square of the half difference of the Extreams. And confequently CR-|-AR, and CR-AR, are the Numbers HL or GP, agreeing to the given Log. CH or CP.

## COROL. III.

It is plain from the Demonstration, that as HD the half Sum of the Logarithmical Ordinates HL, PG, being applied at right Angles to CH, is an Ordinate to the Catenaria; fo alfo the half difference of the fame HL, PG, applied at right Angles to CA in B, is an Ordinate to the Equilateral Hyperbola, whofe Center is C, and its Vertex A; and confequently (by Cor. II. Prop. II.)

. 1.

= the Catena AD. For  $y = \sqrt{2ax_{+}-xx}$ ; and fince it was fhewn in the foregoing Corol. that AR is alfo the half difference of HL and PG; 'tis plain that AR is = the Portion of the Catenaria AD. From whence by the way, we may obferve a Method, how, from the Catena AD given, to find G the Center of the Conterminal Hyperbola, or the point in the Afymptote of the Logarithmical Curve GL. For taking AR = the Catena AD, and joining the points B, R, from the middle of BR credt a Line perpendicular to it, which will meet BA the Axis of the Catena produced, in the point C, fought. Which is evident, fince thus CR will = CB.

# COROL.

Thilized by Google

## COROL. IV.

Hence also it foilows that if the Angle BDT be equal to ACR, the right Line DT touches the *Catenaria* in D. For then it will be (in the fimilar Triangles DBT; CAR) DB: BT:: CA: AR, or CA: Curve AD which is = AR. And confequently DT touches the *Catenaria*, by *Corol. Prop.* I.

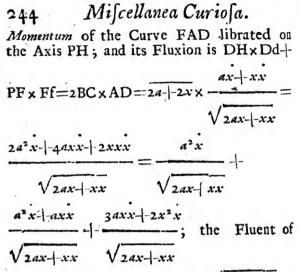
### COROL. V.

It follows alfo that the Space ACHD =the Rectangle  $CA \times AR$ . For becaufe (by *Prop.* IV.),  $AYD = CA \times (AD - BD, = AR - AY, by Cor. III. of this$ *Prop.*=) YR; thething is manifeft. And fince CA is given,its plain that the Space ACHD is as theCurve AD, and the Fluxion of the formerHd, as the Fluxion of the Latter Dd.

# COROL. VI.

If through the point K where CR cuts HD, we draw KZ parallel to PH, meeting AC in Z, and take  $CE = \frac{BC+CZ}{2}$ ; then will E be the Center of Gravity of the Curve FAD. Imagine an upright Cylindrick Superficies erected upon FAD, and to be cut by a Plane palling through PH, and making an Angle of 45 with the Plane of the Curve FAD. This Superficies, will expound the R 2 Momen-

Dia and Google



which,  $a \ge BD - a \sqrt{2ax - xx} - x \sqrt{2ax - xx} =$ CAxBD-| CBxAD. Wherefore CAxBD--CBx AD= (because it arifes together with it) to the foremention'd Cylindrick Superficies = the Momentum of the Curve FAD with respect to the Axis of Libration PH. Whence the diftance of the Center of Gravity of the Curve FAD from the point C,  $\frac{\text{is CA} \times \text{BD} - | \text{ CB} \times \text{AD}}{2\text{AD}}, \text{ or } \frac{1}{2} \frac{\text{CA} \times \text{BD}}{\text{AD}} - | -\frac{1}{2} \text{CB}.$ Farther, becaufe of ZK parallel to AR, 'tis AD: BD:: (AR: ZK::) CA: CZ, whence  $CZ = \frac{CA \times BD}{AD}$ , and therefore CE which by Construction is  $=\frac{1}{2}$  BC  $-\frac{1}{2}$  CZ, shall  $=\frac{1}{2}$  $\frac{CA \times BD}{AD} - \frac{1}{2}$  BC. That is, the Center of Gravity of the Curve FAD, and the point E determin'd

Dented by Google

245

determin'd by this Conftruction, are equally diftant from the point C. But they are alfo pofited in the fame right Line, and towards the fame parts, and therefore they coincide with one another. This Coincidence of the point E as determin'd above, with the Center of Gravity as found at Prop. V. may be thus fynthetically fhewn. By Cor. I. Prop. V. 2BAX = AYD-+ BA×AR. Whence AH -+ 2BAX = (ACHD -+ BA×AR = by Cor. foregoing) AR×CA -+ BA×AR = by Cor. foregoing) AR×CA -+ BA×AR ; that is, BD× AC -+ 2BAX = AR×CB; or BD×AC = AR ×CB - 2BAX. Whence BD×AC -+ AD× BC = (AD×BC -+ AR×CB -- 2BAX = 2AD ×BC -- 2BAX =) 2AD×AC -+ 2AD×AB --2BAX. And dividing by 2AD, we have  $\frac{1}{2}$  $\frac{BD×AC}{AD}$ - $|-\frac{1}{2}BC = (AC-|-\frac{AB×AD--BAX}{AD})$ 

 $CA - \left| -\frac{ARX}{AR} \right|$ . But  $\frac{ARX}{AR}$  is the diffance of the Center of Gravity of the Catena from the Vertex A, determin'd at Prop. V. and confequently, according to the 5th Proposition CA

 $+\frac{ARX}{AR}$  is the diftance of the point E from

C; now  $\frac{1}{2} \frac{BD \times AC}{AD} + \frac{1}{2} BC$ , is the diffance of the point E alfo from the fame point C according to this *Cor*. Whence 'tis manifeft that thefe two Determinations of the point E amount to the fame; becaufe  $CA + \frac{ARX}{AR} =$ 

 $\frac{1}{2} \frac{BD \times AC}{AD} + \frac{1}{2} BC.$ 

R 3

COROL.

# COROL. VII.

The Center of Gravity of the Space PFADH, is in I the middle point of the right Line CE. For fince the Center of Gravity of the Fluxion of AD, or Dd, and Ff, is twice as far diftant from PH, as is the Center of Gravity of the Fluxion of ACHD, or DHhd, and FPpf; and Dd-| Ffx AC is = DdhH+FfPp; 'tis plain that E, the Center of Gravity of the Fluent FAD, is twice as far distant from PH, as I, the Center of Gravity of the Fluent PFADH. But this may be yet shewn otherwife according to the Method us'd before. Imagine an upright Cylinder to be crected upon the Figure PFADH, and to be cut off by a Plane paffing through PH, and making an Angle of 45 with the Plane of the Basis. This Solid will expound the Momentum of the Figure PFADH librated on the Axis PH. And the Fluxion of this Solid or Momentum (viz. the Solids erected on the Bafis PFfp, and HDdh) is produced, by multiplying the Momentum of the Fluxion, or the Fluxion of the Momentum, into 1 AC given. For by Cor. V. of this Proposition HDdh=Ddx AC. Wherefore the Fluent Momentum it felf, is produced by multiplying the Momentum of the Curve FAD with refpect to the Axis PH (as determin'd at Cor. foregoing) viz. CAxBD--CBxAD into 1 AC; which will therefore be  $\frac{1}{2}$  ACx ACx BD- $\frac{1}{2}$  ACx CBx AD. And confequently if this be divided by the librated

librated Figure PFADH (=2CA $\times$ AD, by Cor. V. of this Proposition) there will arife (for the diffance of the Center of Gravity of the Figure PFADH from the Axis PH)  $\frac{1}{4}$ CA $\times$ BD

 $\frac{1}{AD}$  +  $\frac{1}{4}$  CB; which is =  $\frac{1}{4}$  CE determin'd above.

# COROL. VIII.

If through the point N where DT the Tangent to the Catenaria in D, cuts the Line AR, be drawn a Parallel to BC, meeting in O a parallel to AR drawn through E; then will O be the Center of Gravity of the Curve AD. For by Cor. 6. the Center of Gravity of the Curve AD is in the right Line EO. But it shall be demonstrated to be in the right Line NO; and confequently that O it felf shall be the point. Let DA be conceiv'd to be librated upon the Axis HL; then the Momentum of this is the Curve DA multiplied into the diftance of the Center of Gravity from HL. And confequently its Fluxion =  $DA \times Hh$  (Hh being the Fluxion of the diftance of the Axis of Libration from the Center of Gra-

vity) =  $\sqrt{2ax - |-xx|} \times \frac{ax}{\sqrt{2ax - |-x^2|}} = ax$ . And

therefore the Momentum of the Curve DA, with refpect to the Axis HL, is = ax. And confequently the diftance of the Center of Gravity from the fame Axis, is ax divided R 4 by

by AD, or  $\frac{AC \times DY}{AR}$ . But because DT touches the Catenaria, by Cor. 4. of this Proposition, the Angle BDT, or DNY=ACR, and the Angles at A and Y are right ones, therefore in the Equiangular Triangles RAC, DYN, 'tis RA: AC:: DY: YN; whence  $YN = \frac{AC \times DY}{RA}$ , that is YN is the diffance of the Center of Gravity of the Catena AD from the Axis HL; or that Center is in the right Line NO.

# COROL. IX.

If through the point I be drawn a right Line parallel to AR, meeting ON produc'd in W; then W fhall be the Center of Gravity of the Space ACHD. For by Cor. 7. the Center of Gravity of the Space ACHD, is in the right Line IW, but it fhall be demonstrated alfo that 'tis in NW, and confequently W is the point. For (after the fame manner as in Cor. foregoing) the Fluxion of the Momentum of the Space ACHD ponderating upon the Axis HL, will be fhewn to be

 $(ACHD \times Hh = AC \times AD \times Hh =) a\sqrt{2ax + xx}$ 

 $\times \frac{ax}{ax} = a^2 x.$  And confequently the

Viax-+xx

Momentum of the Space ACHD, with refpect to the Axis HL, is the Fluent of this Fluxion,

Data day Google

Fluxion,  $a^2x$ , that is,  $a^2x$ . This therefore divided by the Space ACHD, or

 $a \times \sqrt{2ax} - xx$ , gives the diffance of the Center of Gravity (of the Space ACHD) from the Axis HL, which is  $= \frac{ax}{\sqrt{2ax} - xx}$ 

 $= \frac{AC \times DY}{AR}$  And therefore the Center of Gravity of the Space ACHD, is in the Line NW. And from thefe two laft *Corollaries*, is found the Center of Gravity of any Portion of the *Catena*, though not reaching the Vertex A, or alfo of any Space comprehended under any Portion of the *Catenaria*, and any other right Lines befides those aforefaid.

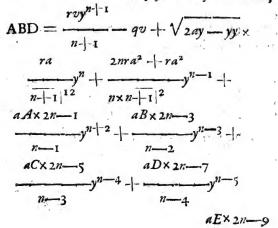
# COROL. X.

Hence are measur'd the Surfaces and Solids generated by the Rotation of the Catena (or a Space comprehended under it, and a right Line) about any given Axis. For a Figure generated by fuch a Rotation, is (as is vulgarly known) equal to the genenerating Figure multiplied into the Periphery defcrib'd by the Center of Gravity in the Rotation, which Periphery is given, fince the Radius or Distance of the Center of Gravity from the given Axis, is given. Thus if the Catena AD roul'd about the Axis AB, then

then  $\frac{\pi}{\bullet}$  AN is the Periphery defcrib'd by the Center of Gravity O ( $\frac{\pi}{\rho}$  denoting the Ratio of the Periphery of a Circle to the Radius) and confequently the Surface generated by the Rotation of the Catena AD =  $(\frac{\pi}{2} \times$  $AN \times AD = \frac{\pi}{\rho} \times AN \times AR$ . That is a Circle, the Square of whofe Radius is double the Rectangle RAN, will = the Surface generated by the Rotation of the Catena AD about the Axis AB. After the fame manner the Solid generated by the Rotation of the Space ACHD about AC, may be fhewn to be equal to a Cylinder, whose Basis is the foremention'd Circle, and its Altitude=AC. Thus also the Surfaces and Solids produced by the Rotation of these Figures about any other given Axis, are meafur'd. For giving the Center of Gravity, they are eafily difcover'd.

# Of the Quadratures of Geometrically irrational Figures. By J. Craig.

**L** E T ACF (Fig. 38.) be a Semicircle, metrically irrational Curve, whofe Ordinate BD cuts the Semicircle in C. The Quantities may be noted thus; The Diameter AF = 2a, the Abfciffe AB = y, the Arc AC = v, the Ordinate BD = z. And let  $z = rvy^n$ a General Equation expreffing the Nature of the Geometrically irrational Curves ADE, in which r denotes any given and determin'd Quantity, and n an indefinite Exponent of the indetermin'd Quantity y. I fay the Arca,



252 Miscellanea Curiosa. 

In this Infinite Series, these things are to be taken notice of: (1.) That the Capital Letters A, B, C, D, E,  $\mathcal{C}c$ . denote the Coefficients of the Terms immediately preceeding them, viz.  $A = \frac{2mraa - |-raa}{n \times n - |-1 \times n - |-1} B =$  $AA \times 2n - 1$   $AB \times 2n - 3$ n - 1,  $C = \frac{aB \times 2n - 3}{n - 2}$ , and fo on. (2.)

That if the Exponent n be an Integer and Positive, or equal to nothing, or if 2n be an odd Number, then the Quadrature of the Space ABD may be exhibited by a finite Quantity: The Series in these Cases breaking off. (3.) That q denotes the Term last breaking off. (4.) That all those Figures in which the Series is broke off have one Geometrically Quadrable Portion very easily affignable from the Series it felf, viz. if you make

the Abfciffe  $y = r^{n+1} + nq + q^{n+1}$ ; there will arife a Geometrically Quadrable Area anfwering to this Abfciffe. (5.) That only the Irrational Terms  $\sqrt{2ay-yy}$  is to be multitiplied into the Terms following it.

Example

#### Example I.

Let z = v, becaufe in this Cafe r = 1, n = o, therefore  $\frac{ra}{n-|-1|^2}y^n$  is the Term laft breaking off, wherefore q = a, whence ABD  $= vy - av - a\sqrt{2ay - y2}$ : And confequently if (by Note 4.) you take the Abfciffe y = a, that is, if the Ordinate pafs through the Center of the Circle, there will arife a Geometrically Quadrable Portion fitting it, viz. Area =  $a^2$ , that is, the Square of the Radius.

#### Example II.

Let  $z = \frac{vy}{a}$ . Becaufe in this Cafe  $r = \frac{1}{a}$ , n = 1, therefore  $\frac{2na^2 - |-ra^2}{n \times n - |-1|^2}$  is the Term laft breaking off, wherefore  $q = \frac{3^4}{4}$ ; whence ABD  $= \frac{vy^2}{2a} - \frac{3av}{4} - |-\frac{y - 3a}{4} \sqrt{2ay - y^2}$ , and confequently, if (by Note 4.) you take  $y = \sqrt{\frac{3aa}{2}}$ , there will arife a Geometrically Quadrable

Data da Google

# Miscellanea Curiosa. 254 Quadrable Area fitting this Abscisse, viz. Area = $\sqrt{:\sqrt{6a^4} - \frac{3a^2}{2}} \times \sqrt{\frac{3a^2}{2}} + \frac{3a}{4}$ . Example III. Let $z = \frac{vy^2}{4a}$ ; In this Cafe $r = \frac{1}{aa}, n=z$ , therefore $\frac{aA \times 2n-1}{n-1} y^{n-2}$ is the Term laft breaking off, therefore $q = \frac{5\pi}{6}$ ; whence by Infinite Series, will ABD = $6vy^3 - 15a^3v - |-2ay^2 - |-5a^2y - |-15^3a\sqrt{2ay - y^2}$ 1842 And confequently, if (by Note 4.) you take $y = \sqrt[3]{\frac{5a^3}{2}}$ , there will arife a Geometrically Quadrable Area fitting this Absciffe, viz. Area = $\frac{2ay^2 + 5a^2y + 15a^2}{18a} \times \sqrt{2ay - y^2}$ .

Secondly, Let ACF (Fig. 39.) be a Parabola, AE its Axis, A the Vertex, and (B a) the Latus Reftum. And let ADG be a Geometrically irrational Curve, whole Ordinate BD cuts the Parabola in C. Let the Abfciffe AB = y,

AB = y, the Ordinate BD = z, the Arc of the *Parabola* AC = v. And let the General Equation expressing the Nature of Infinite irrational Curves be this,  $Z = rvy^n$ , in which r denotes a given and determinate Quantity, and n an indefinite Exponent of the indetermin'd Quantity y. I fay the Area

A B D = $\frac{ry^{n- -1} \times v}{n- -1} - qv -  -\sqrt{2ay}$	$y^2 - y^2 x - y^2 - y^$
$\frac{r}{n- -2\times n- -1 }y^{n- -1} - \frac{ra}{n- -2\times n- -1 ^2}y^n -$	+
$\frac{ra^2 \times 2n - 1}{n \times n - 1 - 2 \times n - 1} y^n - 1 - \frac{aA \times 2n - 1}{n - 1}$	y ⁿ⁻² +
$\frac{aB \times 2n - 3}{n - 2} y^{n-3} - \frac{aC \times 2n - 5}{n - 3} y^{n-4} -$	- &c.

In this Series 'tis to be noted: (1.) That the Capital Letters A, B, C, G'c. denote the Coefficients of the Term preceding them. (2.) That if the Exponent *n* be an Integer and Pofitive, or equal to nothing, or if 2n be an odd Number, then the Quadrature may be exhibited by a finite Number of Terms; the Series in these Cafes breaking off. (3.) That -|-q| is equal to the Term last breaking off. (4.) That of the Terms multiplying the Quantity  $\sqrt{2ay} - \frac{1}{y^2}$ , the last breaking off is to be doubl'd. (5.) That all those Figures in which *n* is an Integer, Positive and an odd Number, or more generally, all those Figures in which the last Term breaking off has

has an Affirmative Sign or +, have one Geometrically Quadrable Portion, and affignable from the Series it felf, by taking the Absciffe as in the fourth Note of the preceding Series.

#### Example I.

Let z = v, becaufe in this Cafe r = 1, n = o, therefore the Term laft breaking off is  $\frac{ra}{n-|-2\times n-|-1|^2}y^n$ , whence  $-|-q = \frac{a}{2}$  (by Note 3.) and becaufe in this Cafe  $-\frac{a}{2}$  is the laft Term to be multiplied into  $\sqrt{2ay-y^2}$ , therefore ABD =  $vy + \frac{av}{2} + \sqrt{2ay + y^2} \times \frac{1}{2}y - a$ .

### Example II.

Let  $z = \frac{vy}{a}$ , because in this Case  $r = \frac{1}{a}$ , n = 1, therefore the Term last breaking off is  $\frac{a^2 \times 2n - |-1|}{n \times n - |-2 \times n - |-1|^2} y^{n-1} = \frac{a}{4}$ , whence  $q = \frac{1}{4}a$ , and  $\frac{1}{2}a$  is the last Term to be multiplied by  $\sqrt{2ay - |y^2}$ , therefore

ABD

Diament in Google

Mifcellanea Curiofa. 257  $ABD = \frac{vy^2}{a^2} - \frac{av}{4} + \sqrt{2ay} + \frac{y^2}{2ay} - \frac{y^2}{6a} - \frac{7}{12} + \frac{a}{2}$ And if you take  $y = \sqrt{\frac{1}{2}aa}$ , there will arife a Geometrically Quadrable Area fitting this Abfcille, viz. Area =  $\frac{1}{12}$ 

$$V: \sqrt{2a^4} - \left| -\frac{a^2}{2} \times 5a - \sqrt{\frac{1}{2}a^2} \right|$$

I have other Theorems of this Nature, for Figures depending on the Circle and Parabola; but thefe two may fuffice as a Specimen to fhew the Ufe of my Method publifh'd in my *Treatife of Quadratures*, in determining the Quadratures of Irrational Figures, for which there has been no Method (as far as I know) as yet made Publick.

That the Reader may the more eafily come at the Invention of these and the like Theorems, I shall subjoin another, and more hereafter, if need be.

Let therefore (Fig. 40.) ACF be a Semicircle, ADE a Geometrically Irrational Curve, whofe Ordinate BD cuts the Semicircle in C. Let the Quantities be denoted as before, viz. the Diameter AF = 2a, the Abfciffe AB = y, the Arc AC = v, the Ordinate BD = z; and let  $z = rv^2y^n$ , an Equation expreffing the Nature of the Curves ADE, in which r denotes any given and determin'd Quantity, and n an indefinite Exponent of the indetermin'd Quantity y. I fay the Area

ABD

58 Mifcellanea Curiofa.  
ABD = 
$$\frac{rv^2y^{n+1}}{n-1}$$
  $qv^2 + v\sqrt{2ay - y^2} \times \frac{2ra}{n-1} y^n + \frac{2ra^2 \times 2n-1}{n \times n-1} y^{n-1} + \frac{aA \times 2n-1}{n-1} y^{n-2} + \frac{aB \times 2n-3}{n-2} y^{n-3} + \frac{aC \times 2n-5}{n-3} y^{n-4} + \frac{aD \times 2n-7}{n-4} y^{n-5} + \frac{aE \times 2n-9}{n-5} y^{n-6} \&c. - \frac{2ra^2}{n-1} y^{n-1} - \frac{2ra^3 \times 2n-1}{n^2 \times n-1-1} y^n - \frac{a^2 A \times 2n-4}{n-1} y^{n-1} - \frac{a^2 B \times 2n-3}{n-2} y^{n-2} - \frac{a^2 C \times 2n-5}{n-3} y^{n-3} \&c.$ 

In this Theorem these Things are to be taken notice of; (1.) That 'tis made up of two Infinite Series, the former of which (connected by the Sign -]-) is multiplied into  $v\sqrt{2ay-y^2}$ ; but the Terms of the latter (affected by the Sign —) are Absolute. (2.) That in the former Series, the Capital Letters, A, B, C,  $\mathfrak{S}c$ . denote the Coefficients of the Terms respectively preceding them; and in the latter have the same Values as in the former

former. (3.) That the Quadrature may be express'd by a finite Quantity, when n is a politive Integer, or equal to nothing, or if 2n be an odd Number; for in these Cafes each Series is broke off. (4.) That 2q is equal to the last Term breaking off, of the former Series.

# Example I.

Let  $z = \frac{v^2}{a}$ . Because in this Case  $n = a_1$ 

 $r = \frac{1}{2}$ , therefore shall the Area ABD  $= \frac{yv^2}{a}$ 

v²-1-2v V 2ay-y²-2ay.

#### COROLLARY.

The whole Figure AFE is equal to twice the Square, whole fide is ACF, lefs the Square of the Diameter.

### Example II.

Let  $z = \frac{yv^2}{a^2}$ , becaufe in this Cafe n = 1,

 $r = \frac{1}{r^2}$ , therefore fhall the Area ABD =

S 2

y2 22

260 Miscellanea Curiosa.  $\frac{y^2 v^2}{2a^2} - \frac{1}{4} v^2 + v \sqrt{2ay - y^2} \times \frac{y}{2a} + \frac{1}{2} - \frac{1}{4} y^2 - \frac{1}{4} y^2 - \frac{1}{4} - \frac{1}{4} y^2 - \frac{1}{4} - \frac$ 3 ay Example III. Let  $z = \frac{y^2 v^2}{z^3}$ , because in this Case n = 2,  $r = \frac{1}{a^3}$ , therefore fhall the Area ABD  $= \frac{y^3 v^2}{2a^3}$  $-\frac{5}{6}v^{2} + v\sqrt{2ay-y^{2}} \times \frac{2y^{2}}{9a^{2}} + \frac{5y}{9a} + \frac{5}{3} - \frac{5}{3$  $\frac{2y^3}{27a} - \frac{57^2}{18} - \frac{5ay}{3}$ 

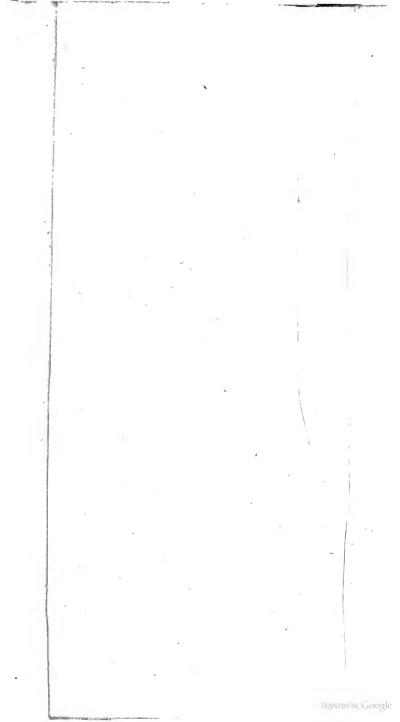
While I was writing this, I receiv'd the late Months of the Lipfick Acts, in which I read, with a deal of pleafure, feveral excellent things for promoting Geometry; and among them fome Remarks of Mr. Leibnitz, and Mr. J. Bernouilli, upon my Method of Quadratures. In the Acts of April, An. 1695. Mr. Leibnitz informs us that he has a Method fomewhat like ours; and truly, I mightily Congratulate my felf, that any thing of mine could have the leaft likenefs to the Thoughts of fo great a Geometer. But whereas he fays his own is much more General, and fhorter than mine; I make no doubt of that. It were to be wifh'd, he would

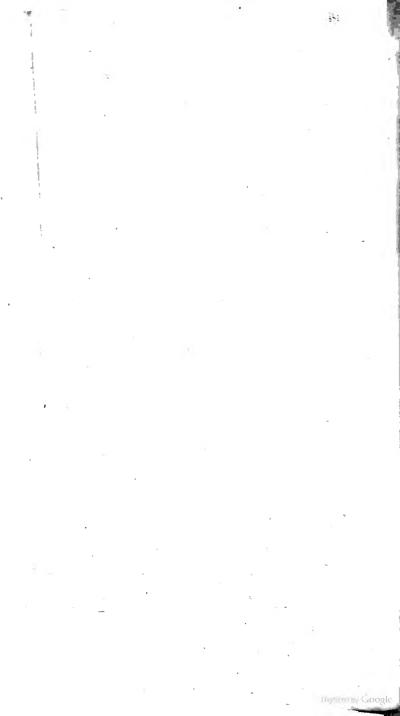
would no longer suppress this Method of his, and feveral other things he has, efpecially relating to his Differential Calculus, but rather, as foon as his Leifure permits, publish them for the Good of the Commonwealth of Learning. In the mean while we hope the Illustrious Marquess De PHospital will fpeedily make publick what is neceffary to perfect that Calculus, in the latter part of that excellent Work of his, which (in the Preface to the former part) he informs us, he has compos'd upon the Integral Calculus. We expect alfo, with fome Impatience, that other Section, in which that Noble Author promifes he will fhew the Ufe of his Calculus in Physicks and Mechanicks. For whatever he has publish'd, as well those Specimen to be found fcatter'd in the Lipfick Acts, and elsewhere, as that excellent Book of his (Intitul'd, Analyse des Infiniment petits) cause us to expect great Things from that Noble Marquefs.

Whereas the Ingenious Mr. J. Bernouilli has thought fit (in the Acts of February and August, An. 1695.) to fay my Method is not General, I freely confess it, as that Sagacious Person might easily perceive in the Course of my Examples. In a Matter so Intricate I took what Steps I could; and if deter'd with the length and difficulty of the Journey, I then made no farther Progress. I might fairly make a Step where I please, fince my Application to these Mathematical Studies is only by the by. Mr. Bernouilli has partly hinted where my Method is at a S 3 Stand,

Stand, though he feems not to have taken up the whole Matter. In the mean while I acknowledge my felf highly oblig'd, that he has honour'd my Treatife with his Animadversions; but much more fo, that he was willing to free me of my Mistakes, with fo much Candor and Humanity.

Con-





Concerning the apparent Magnitude of the Sun and Moon, or the aprent distance of two Stars when nigh the Horizon, and when higher elevated.

Do not defign fo much to establish any thing of my own that may be fatisfactory in folving this admirable appearance, as to detect the Errors of those that have offered at a Solution thereof, and have come short (as I conceive) of being fatisfactory; that thereby I may again fet the minds of Philofophers on work, and roufe them up to enquire anew after this furprizing Phanomenon. That I may do this the more effectually, I shall briefly declare the Matter of Fact, and then proceed to the Reafon thereof, given by feveral, and to their Confutations.

First therefore it is well known that the mean apparent Magnitude of the Moon is 30 m. 30 f. we will take it Numero Rotundo to de 30, that is, an Arch of a great Circle in the Heavens of 30 Minuts is covered by her Diameter; and this we'll suppose to be her apparent Diameter, at a full Moon in the midft of Winter, and when the's in the Meridian, and at her greatest Northern Latitude, and confequently the utmost that she can be elevated in our Horizon : 'Tis as well known

S 4

known alfo that when fhe is in this pofture, being looked upon by the naked Eye fhe appears (that we may accommodate all to fenfible Measures) to be Magnitudinis Pedalis. about a foot broad. But the fame Moon being looked upon just as the rifes, the appears to be three or four foot broad, and yet if with an Inftrument we take her Diameter. both in one posture and t'other, we shall find that still she shall be but 30 Minutes; the feveral ways of trying this I will not mention, they being as various as are the Methods of taking the Moons apparent Diameter, common enough among the Aftronomers; neither will I infift upon the truth of the Matter of Fact, for that I think cannot reasonably be questioned, after fo many trials and fo many experiments thereof, faithfully recorded by undoubted Witnesses; and it would be very unreafonable to imagine that fo many Authors should rack their Brains for folving an appearance wherein they were not certain of the matter of Fact. But because of Nullins in Verba, I can affert that I have accurately try'd it my felf, and I have fo found it : One of the ways I proceeded was thus; I took a very good Telescope of about 6 foot long, in the inward Focus of whole Eye-Glafs I apply'd a very fine Lattice made of the fingle hairs of a Man's Head; then looking with this at the Moon when the was just rifen and looked extraordinary big, I obferved what number of the squares of the Lattice were occupy'd by her Body; then obferving her again, when more elevated and free from all extravagant Greatness, I still found

265

found the fame fquares of the Lattice poffeffed by her. This way is equivalent to that now more ufed, of taking her Diameter by Mr. *Townly's Micrometers*; but I have alfo tried and found the fame thing by an accurate Sextant, taking the diftance of the Moons opposite Limbs.

Now this Phanomenon affords two things to be confidered, first why the Moon (I still name the Moon as being an Object more adapted for our fight, for the fame thing holds in the Sun) should feem bigger about the Horizon, then when more elevated ; and fecondly, fhe appearing bigger, how comes it to pass that her Diameter being taken, it is no greater than when the appears lefs. But the Difquifition concerning this latter being likely to comprehend the former, I shall not divide my Discourse into two Branches, but proceed in the Method proposed. Only I defire it may be noted, that I suppose the Horizontal and Meridional Moon to be found both of the fame Angle, whereas in truth the the Meridional Moon (tho' appearing lefs) fhall be found of the greater Angle : which increafeth the Wonder. But this proceeding from the different distances that one and the other is looked at (the Meridional Moon being nigher us by almost a Semidiameter of the Earth) and confequently eafily folved that way; I have therefore chofen to put between them a plain equality, for avoiding Confusion and Intricacy in Difcourfe.

Wherefore let us hear what the Ingenious of these latter days can say to this appearance. And first we find the Celebrated Des-Cartes attributing

tributing this appearance rather to a deceived Judgment than to any Natural Affection of the Organ or Medium of fence; for the Moon (fays he) being nigh the Horizon, we have a better opportunity and advantage of making an Estimate of her, by comparing her with the various Objects that incur the fight, in its way towards her; fo that tho' we imagine fhe looks bigger yet 'tis a meer deceit; for we only think fo, becaufe the feems nigher the tops of Trees or Chimneys or Houfes or a space of Ground, to which we can compare her, and effimate her thereby ; but when we bring her to the Teft of an Inftrument that cannot be deluded or imposed upon by these appearances, then we find our Estimate wrong, and our Senfes deceived. Thefe Thoughts, methinks, are much below the accustomed accuracy of the noble Des Cartes; for certainly if it be fo, I may at any time increase the apparent bigness of the Moon, tho' in the Meridian; for it would be only by getting behind a Clufter of Chimneys, a Ridg of a Hill, or the top of Houses, and comparing her to them in that posture, as well as in the Horizon; belides if the Moon be look'd at just as she is Rising from an Horizon determined by a fmooth Sea, and which has no more Variety of Objects to compare her to, than the pure Air; yet fhe will feem big, as if lookt at over the rugged top of an uneven Town or rocky Country. Moreover, all variety of adjoining Objects may be taken off, by looking through an empty Tube, and yet the deluded imagination is not at all helped thereby. I come next to the folution hereof given

267

given by the famous Thomas Hobbs; and for this we shall stand in need of Figure 41. wherein, fays he, let the point G be the Center of the Earth, and F the Eye on the furface of the Earth ; on the fame Center G let there be ftruck the two Arches, EH determining the Atmosphere, and A D to reprefent that blue furface in which we imagine the fixed Stars; and let FD be the Horizon. Divide the Arch A D into three equal parts by the lines BF, CF, it is manifest that the Angle AFB is greater than the Angle BFC. and this again greater than the Angle CFD. Wherefore fays he, to make the Angle CFD equal to the Angle CFD, the Arch CD must be greater than the Arch CB; and confequently, that the Moon may in the Horizon appear under the fame Angle as when elevated, fhe must cover a greater Arch, and therefore feem greater; that is, the Moon in the Meridian appearing under the Angle BFC, that fhe may appear under an equal Angle in the Horizon, as suppose CFD, 'tis neceffary the Arch CD should be greater than CB; and confequently tho' fhe appear to fubtend a greater Arch when in the Horizon then when elevated, yet she appears under the fame Angle. And all this without Refraction. The Geometry of this Figure is most certainly true and demonstrable. At this I quarrel not ; but it makes no more in our prefent Difficulty than if nothing had been faid; for the Philosopher has here made a Figure of his own, and from thence he argues as confidently, as if Nature would accommodate her felf to his Scheme, and he not

not oblig'd to accommodate his Scheme to Nature; for here he has made the Circle GF reprefenting the Earth very large in proportion to the Circle AD; and then indeed taking the point F in the Earth's furface, and by lines from thence dividing the Angle AFD into what ever equal parts the intercepted Arches AB, BC, CD, shall be unequal. But if he had confidered, that the Earth is as it were a point in respect of the Sphere of the fix'd Stars, nay the very annual Orbit of the Earth is almost if not altogether imperceptible (faving the truth of Mr. Hook's Attempt) he would have found that the Lines FB, FC, FD, must be all conceived as drown from the point G, and then equal Angles will intercept equal Arches, and equal Arches equal Angles: And fo it happens (at leaft beyond the poffibility of difcovery of fenfe) to the Eye on the furface of the Earth. And befides he should have confidered, that all. Obfervations Aftronomical are performed as from the Center of the Earth, and therefore it is that they keep fuch a ftir about a Parallax; fo that his drawing his lines fo far from G as F is, and to another concentrick Circle fo nigh as AD, deceived him in this Point.

The famous Gaffendus has written 4 large Epiftles on this Subject, the fubftance of all which is, that the Moon being nigh the Harizon and looked at through a more foggy Air, cafts a weaker Light, and confequently forces not the Eye fo much as when brighter; and therefore the Pupil does more inlarge it felf, thereby transmitting a larger Projection on

Walland by Google

269

on the Retina. In this Opinion I do find he is not alone, for in the Journals des Scavans this Difquifition being again revived by a French Abbe, he therein follows this Sentiment of Gassendus. It was first published in the 3d Conference presented to the Dauphin in August 1672. but by reason of an Objection moved by Father Pardye, it was fain to be re-published with fome additions and amendments in Octob. 1672. The addition was, that this contracting and enlarging of the Pupil causeth a different shape in the Eye; an open Pupil making the Crystalline flatter and the Eye longer, and the narrower Pupil fhortning the Eye, and making the Crystalline more convex, the first attends our looking at Objects which are remote or which we think fo; the latter accompanies the viewing Objects nigh at end. Likewife an open Pupil and flat Crystalline attends Objects of a more fedate Light, whilft Objects of more forcible Rays require a greater Convexity and narrow Pupil. From these Politions the Abbe endeavoured to give an account of our Phanomenon as follows.' When the Moon is nigh the Horizon, by comparison with interposed Objects, we are apt to imagine her much farther from us then when more elevated, and therefore (fays he) we order our Eyes as for viewing an Object farther from us; that is, we fomething enlarge the Pupil, and thereby make the Crystalline more flat; moreover the duskishness of the Moon in that posture does not fo much strain the fight; and confequently the Pupil will be more large, and the Crystalline more flat: Hence a larger Image fhall

270

shall be projected on the Fund of the Eve. and therefore the Moon shall appear larger. And this disposition of the Eye that magnifies her, magnifies also the divisions of our forementioned Lattice, and confequently she by her Body shall posses no more of the divisions than when the feems lefs. These two forementioned accidents, viz. the Moons imaginary distance and duskishness, gradually vanishing as the rifes, a different Species is hereby introduced in the Eye, and confequently the feems gradually lefs and lefs, 'till again the approaches nigh the Horizon. Thefe two Opinions of Gallendus and the Abbe being fo nigh a-kin, I shall consider them both together, and first I affert that a wider or narrower Aperture increases not, neither diminifhes the projection on the Retina. I know Honoratus Faber in his Synopfis Optica endeavours to prove the clear contrary to this my Affertion, and that after this manner. Fig. 42. A B is an Object, E F the greater aperture of the Pupil, admitting the projection KI on the Retina, whereas the leffer aperture CD admits only the projection GH; but GH is less than KI, wherefore a lesser aperture diminishes the projection. I admire that any Man that undertook (as Honoratus Faber) to write of Opticks more accurately than all that went before him, should be guilty of fo very grofs an Error; and I do more admire that the celebrated Gaffendus, and with him the noble Hevelins fhould be of the fame Opinion: For tho' the 'forefaid Figure and Demonstration hold most certainly true in direct projections, as in a dark Room with

Danied by Google

271

with a plain hole; yet it will not hold in Projections made by Refraction, as it is in those on the Retina in the Eye, by means of the Crystalline and other Coats and Humours of the Eve. For a Demonstration of this obferve Fig. 43. wherein let A B be a remote Object, and EF the Crystalline at its large aperture, projecting the Image 1M on the Retina. Let then CD be the lesser aperture of the Pupil before the Crystalline : I fay the Image IM shall be projected as large as before, for the Cone of Rays EAF confifts partly of the Cone of Rays CAD, therefore where the former EAF is projected, the latter CAD, as being a part of the former, shall be projected alfo. So that no more is effected by this narrow Aperture, but that the fides of the radiating Cones are intercepted, and confequently the Point I shall be affected with lefs light, but it shall still be in the fame place: What is faid of that Cone and that Point may be faid of all other Cones and other Points of the Object. From hence appears first, the Invalidity of the Account given of the Moons appearance by Gaffendus from this Reafon. 2dly, The Reafon appears why a Telescopes greater or leffer Aperture, makes no difference in the Angle it receives; for imagine EF to be an Objectglass of a Telescope, and 'tis plain. 3dly, 'Tis evident why a greater or lefs Aperture on a Telescope should make the Objects appear Lighter or Darker, for thereby more or lefs Rays are admitted to determine on the Projection of each Point. But all this by the by. And this is fufficient for a Confutation of

of Gallendus and Faber : But our forementioned Abbe fuperadds to a greater or leffer Aperture of the Pupil, as a necessary Confequent. a greater and leffer Convexity of the Cryftalline, as alfo a lengthening and fhortening the Tube of the Eye: And this I must confess would do fomething if we find it true in our Cafe; and this let us try. First, fays he, the duskishness of the Moon nigh the Horizon admits the Pupil to enlarge it felf, the Crystalline to flatten, and the Eye to lengthen. But what if we change our Object, and instead of the Moon take the distance between fome of the fixt Stars; as fuppofe those of Orions Girdle) we shall find the fame Phanomenon in them, and yet I hope neither he nor Gaffendus will affert, that they at one time ftrain the Eye more than at another, or that at any time their fulgur strains the Eye at all; if he do, let him take Stars of the leffer Magnitudes, nay even those that can but just be perceived, and then he will be convinced : Or let him confider whether this will hold in looking at the Sun through very dark Glasses, which render the Sight thereof as inoffensive to the Eye, as that of a green Field. But perhaps he will then fay that this other Reason holds, which is 2dly, That the greater imaginary diftance at which we think the Moon near the Horizon, than when more elevated, makes us Contemplate her as if really the was fo, viz. with ample Pupils, Gr. but this I have fufficiently overthrown in my Remarks against Des Cartes; therefore I pass it over, only fubjoining that if there were any thing in this Surmife, my-thinks the Horizontal

273

rizontal Moon fhould be fancied nigher to us than farther from us; for if we are for trying natural Thoughts, let us take Children to determine the Matter, who are apt to think that could they go to the edge of that fpace that bounds their Sight, they fhould be able (as they call it) to touch the Sky; and confequently the Moon feems then rather nigher to us than farther from us.

After I had writ thus far I accidentally caft my Eye upon Riccioli's Treatife of Refraction, at the end of his 2d. Volume of the Almagest, Lib. 10. Sect. 6. Cap. 1. Queft. 13. wherin he speaks of our present Difficulty; but to my wonder I find him affert, that he and Father Grimaldi had often taken the Horizontal Sun and Moons Diameter by a Sextant, when to the naked Eye they appeared very large ; (Grimaldus directing his Sight to the left edge, and Ricciolus to the right,) and that even by the Instrument they always found the Diameters greater than when more elevated, the Sun often fubtending an Angle of almost a Degree, and frequently 45 Minutes, the Moon alfo 38 or 40 Minutes. This is down right contrary to the matter of Fact which I have before alledged, and directly repugnant to the matter of Fact afferted by the French Abbe in the forecited Journal. Whether of us be in the right I leave to accurate Experiment to determine, and fubmit the whole to the decision of the Illustrious Royal Society. Only give me leave to add one word against Riccioli, for had his Experiments been accurately profecuted, he should Т have

Digitized by Google

have tryed them when the Horizontal Moon had look'd ten times more large in Diameter than ordinary; and then if it be true, that even by an Inftrument fhe will be found proportionally broader than really fhe fhould fubtend an Angle of 300 Minutes, or 5 Degrees: for very often I have feen the Moon when fhe appeared 10 times broader than ordinary, which the fmall addition of 8 or 10 Minutes to her ufual Diameter will never Caufe.

Lastly, as an Apology for my reviving this disquisition to that Noble Company of English Philosophers, I shall only imitate the words of the forementioned Abbe's Letter. Pour la Raison de cette Apparence, & de la tromperie de nos Sens, je la tiens plus Difficile a trouver, que les plus grands Equations d'Algebre, & quand vous y aurez bien pense, vous m'Obligerez de m' en dire vostre Sentement, &c.

After which I have only to fubscribe my felf an unworthy Member, and an humble Servant and Admirer of that Illustrious Company.

Dublin March 10. 85.

William Molyneux.

The

The Sentiments of the Reverend and Learned Dr. John Wallis R. S. Soc. upon the aforefaid Appearance, communicated in a Letter to the Publisher.

A S to the laft Inquiry (concerning which, you fay, the Royal Society would be glad to know my Opinion;) about the apparent Magnitude of the Sun near the Horizon, greater than when confiderably high:

The Inquiry is Ancient: And, I remember, I difcourfed it near forty Years ago with Mr. Foster, then Professor of Astronomy in Gresham College. Who did then assure me (from his own Observation, I suppose, for I have never examined it my felf,) that the apparent Magnitude taken by Instrument (however the Fancy may apprehend it) is not greater at the Horizon, than when higher. And Mr. Casmel (when your Letter was communicated to our company here) assirted the fame.

And (though I have not my felf made the Obfervation) I do not doubt but the thing is fo. For it is agreed, That Refraction near the Horizon, though (as to appearance) it alter the Altitude of the thing feen; yet it alters not the Azimuth at all.

T 2

And

And it must needs be fo. For, fince this equally respects all points of the Horizon; let the Refraction be what it will, the whole Horizon can be but a Circle: So that there is no room for the breadth of a thing (as to the Angle at the Eye) to be made greater, whatever its Tallness may (the Refraction not equally affecting all parts in the Circles of Aluitude.) Nor is there any reason why this should rather thrust the other, than that the other thrust this, out of place.

Whereas, in the Altitude, it is otherwife: For while what is near the Horizon is inlarged, that which is further off is thereby contracted: which as to the Azimuth or Horizontal Polition cannot be.

In Spectacles indeed it is otherwife; for they reprefent the Object every way enlarged; and do thereby hide the adjacent parts. But in Refraction by Vapours, fuppoling all parts of the *Horizon* equally affected by them, one part cannot be expanded in breadth (whatever it may be as to the heighth) without thrufting out another (for the whole *Horizon* can be but a Circle) and, why one part rather than another ?

Unlefs we would fay (as perhaps we may, if there shall appear a necessity for it) That the Rays of a lucid Body do expand themfelves every way to the prejudice of the parts adjacent, by covering them.

But fuppofing (which I am apt to believe, till the contrary fhall be evinced by Experiment) that the Sun or Moon's apparent Diameter taken by Inftrument near the Horizon, is the fame as taken in a higher Pofition, (I mean

philed by Google

mean its Horizontal Diameter, or that parallel to the Horizon; for the erect Diameter, in a Circle Perpendicular to the Horizon, may by the Refraction be varied, and thereby made, not greater, but lefs than when higher; as hath been noted in the Name of Sol Ellipticus at the Horizon.) Supposing, I fay, that the Sun's apparent Diameter Horizontal, taken by Instrument, is the fame near the Horizon, as in a higher Position, I take its Imaginary greatness which is fansied near Horizon, to be only a deception of the Eye; or rather the Imagination from the Eye.

For fure it is, that the Imagination doth not estimate the greatness of the Object seen, only by the Angle which it makes at the Eye; but, by this compared with the supposed distance.

True it is that, *Cateris paribus*, we judge that to be the greater Object, which makes at the Eye the greater Angle: But not fo if apprehended at different Diftances.

For if through a Cafement (or leffer aperture) we fee a Houfe at 100 Yards diftance; this Houfe (though feen under a lefs Angel) doth not to us feem lefs than the Cafement through which we fee it, (or this greater than that, becaufe it makes at the Eye the greater Angle:) But the Imagination makes a comparative Eftimate from the Angle and Diftance jointly confidered.

So that, if two things feen under the fame or equal Angles, if to one of them there be ought which gives the apprehension of a greater Distance, that to the Imagination will appear greater.

T 3

Now

Now fure it is, that one great advantage for Estimating of a thing seen, is, from the variety of intermediate Objects between the Eye and the thing seen. For then the Imagination must allow room for all these things.

Hence it is that if we fee a thing over two Hills, between which there lies a great Valley unfeen, it will appear much nearer than if we fee the Valley alfo: And it will appear as just beyond the first Hill. And if we move forward to the top of the nearest Hill (that fo the Valley may be feen) it will then appear much further than before it did.

And on this account it is, that the Sun fetting, appears to us as if it were but juft beyond the utmost of our visible Horizon; because all between that and the Sun is not seen. And, upon the same account, the Heaven it felf seems Contiguous to the visible Horizon.

Now when the Sun or Moon is near the Horizon, there is a profpect of Hills, and Vallies, and Plains and Woods, and Rivers, and variety of Fields, and Inclosures, between it and us : which prefent to our Imagination a great Diftance capable of receiving all these. Or, if it so chance that (in some Position) these Intermediates are not actually seen : Yet having been accustomed to see them, the Memory suggests to us a view as large as is the visible Horizon.

But when the Sun or Moon is in a higher Polition; we fee nothing between us and them (unlefs perhaps fome Clouds) and therefore nothing to prefent to our Imagination fo great a Diftance as the other is. And

Digitized by Google

And therefore, though both be feen under the fame Angle, they do not appear (to the Imagination) of the fame bignefs, becaufe not both fanfied at the fame Diftances: But that near the *Horizon* is judged bigger (becaufe fuppofed farther off) than the fame when at a greater *Altitude*.

'Tis true, that as to fmall and middling Diftances (befides this Effimate from Intermediates) the Eye hath a means within it felf to make fome Effimate of the Diftance. As, when we already know the bignefs of a thing feen, to which we have been accuftomed; as a Man, a Tree, a Houfe or the like: If fuch thing appear to us under a fmall Angle, and indiftinct, and faintly coloured; the Imagination doth allow fuch Diftance, as to make fuch a thing fo to appear. And, if this, thro' a Profpective Glafs, be repefented to us under a bigger Angle, and more diffinct: It is accordingly apprehended as fo much nearer.

But the cafe is otherwife, when we do not by the known bignefs, judge the Diftance; but, by the fuppofed Diftance, judge of the bignefs; as in the Cafe before us.

And accordingly, different Perfons, according to different fancied Diftances, judge very differently. As, if two Stars be fhewed to ignorant Perfons, and you ask how far they feem to be afunder: one perhaps will fay a Foot; another a Yard, or more: And one fhall fay, the Sun appears to him as big as a Bulhel; another, as big as a *Holland* Cheefe: Each effimating according to the fancied Diftance.

279

Again ; in our two Eyes (when the Object is feen by both) there is yet another means of estimating how far off it is. (And it is this by which we judge of Distances.) Namely, there are, from the fame Object, two different visual Cones, terminated at the two Eyes: Whose two Axes contain, at the Object, different Angles, according to different Distances: An accuter Angle at a great Distance, and more obtuse when nearer.

Now, that fuch Object may be feen by both Eyes, clearly; it is requifite that the Eyes be put into fuch a Polition, as that the Sight of each Eye receive the refpective Axe at right Angles. Which requires a different Polition of the two Eyes, according to the different Diftance of the Object.

As will manifeftly appear; if we look, with attention, on a Finger (or other fmall Object) at two or three Inches diftance from the Eye; and then upon another like Object at three or four Yards beyond it: (and this alternately feveral times. For 'twill be manifeft, that while we look intently on the one, we do not fee the other (or but confufedly) though both be juft before us. And, as we change our view, from the one to the other, we manifeftly feel a Motion of the Eyes (by their Mufcles) from one pofture to another.

And according to the different pofture in the Eyes, requisite to a clear Vision by both, we estimate the Distance of the Object from us.

And hence it is, that they who have lost the Sight of onc Eye, are at a great difadvantage,

35

280

as to estimating Distances, from what they could do while they had the use of both.

But now when the Diftance grows fo great, as that the Polition of these visual Axes become Parallel, or fo near to Parallel, as not to be diffinguishable from it: This advantage is lost, and we can thenceforth only conclude, that it is far off; but not how far.

Hence it is, that our view can make no diffinction of the Moon's Diffance, from that of the other Planets, or even of the fixed Stars: But they feem to us as equally remote from us; though we otherwife know their Diffances from us to be vaftly different. Becaufe the Parallax (as I may fo call it) from the different Polition of the two Eyes, isquite loft, and undifcernable, in Diffances much lefs than the leaft of thefe.

And fo, of the fixed Stars among themfelves : Which, though they feem equally remote from us; many (for ought we know) be at Diftances vaftly different. Nor can we tell, which of them is neareft : (unlefs perhaps we may reafonably guefs, those to be neareft, which feem biggeft.) Becaufe, here not only the Parallax from the Diftance of the two Eyes; and that from the Earths Semidiameter; but even that from the Semidiameter of the Earths great Orb, is quite loft; and none remaining, whereby to effimate their Diftance from us.

But (to return to our cafe in hand;) tho' as to finall Diftances, we may make fome eftimate from the known *Magnitude* of the Object: And, as to middling diftances, from the Parallax (as I may call it) arifing from the

the interval of the two Eyes: Yet even this latter will hardly reach beyond, if fo far as the visible *Horizon*: and all beyond it, is lost.

So that, there being nothing left to affift the fancy in effimating fo great a diffance, but only the intermediate Objects: Where thefe intermediates appear to the Eye, (as, when the Sun or Moon are near the Herizon .) the diffance is fancied greater, than where they appear not, (as when farther from it:) and confequently (though both under the fame or equal Angles) that near the Horizon is fancied the greater. And this I judge to be the true reafon of that appearance.

You will excufe (I hope) what excursion I have made; because though fome of them might have been spared, as to the present case; yet they are not impertinent to the bufiness of Vision; and the estimate to be thence made, of *Magnitudes* and Distances, by the Imagination.

The Sun's Eclipfe May 1f, was here obferved about  $\frac{1}{2}$  a Digit; between one and two a Clock after noon.

A Demonstration of an Error committed by common Surveyors in comparing of Surveys taken at long Intervals of Time arising from the Variation of the Magnetick Needle, by William Molyneux. Efg; F. R. S.

THE Variation of the Magnetick Needle is fo commonly known, that I need not infift much on the Explication thereof; 'tis certain that the true Solar Meridian, and the Meridian fhewn by an Needle, agree but in a very few places of the World; and this too, but for a little time (if a moment) together. The Difference between the true Meridian and Magnetick Meridian perpetually varying and changing in all Places and at all Times; fometimes to the Eaftward, and fometimes to the Weftward.

On which account 'tis impossible to compare two Surveys of the fame place, taken at distant times, by Magnetick Instruments, (fuch as the Circumferentor, by which the Down Survey, or Sir William Petty's Survey of Ireland was taken) without due allowance be made for this Variation. To which purpose we ought to know the Difference between the Magnetick Meridian and true Meridian

at

at that time of the *Down* Survey, and the faid difference at the time when we make a new Survey to compare with the *Down* Survey.

But here I would not be underflocd as if I proposed hereby to shew, that a Map of the fame place, taken by Magnetick Instruments at never so distant times, should not at one time give the fame Figure and Contents as at another time This certainly it will do most exactly, the variation of the Needle having nothing to do either in the Shape or Contents of the Survey. All that is affected thereby, is, the Bearings of the Lines run by the Chain, and the Boundaries between Neighbours. And how this may cause a confiderable Error (unless due allowance be made for it) is what I shall prove most fully.

In order to which, let us fuppofe that about the Year 1657. (at which time the *Down* Survey was taken) the Magnetick Meridian and true Meridian did agree at *Dublin*, or pretty nigh all over *Ireland*; that is to fay, that there was no Variation. And indeed by Experiment it was at that time found, as I am well affur'd, that at *Dublin* it was hardly half a Degree.

Let us fuppofe that in the year, 1695. the Variation was 7 Degrees from the North to the Weftward; that it was really fo, I balieve I am pretty well affured, from an Experiment thereof made by my felf with all diligence. But this is not material, let us now only fuppofe it.

Let

285

Let AB represent the Survey of two Town-Lands, one in the possession of A, and t'other in the possession of B, which we call A Town-Land and B Town-Land, taken by the *Down*-Survey, Anno 1657. when there was no Variation.

Let the Line NS running through the Point P be the true Meridian, and confequently the Magnetick Meridian alfo at that time, becaufe of the fuppofed no Variation, and let this Line NS be alfo the Boundary between the two Town-Lands A and B.

In the year 1695, when the Variation is 7 Degrees from the North to the Westward, B having a Map of the Down Survey, and being fuspicious that his Neighbour A had incroached on him by a Ditch PQ, imploys a Surveyor to inquire into the matter: The Surveyor finds by his Map that the Boundary between B and his Neighbour A run from the Point P through a Meadow directly according to the Magnetick Meridian SPN; but obferving the Ditch P Q caft up much to the Eastward of the prefent Magnetick Meridian, he concludes that A has incroached on  $B_{1}$ , and that the Ditch ought to have been caft up along it the Line Pq, the Angle Q Pqbeing an Angle of 7 Degrees, that is the prefent Variation of the Needle; and the Line Pq the prefent Magnetick Meridian: For which Variation, not making any al-lowance, he politively determines that B has all the Land in the Triangle Q P q, more than he ought to have; and that his Ditch ought to run along  $\mathfrak{l}$  the Line Pq.

'Tis

#### 286

Miscellanea Curiosa.

Tis true indeed, if the Surveyor go the whole furround of the Lands A and B, he will find their Figure and Contents exactly agreeable to the Map here expressed. But then the Bearings of the Lines are all 7 Deprees different from the Bearings in the Map. and they will run in and out upon the adjacent Neighbouring Lands, and caufe endlefs Differences between their Possessions; as is manifest from the Figure : wherein the prickt Lines represent the Disagreement in the Bearings of the Lines, protracted from the Point P; and we fee A incroaching on his Neighbours on the Weftward, as he incroaches on B, and B's Eaftward Neighbours incroaching on him, and fo forward and clear tound. Whereas, by a due allowance for the Variation of the Needle, all this Confusion and Difagreement is avoided, and every thing hits right.

Thus for inftance in the Cafe before us, knowing that the Magnetick Variation has caufed the prefent Magnetick Meridian to fail in the Line n q P s, 7 Degrees from the North to the Weftward; to reduce this to the Magnetick Meridian at the time of the Down Survey, I must make the Meridian of mv Map to fall 7 Degrees to the Eastward of my Magnetick Meridian; as we see the Line P Q falls 7 Degrees to the Eastward of the Line P q.

What is here faid on fuppolition that the Magnet had no Variation at the time of the first Survey taken, and that it had 7 Degrees variation Westward at the time of the accord Survey, may easily be accommodated to

to the fuppofal of any other Variations at the first and second Surveys, Mutatis mutandis, for knowing the Variations we know their Difference; and if we know their Difference, this gives us the Angle Q P q, by which we reduce them to each other.

The best way therefore to make Maps invariable, conftant and everlasting, were for the Surveyors, who use Megnetick Instruments to make always allowance for the Magnetick Variation, and to protract and lay down their Plats by the true Meridian. This the wary Sailer is fully convinced of : and therefore in Steering his Course, he conftantly allows for the prefent Variation, which he observes by the Azimuth Compass, or elfe he would miss his appointed Harbour oftner then he would hit it: For no two Points on the Globe keep the fame Bearing to each other by the Magnetick Meridian for any time together. And though the Variation be flow, yet in a long Courfe, or in times pretty diftant, it may caufe vaft Errors, unlefs allowed for. Thus for inftance, fuppofe in the year 1660. a Sailor had fteered from the Land's end of England to Cape Finifter in Spain, by his Magnetick Compass a direct South Courfe; and that at that time there were no Variation. Afterwards Anno 1700. when there was (fuppofe) 7 Degrees of Variation from the North to the Westward, another Sailor intending to make the fame Paffage, fteers directly the fame Southerly Courfe by his Magnetick Compass: I fay, this last Seaman will be carried far into the Bay of Biscay to the Eastward, and will miss of

of his defired Port by many Leagues; but if in his Courfe he hath allowed for this Variation, and inftead of failing a direct Southerly Courfe by his Compafs, he had fteer'd 7 Degrees from the South to the Weftward, he had hit his Point. Whether thefe be the true Bearings of thefe two Places, it matters not: We go on to the Supposition that they are.

Perhaps it may be objected, That Surveys may be taken without Magnetick Instruments, and that therefore this Error arising from the Megnetick Variation, and Change of the Bearings of Lines, may be avoided. To which I answer, first, That granting a Survey may be taken without Magnetick Instruments, this is nothing against what we have laid down relating to Surveys that are taken with Magnetick Instruments, as the Down Survey actually was, and most Surveys at prefent actually are taken therewith. Secondly, Though a Survey may be taken truly without Megnetick Instruments, fo as to shew the exact Angles and Lines of the Plat, and confequently the true Contents, yet this will not give the true Bearings of the Lines, or fhew my Position in relation to my Neighbours, or the other parts of the Country. This must be supply'd by the Magnet, or fomething equivalent thereto, as finding a true Meridian Line on your Land by Celestial Observations. And I doubt not but the ancient Egyptians, before the difcovery of the Magnet were forced to fome fuch Expedient in their Surveys and Applotments of Lands between Neighbour and Neighbour, after the lnundations of the Nile, which, we are told

0002

289

told, gave the first Original to Geometry and Surveying. Absolute Necessity and Use having introduced these, as Delight and Diversion introduced Astronomy amongst the Chaldeans.

And this brings me to another Objection which may be made against the Instance before laid down : It may be faid, That certainly the Surveyor which B imployed was very ignorant, who would choose to judge of the Line PQ, rather by its bearing than by determining the Point Q, by measuring from H and G. To this I answer, What if both the Points Hand G were vanish'd fince the Down Survey was taken? What if the whole face of the Country were chang'd, fave only the Point P and the Line P Q? How shall the Surveyor then judge of the Line PQ but by its bearing? That this is no extrava-gant Supposition, we have an Example in Egypt above-mentioned, where the Nile lays all flat before it, and fo uniformly covers all with Mud, that there is no diffinction. fuch a Cafe your bearing must certainly help you out, there is no other way.

But I answer secondly, To say that the Surveyor might have determin'd the Point Qby admeasurement from G and H, or any other adjoining noted Points, as from F, K, I, &c. 'tis very true; But then 'tis against our Supposition. I am upon shewing an Error that arises from judging of the Line P Q by Magnetick bearing, and to tell me that this might be avoided by another way, is to say nothing. I my felf shew how it may be avoid-

ea

ed by allowing for the Variation; but still it is an Error, till it be avoided.

But thirdly, if B's Surveyor do not allow for the Variation of the Needle, he will never exactly determin even the Points G, F, H, K, &c. or any other Points in the Plat; but inftead thereof will fall on the Points g, h, f, k:

From what has been laid down, we may fee the abfolute neceffity of allowing for the Variation of the Magnet, in comparing old Surveys with new ones; for want of which great Difputes may arife between neighbouring Proprietors of Lands: and it were to be wish'd that our Honourable and Learned Judges would take this Matter into their Confideration whenever any Business of this kind comes before them. Hitherto an abfolute Acquiescence in the Down Survey, without any of the fore-mention'd Allowance, has been agreed upon as a standing Rule in our Courts of Judicature in Ireland; but that many Men may be injured thereby, I fuppofe is manifest from what foregoes.

I have only this to add, That leaft I be thought herein to ftrike at the Truth or Exactness of the Down Survey, 'tis not at all the intention of this Paper, but rather to confirm it, by fhewing which way Men ought to Examine it truly, and not by the common ways used by them, which rather confound it, and all that claim under it.

#### See the Table Fig. 44.

Dig Red & Googl

Al-

291

Although this Paper was chiefly defigned for the ending of Contefts in the Kingdom of *Ireland* about the interefts of fome of those whose Lands are Neighbouring, and have been furveyed by Magnetick Instruments, yet confidering its universal Use, it was thought it would be very grateful to the Curious to publish it here.

U

A Proposal concerning the Parallax of the fixed Stars, in Reference to the Earths Annual Orb. In several Letters of May the 2d. June 29. and July 20.1693. from Dr. John Wallis to William Molineux Esq;

SIR,

292

Am obliged to you for two Books which you have been pleafed to fend me, that of your Sciothericum Telefcopicum, and that of Dioptricks; which you have performed fo well, that I have not been better fatisfied with any that I have feen of that Subject. I should not fo long have neglected to return my Thanks for them, but that I thought a Letter of bare Thanks to be too empty, unless I had somewhat else to fend with it.

You will, I hope, give me leave (though I have not the opportunity of being perfonally known to you) to fuggeft a Speculation, which hath been in my Thoughts thefe Forty Years or more; but I have not had the opportunity of reducing it to Practice, as being not fo well flored with neceffary Inftruments of that kind, nor much exercifed to Telefcopick Obfervations. And though I have many Years fince fuggefted it to others, yet neither

Data day Google

292

ther have they had leifure of convenience of putting it in Practice.

1008

It is concerning the Parallax of the fixed Stars, as to the Earths Annual Orb.

Galileo complains of it a great while fince (in his Syftema Cofmicum) as a thing not attempted to be observed with fuch diligence as he could wish, and I doubt we have the fame cause of complaining ftill. I know that Dr. Hook and Mr. Flamstead have attempted somewhat that way, but have defisted before they came to any thing of Certainty. What hath been done to that purpose abroad I know not.

Galileo hath fuggested divers things considerable in order to it.

As to the times of Obfervation; That it fhould be when the Sun or Earth are in the Tropicks, or as near thereto as may be: Becaufe at those times, if any, will be the greatest difference observable in their meridional Altitude.

As to the Stars to be obferved, That they fhould be fuch as are as near as may be to the Pole of the Ecliptick: For fuch as are in the Plain of the Ecliptick, or near unto it, though they may be fometime nearer, fometime farther from us, (which might fomewhat alter their apparent Magnitude, if it were fo much as to be obfervable) yet it would little or nothing alter the Parallactick Angle, as *Galilelo* doth there demonftrate.

He notes alfo, that in a business fo nice, the ordinary Instruments of Observation (though pretty large) would be insufficient

U₃

(he doubts) for this purpofe, and doth propofe, that by the fide of fome Edifice or Mountain, at fome Miles diftance, the fetting of fome noted Star (as that of *Lucida Lyra*) might be observed at those different times of the Year, which might be equivalent to an Inftrument whose Radius were fo large.

Which were a good Expedient if Practicable; but I doubt the Denfity of our Atmofphere is fo great, as that it will be hard to difcern a Star just at the Horizon, or even within fome few Degrees of it: And that the Refraction would be there fo great, and fo uncertain, as not to comply with fo curious an Obfervation.

That which occurred to my Thoughts upon thefe Confiderations, was to this purpole : That fome Circumpolar Stars (nearer to the Pole of the Equator than is your Zenith, and not far from the Pole of the Zodiack) should be made choice of for this purpofe. And in cafe the Meridinal Altitude be difcernably different at different times, fo will alfo be their utmost East and West Azimuth, which may be better observed than their Rising or Setting: And this will be not obnoxious to the Refraction, as is the Meridional Altitude; (for though the Refraction do affect the Altitude, yet not the Azimuth at all); and we may here have choice of Stars for the purpofe; which in Obfervations from the bottom of a Well we cannot have; being there confined to those only which pass very near pur Zenith, though very fmall Stars.

I

294

I would then take it for granted, as a thing at leaft very probable, that the fixed Stars are not at all (as was wont to be fuppofed) at the fame diftance from us; but the diftance of fome, vaftly greater than of others; and confequently, though as to the more remote, the Parallax may be undifcernable; it may perhaps be difcernable in those that are nearer to us.

And those we may reasonably guess (tho' we are not fure of it) to be nearess to us, which to us do appear biggess and brightess, as are those of the First and Second Magnitude; and there are at least of the Second Magnitude, pretty many not far from the Pole of the Ecliptick, (as that in particular, in the Shoulder of the lesser): And in case we fail in one, we may try again and again on some other; which may chance to be nearer to us than what we try first. And Stars of this bigness may be discerned by a moderate Telescope, even in the day-time; especially when we know just where to look for them.

The manner of Obfervation I conceive, may be thus: Having first pitched upon the Star we mean to obferve, and having then confidered (which is not hard to do) where fuch Star is to be feen in its greatest East or West Azimuth; it may be then convenient to fix very firm and steadily on fome Tower, Steeple, or other high Edifice (in a convenient fituation) a good Telescopick Object-glass in fuch position, as may be proper for viewing that Star. And at a due distance from it near the Ground, build on purpose (if alrea-U 4 dy

295

296

dy there be not any) fome little Stone Wall, or like Place, on which to fix the Eye-glafs, fo as to answer that Object-glass: And having to adjusted it, as through both to fee that Star in its defired Station, (which may beft be done while the Star is to be feen by Night in fuch lituation, near the time of one of the Solftices), let it be there fixed fo firmly, as not to be disturbed, (and the place fo fecured, as that none come to diforder it), and care be taken fo to defend both the Glasses, as not to be endangered by Wind and Weather. In which contrivance I am beholden to Mr. John Cafwel M. A. of Hartball in Oxford, for his Advice and Affiftance ; with whom I have many Years fince, communicated the whole matter.

This Glass being once fixed (and a Micrometer fitted to it, fo as to have its Threds perpendicular to the Horizon, to avoid any inconvenience which might arife from diversity of Refraction if any be) the Star may then be viewed from time to time (for the following Year or longer) to fee if any change of Azimuth can be obferved.

This I thought fit to recommend to your Confideration, who do fo well understand Telescopes, and the managery of them; not knowing any who is more likely to reduce it to Practice. If you shall think fit to give your felf the trouble of attempting the Experiment, and that it fucceed well, it will be a noble Observation, and worth the Labour: And, if it should miscarry, the charge hope would not be great.

But

But when I fuggeft (as a convenient Star for this purpole) the fhoulder of the leffer Bear (as being the neareft to the Pole of the Zodiack of any Star that is of the first or fecond Magnitude), I do not confine you to that Star; but (without retracting that) fuggest another; namely, the middle Star, in the Tail of the great Bear, which (tho' fomewhat farther from the Pole of the Zodiack) is a brighter Star than the other, and may be nearer to us.

But I do it principally upon this Confideration : namely, That there is adhering to it a very fmall Star, (which the Arabs call Alcor, of which they have a Proverbial faving, when they would defcribe a fharp-fighted Man; That he can discern the Rider on the middle Horfe of the Wayn; and of one who pretends to fee small things but over-look much greater; Vidit Alcor at non Lunam plenam): Which Hevelius in his Obfervations, finds to be diftant from it about 9 Minutes, and 5 or 10 Seconds: So that belides the advantage of difcovering the Parallax of the greater Star, if difcernable. Their difference of Parallax of that and of the leffer Star (being both within the reach of a Micrometer) may do our Work For if that of the greater Star be as well. discernable, but that of the lesser be either not discernable, or less discernable. Their different diftances from each other at different times of the the Year, may, perhaps (without farther Apparatas) be difcerned by a good Telescope of a competent length, furnished with a Micrometer, if carefully

pre-

297

preferved from being difordered in the Intervals of the Observations; and discover at once, both, that there is a Parallax, and that the fixed Stars are at different diffances from us, wherein, that I be not miltaken, my meaning is not, that the Inftrument or Micrometer should be removed for the observing of the leffer Star; but that (when the Azimuth of the greater Star is taken) by a Micrometer (confifting of divers fine Threads parallel and transverse) may (at the fame time) be observed the Distance of the two Stars, each from other, in that Polition (both being at once within the reach of the Micrometer ;) which distance (the Inftrument remaining unmoved) if it be found (at different times of the Year) not to be the fame; this will prove, that there is a different Parallax of these two Stars.

This latter part of the Observation (of their different distances at different times) I fuggest, as more easily practicable though not fo nice as the former. For it may be done I think, without any further Apparatas there than a good Telescope, of ordinary form, furnished with a Micrometer, (this being carefully kept unvaried during the Interval of these Observations. And if this part only of the Observation (without the other) be purfued ; it matters not though the two Obfervations (near the two Solftices) be, one at the Eastern, the other at the Western Azimuth (whereby both may be taken in the Night-time,) for the distance must (at both Azimuths) be the fame, if after observing the Azimuth of the greater Star it be necef-

Ingitized by Google

ceffary to move the Micrometer for meafuring its diftance from Alcor that may be done another Night (and it is not neceffary to be done at one Obfervation) for that diftance, and cannot be difcernably varied in a Night or two.

I shall give you no farther trouble at prefent, but subscribe my felf, Sir,

Yours, &c.

299

# A Difcourfe on this PROBLEM;

Why Bodies diffolved in Menstrua Specifically lighter than themfelves, swim therein?

#### By Mr. WILLIAM MOLYNEUX, of Dublin, Member of the Royal Society.

HE Liberty of Philosophising being now universally granted between all Men, I am fure that a difference in Opinion will be no breach of affection between two intirely Loving Brothers: And therefore I shall take the freedom to propose my own Thoughts in a matter wherein my Brother Mr. Thomas Molyneux hath appeared publickly in the Novelles de la Republique des Letres, Mois d' Aout 1684. Art 4. and Mois de Janvier 1685. Art 7. The Problem proposed is, Why Bodies diffolved float in Liquors lighter than themselves ? as for Example : Mercury diffolved in ftrong Spirit of Nitre fwims therein, tho' each fmall Particle of Mercury, be far heavier than fo much of the Liquor whofe place it occupies. This, fays he, cannot be folved by the prime Law of Hydroftaticks, which is, that a Body which is an equal quantity is heavier than a like quantity of Liquor, finks in that Liquor ;

201

quor; thus a Cubick Inch of Iron being heavier than a Cubick Inch of Aqua-Fortis, and each Particle (how fmall foever) of Iron being heavier than a like Particle of Aqua-Fortis; Iron being put into Aqua-Fortis should fink, and yet we find, that Iron being diffolved in a convenient quantity of Aqua-Fortis floats therein, and does not fall to the Bottom. The Reafon which my Brother gives for this is, That the internal Motion of the Parts of the Liquor, does keep up the Particles of the diffolved Solid, for they being fo every Minute, are movable by the leaft force imaginable, and the Action of the Particles of the Menstruum, is sufficient to drive the Atomes of the diffolved folid Body from place to place; and confequently, notwithstanding their Gravity, they do not fink in the Liquor lighter than themfelves. As a Proof of this in the 7th Article of Janvier 1685. he offers an Experiment known in Chymistry, that a Menstruum over a digesting Fire (as the Chymist speaks) will diffolve a greater quantity of a Body put into it, than when 'tis off the Fire, and if it be taken off the Fire, and fuffered to cool, a great Portion will precipitate of that which was perfectly diffolved. whilst the Menstruum continued hot. For. fays he, the Particles of the Menstruum acquire a more violent agitation by the Fire, and are therefore able to raife and keep up a greater Quantity of the diffolved Body, or hereby they are able to refift a greater Gravity.

It has been objected against this Notion, that the common Experiment of precipitation, tion, by mixing an Alkaly with an Acid feems to contradict this; for thereby the Fluidity of the Menstruum is not taken away, and confequently, the internal Agitation of its Parts is not diminished, and yet thereupon. the Particles of the diffolved Body precipitate all to the Bottom. To this he answers in the forecited Article of January, that all Mixtures of different Liquors introduce in each a different Conformation of Pores, and therefore the Infusion of a new Liquor, drives the infelible Parts of the diffolved Body from their Places, and forces them to ftrike againft each other, and cling together, and fo becoming more big and heavier than formerly, the internal Agitation of the Liquor is no longer able to move and fuftain them. and confequently they fall to the Bottom.

This, as fairly and fhortly as I can propole it, is his Sentiment of this Phænomenon.

But I conceive an other Account may be given of this Appearance, and that the forefaid Law of Hydrostaticks is a little deficient. 'Tis true indeed, if we confider only the fpecifick Gravity of a Liquor, and the specifick Gravity of a folid Particle floating therein, the forementioned Rule is exact; but in finking there is requifite a feparation of the Parts of the Liquor by the finking Body; and there being a natural Inclination in the Parts of all Liquors to Union arifing from an Agreement or Congruity of their Parts, there is a refistance therein to any thing that feparates this Conjunction : Now unlefs a Body have weight enough to overcome this Congruity or Union of Parts, fuch a Body will

202

will float in a Liquor specifically lighter than it felf. But that a heavy Body, as Mercury or Iron may have its Parts reduced to that Minutenefs, that their Gravity or Tendency downwards, is not ftrong enough to feparate the Cohefion or Union of the Parts of a Liquor, will be manifest, if we confider, that the Reliftance made by the Medium to a falling Body, is according to the Superficies of the Body; but as the Body decreafes in Bulk, its Superficies does not proportionably de-crease, thus a Sphere of an Inch Diameter, has not eight times lefs Superficies than a Sphere of two Inches Diameter, tho' it have eight times less Bulk, and confequently paffing through a Medium, as fuppole Air or Water, the Sphere of an Inch Diameter is, proportionably to its Bulk, more refifted, than a Sphere of two Inches Diameter in proportion to its Bulk, and hence it will come to pass, that at last a Body may be reduced to that Minutenefs, that its Gravity preffing downwards (which is according to its Bulk) may be lefs than the refistance of the Medium, which operates on the Surface of the Body; feeing as I faid before, the Surfaces of Bodies do not decrease so fast as their Bulks, these decreasing in a Triplicate, but those in a Duplicate Ratio of the Bodies Diameters.

This Account does not at all oppose the Experiment of a *Menstruum* over the Fire, being able to dillolve or fustain a greater Quantity of a heavy Body; for the Reason, of this, as'tis given by my Brother, does not Contradict my Notion. The Account likewife

But because in the beginning of my Difcourfe, I fay that the forementioned Law of Hydroftaticks is a little defective, I defire to explain my felf a little further in that Point. * In Weights falling through the Air, were Gravity only confider'd, the Proportions of their Descents would be exactly as Galileo has demonstrated; but it is allow'd by all, that the Reliftance of the Air, not being conlider'd in those Demonstrations, they are not Mathematically true in Practife, but that really there is fomething of that proportion hindred by the Airs Refiftance. Now, what is this lefs than to fay, that the Refiftance of the Air takes off fome of the Operation of Gravity, or is able to withstand or oppose part of its Action ? And if fo, what shall we fay were an Iron Sphere let through a Medium of Water? Surely the Proportions of its descents would be much more disturbed herein, as Water is much more Solid and difficult to be feparated or passed through than Air, and confequently we must needs grant, that more of the Operation of Gravity, is taken off or relifted by this Opposition of the Water, than that of the Air. And if fo, furely there may be a certain degree of Gravity, that may be quite taken off by the reuftance of the Water : Were a Pistol Bullet let fall through the Air, it would defcend imperceptibly nigh the Proportions that Galileo has affigned, but were a fingle grain of Sand fo let fall, it would be much hindred ia.

205

in its Courfe, and the half of this Grain would be more obstructed; what shall we then fay of the ten thousandth part, or of a part the ten thousand millioneth of this, and again of the Infinits Subdivisions of that, 'till at last we come to a part that would be wholly refissed, or kept up; such as I conceive the minute Particles of a Body dissolved in a Menstruum?

On this account 'tis, I fay, that the forementioned Principle of *Hydroftaticks* is a little defective; for it confiders not the natural Congruity of the Parts of a Liquor, whereby they defire, as 'twere, to unite and keep together, juft as we fee two drops of Water on a dry Board being brought together do jump and coalefce, and therefore Liquors have an innate power of refifting a certain degree of force that would feparate them; fuch as I fuppofe the degree of Gravity, in the moft minute Particles of a Body diffolved in a *Menftruum*.

The fore-mentioned Rule holds true to the most nice Sense in great Bodies, but in those that are by many Millions of Divisions smaller, it feems to fail.

This, in fhort, is my Conjecture in this matter, which I propole, as my Brother did his, with all fubmiffion imaginable, and thereby to give occafion to others to enquire into the Caufes of this appearance, rather than to publifh my own Sentiments as the undoubted Solution thereof.

But this I must acknowledge, that the internal motion of the parts of a Liquor feems fo very agreeable to truth, and explicates fo X many many Phænomena eafily and plainly, that I would not be thought to deny it. Neither would I be thought wholly to reject my Brothers Solution of this Problem; for certainly that Motion (whatfoever it is) in a Menfruum, which is able to diffolve fuch a folid Body as Iron, that is, which is able to difturb the clofe and ftrong Cohefion of the Parts of Iron, may very well be fuppofed fufficient to diffurb or keep up these parts from refting in the bottom of the Veffel, wherein the folution was made: And certainly no better account can poffibly be given of fuch Solutions, than by supposing such an internal motion in the parts of the Menstruum inlinuating themfelves into the folid Body, and loofening its parts. And tho' it may be objected, that in the parts of Water there may be fupposed as violent an internal motion, as in the parts of Aqua-Fortis, and yet we fee Water will not diffolve Iron as Aqua-Fortis does, and common Bees-Wax is diffurbed by neither of them, I leave the nice enquiry after this Point to others, viz. What kind of Motion and peculiar Conformation of parts is requisite both in the Menstruum and in the diffolved Body, that a Solution may refult from their Commixture.

# Some Reflections on the foregoing Paper by Mr. T. M.

What my Brother has laid down in this Difcourfe, I think does most undeniably evince that the received Law of Hydrostaticks is somewhat defective. For Liquors, tho' they are

206

307

are Fluid yet they are Bodies, and therefore confift of parts united ; which Union, tho' it be eafily deftroy'd, yet of necessity it requires fome degree of Force for the effecting it; nor is it more manifest, if rightly confidered, that a Flint requires Force for the feparation of its parts, than that Fluids do for theirs. But however, I imagine, this Property ought not to be relied upon as the fole Caufe of this appearance, to which my Brother has apply'd it; nay perhaps does not fo much as concur the least in the producing this effect; my Reason in short is this: Whatever is of fusicient Power to raife the minute Particles of a heavy Body in a light Fluid, is certainly a fufficient caufe to keep them in that state : Now my Supposition may give fome account of this, what my Brother fays, never can; for he must necessarily suppose them first raifed; and then he gives the reafon of their not finking: Whereas 'tis not to be queitioned but that that Force which raifed them, is the fame which keeps them from falling to the bottom.

But these Conjectures (for I esteem them' no more) I leave to the Confideration of those that desire to enquire further in this Matter.

X

208

Of the weight of a cubic foot of divers grains, &c. try'd in a Veffel of well-feafon'd Oak, whofe concave was an exact cubit foot. By the direction of the Philofophical Society at Oxford.

T H E following Bodies were poured gently into the Veffel, and those in the 12 first Experiments were weighed in scales turning with 2 ounces, but the last 7 were weigh'd in scales turning with one ounce. The pounds and ounces here mentioned are Avoirdupois.

- 1. A foot of Wheat (worth 6 s. a Bushel) weigh'd of Avoirdupois weight. 47.
- 2. Wheat of the best fort (worth 6 s. 4 d. a Bushel) 48. 4.
- 3. The fame fort of Wheat measured a fecond time. 48. 2.
- Both forts were red Lammas Wheat of the last year.

4. White Oors of the last year. 29. 8. The best fort of Oats were 2 d. in a

Bushel better than these.

- 5. Blue *Peafe* (of the last year) and much worm-eaten. 49. 12.
- 6. White Peafe of the last year but one. 50. 8.

7: Barley

8.

309

	π.	3.
7. Barley of the last year (the best		-
fort fells for 1 s. 6 d. in a Quarter		
more than this.	4.1	2.
8. Malt of the last years Barley, made	41.	£.•
2 Months before.		
	30.	4.
9. Field-Beans of the laft year but one.		
10. Wheaten Meal (unfifted).		0.
11. Rye Meal (unlifted.)	28.	4.
12. Pump Water.	62.	8.
13. Bay Salt.	54.	Ι.
14. White Sea-Salt.	43.	
15. Sand.	85.	
16. Newcastle Coal.	67.	
17. Pit-Coal from Wednesbury, 63.		
but this is very uncertain in the		
filling the Interflices betwixt the		
greater pieces.	6.	- 7
	63.	0.
18. Gravel.	09.	
19. Wood-Ashes.	58.	5.

A further List of the Specifick Gravities of Bodies, being in proportion as the following numbers.

DUmp-water	1000
<b>F</b> ir dry	546
Elm dry	600
Cedar dry	613
Walnut-tree dry	631
Crab-tree meanly dry	765
Ash meanly dry, and of the out-fide	lax
part of the Tree	734
Afhmore dry, but about the Heart	. 845
Maple-tree	755
X 3	Yew

310 Miscellanea Curiosa.	
Yew of a Knot or Root 16 years old	760
Beech meanly dry	854
Oak very dry, almost Worm eaten	753
Oak of the out-fide fappy part, fell'd	a
year fince	- 870
Oak dry, but of a very found clofe to	ex-
ture	929
The fame tried another time	932
Logwood	913
Claret	993
Moil Cyder not clear	1017
Sea-water fetled clear	1028
College plain Ale the fame	1028
Urine	1030
Milk	1031
Box the fame	1031
Redwood the fame	1031
Sack	1033
Beer Vinegar	1034
Pitch	1150
Pit-Coal of Stafford-shire	1240
Speckled wood of Virginia	1313
Lignum Vitæ	1327
Stone-bottle	1777
lvory	1826
Alabafter	1872
Brick	1979
Heddington-stone, the foft lax kind	2029
Burford-stone, an old dry piece	2049
Paving-ftone a hard fort from about Blaidon	1 (F)
Flint	2460
	2542
Glafs of a quart Bottle	2666
Black Italian Marble	2704
White Italian Marble tried twice	2707
111 - 5r	

White

Miscellanea Curiosa.	311
White Italian Marble of another	
of a visibly closer texture	2718
Block-tin	7321
Copper	88+3
Lead	11345
Quick-filver	14019
Quick-filver	13593

The laft Experiment was tried with another quantity of Quick-filver, which had been ufed in Water in the preceding Experiment : However I rather truft the laft, for that I found a fmall miftake (tho' here in the calculation allowed for) in the weight of the Glafs containing the Quick-filver in the trial before.

The Solids here mentioned were examined hydrostatically by weighing them in Air and Water; but the Fluids, by weighing an equal portion of each in a Glass holding about a quart. The numbers flew the proportion of gravity of equal portions of these Bodies; but if of these Bodies we take portions equally heavy, their magnitudes will be reciprocally proportional to their correspondent numbers, e.g. a cubic foot of water is to a cubic foot of Alabaster in gravity as 1000 to 1872; but a pound weight of water, is to a pound weight Albaster in magnitude as 1872 to 1000. So that knowing by the former Table the weight of a cubic foot of Water, and by this, the proportion in gravity betwixt Water and Alabaster, we may by the Rule of Three find the weight of a cubic foot of Alabaster, and fo of any other of these Bodies; or we may know their magnitude by knowing their X 4 gra-

gravity. So that an irregular piece or quantity of thefe Bodies being offered, 'tis but weighing them, and we may know their just magnitude without further trouble.

Observations of the Comparative, Intensive or Specific Gravities of various Bodies. Made by Mr. J. C.

DUmp-water,	1000
Cork,	237
Saffafras Wood,	482
Juniper Wood (dry)	556
Plum-tree, (dry)	663
Maftic,	849
Santalum Citrinum,	809
Santalum album,	1041
Santalum rubrum,	1128
Ebony,	1177
Lignum Rhodium,	1125
Lignum Afphaltum,	1179
Aloes,	1177
Succinum pellucidum,	1065
Succinum pingue,	1087
	1238
Jet, The top part of a Rhinocero's horn,	1242
The top part of an Ox horn,	1840
The (Blade) bone of an Ox,	1656
An human Calculus,	1240
Another Calculus humanus,	1433
	1664
Another Calculus, Brimftone, fuch as commonly fold,	1811
	1720
Borax,	1822
A spotted factitious Marble,	1928
A Gally-Pot,	Oyfter-
	~1

Miscellanea Curiosa.	313
Oyfter-shell,	
Murex-shell,	2092
Lapis manati,	2590
Selenitis,	2270
Wood petrefied in Lough-Neagh,	2322
Onyx-stone,	2341
Turcois-stone,	2510
Englifh Agat	2508
Grammatias lapis,	2512
A Cornelian,	2515
Corallachates,	2568
Talc.	2605
Coral,	2657
Hyacinth (spurious)	2689
Jafper (fpurious)	2631
A pellucid Pibble,	2666
Rock Cryftal,	264 E
Crystallum Disdiaclasticum,	. 2659
A red Pafte,	2704
Lapis Nephriticus,	2842
Lapis Amiantus from Wales,	2894
Lapis Lazuli	2913
An Hone,	3054
Sardachates,	3288
A Granat,	3598
A Golden Marcafite,	3978
A blue Slate with fhining Particles,	4589
A mineral Stone, yielding 1 part	. 3500
160 Metal,	in
The Metal thence extracted,	2650
The (reputed) Silver One of We	8500
The (reputed) Silver Ore of Wales,	7464
The Metal thence extracted, Bifmuth,	11087
Spelter,	9859
Spelter Soder,	7065
Iron of a Key,	8362
non of a fity,	7643
	Steel,

•

•

314 Miscellanea Curiosa.	
Steel,	7852
Caft Brafs,	8100
Wrought Brafs,	8280
Hammer'd Brafs,	8349
A false Guinea,	9075
A true Guinea,	18888
Sterling Silver,	10535
A brafs Half-Crown,	9468
Eletrum, a British Coin,	12071
A Gold Coin of Barbary,	17548
A Gold Medal from Merocco,	18420
A Mentz Gold Ducat,	18261
A Gold Coin of Alexanders,	18893
A Gold Medal of Queen Mary,	19100
A Gold Medal of Queen Elizabeth,	19125
A Medal esteem'd to be near fine Gold,	19636

•

ŕ

.

# A Letter of Dr. Wallis to Dr. Sloane, concerning the Generation of Hail, and of Thunder and Lightning, and the Effects thereof.

#### Oxon. July 26. 1697.

SIR,

Thank you for the Transactions of June which you fent me; wherein I am well pleafed with Mr. Halley's Remarks on the Torricellian Experiment at the top of Snowdon-hill in Wales, at the height of 1240 yards perpendicular. Where the height of that Quickfilver in the Barofcope was 3 Inches and to lefs than below at the Sea-fide; which is an Obfervation of good use; and would have been more fo, had he had the leifure to make like Obfervations at feveral other perpendicular heights in the Afcent. For from fuch comparative Obfervations we are to make an Estimate, at what proportion the height of the Quickfilver doth decreafe in reference to the height of the place. I mean whether in the fame Proportion, or the Duplicate, Sub-duplicate, or how otherwife Complicate thereof. From whence we may make a Judgment of the height of the Atmofphere, if at least it have a determinate height. I did once attempt (a great while fince) a Computation of it; but wanted a fufficient number of Data to proceed upon. But

But that which is most furprizing in those Transactions is, the prodigious Hails there mentioned; which happen'd at many Places, on different Days, and all within the compass of lefs than fix Weeks. I have been told of the like in other Places about the fame time, in Lincolnshire, Hampshire, and elfewhere; whether or no on the fame Days which you mention, I cannot tell; nor can I give a pareicular Account of them. But it would be kind in those who can, to give you like Accounts thereof with those you have Published, for a like publick Information.

I find it is thought very ftrange, what should cause so fudden a Congelation of Hailftones to fo great a bignefs before they fell. And it is indeed very strange. But it is not neceffary that the whole bignefs be attained before they begin to fall, but the freezing may continue during the Fall, to increase the Bulk. For I remember that (many Years fince) I observed here at Oxford a strange shower of Hail, wherein (besides the formed Stones that fell on the Ground, there did hang on the Trees a great deal in the Form of Icicles (a Foot or more in length) fo many and heavy, as to break off fome Boughs with their weight; and I was then told, that in fome places great Branches of Trees were fo broken off; which must needs be from the continuing to freeze during the fall.

And truly the Generation of Hail in general, is a thing which deferves to be farther inquired into, than (I think) hath been yet done. I find Mr. Halley (in his Narration) afcri-

Dipleted & Google

217

ascribing it to Vapour disposing the Aqueous Parts so to congeal. And not unlikely.

If I may interpofe my Opinion, you may take it thus:

Thunder and Lightning are fo very like the Effects of fired Gun-powder, that we may reafonably judge them to proceed from like Caufes. The violent Exploiton of Gunpowder, attended with the Noife and Flafh, is fo like that of Thunder and Lightning, as if they differed only as Natural and Artificial; as if Thunder and Lightning were a kind of natural Gun-powder, and this a kind of artificial Thunder and Lightning.

Now the principal Ingredients in Gunpowder are, Nitre and Sulphur (the Admiftion of Charcole being chiefly to keep the Parts feparate for the better kindling of it.) So that if we fuppofe in the Air, a convenient mixture of Nitrous and Sulphorous Vapours, and those by Accident to take Fire; fuch Explosion may well follow, with fuch Noife and Light, as in the firing of Gun-powder. And being once kindled, it will run on from Place to Place as the Vapour leads it, as in a Train of Gun-powder, with like Effects.

This Explosion, if high in the Air, and far from us, willdo no Mischief, or not confiderable; like a parcel of Gun-powder fired in the open Air, where is nothing near to be hurt by it: But if near, to us (or among us) it may kill Men or Cattle, tear Trees, fire Gunpowder, break Houses, or the like; as Gun-powder would do in like Circumstances.

Now

Now this nearnefs or farnefs may be eftimated by the Diftance of Time between feeing the Flash of Lightning, and hearing the Noise of the Thunder. For though in their Generation, they be fimultaneous; yet (Light moving faster than Sound) they come to us fucceffively. I have observ'd that, commonly, the Noife is about Seven or Eight Seconds after the Flash (that is, about half a quarter of a Minute); but fometimes much fooner, in a Second or Two, or lefs than fo, and almost immediately upon the Flash. And at fuch time, the Explosion must needs be very near us. or even amongst us. And, in fuch Cafes, I have (more than once) prefaged the Expectation of Mischief, and it hath proved accordingly, in the Destruction of Men or Cattel. and the like. (As once at Oxford; when, within half an Hour after fuch Prefage, I heard of one killed at Medly, hard by, and others endangered; and another time at Towcester, when within a few Hours after, we heard of Five Perfons kill'd at Everton, about Four or Five Miles from us. and others wounded; beside other Hurt done.)

Now, that there is in Lightning a Sulphorous Vapour, is manifelt from the Sulphorons Smell which attends it, efpecially when Hurt is done; and even where no Hurt is done, from the Lightning it felf, more or lefs difcernable. And a fultry Heat in the Air, is commonly a Fore-runner of Lightning foon after.

And that there is alfo a Nitrous Vapour with it, we may reafonably judge, becaufe we do not know of any Body fo liable to a fuddain and violent Explosion.

Now

Now thefe Materials being admitted, it remains to be confidered, how they may be kindled in order to fuch Explosion. As to which, I have been told from Chymists (though I have not feen it tried) That a Mixture of Sulphur, Filings of Steel, with the Admission of a little Water, will not only cause a great Effervescence, but will of it felf break forth into an actual Fire.

So that there wants only fome Chalybeat or Vitriolick Vapour (or fomewhat equivalent) to produce the whole Effect (there being no want of Aqueous Matter in the Clouds.)

And there is no doubt, but that amongst the various *Effluvia* from the Earth, there may be copious Supplies of Matter for fuch Mixtions.

And 'tis known, that Hay, if laid up too Green, will not only heat, but take Fire of it felf.

And while we are difcourfing of this, it may fuggeft fomewhat as to the Generation of Hail which is very oft an attendant of Thunder and Lightning. 'Tis well known, in our artificial Congelations, that a Mixture of Snow and Nitre (or even common Salt) will caufe a prefent and very fuddain Congelation of Water. And the fame in Clouds may caufe that of Hail-ftones. And the rather, becaufe (not only in those prodigioufly great, but in common Hail-ftones) there feems fomewhat like Snow rather than Ice, in the midft of them.

And, as to those in Particular (of which we are now speaking) fo very large (as to weigh Half a Pound, or Three Quarters of a Pound) fup-

Conditional by Google

220

fuppofing them to fall from fo great a Height, as 'tis manifeft they did by the Violence of their Fall: 'Tis very possible, that though their first Concretion, upon their fuddain Congelation, might be but moderately great, as in other Hail; yet, in their long Descent, if the Medium through which they fall were alike inclined to Congelation, they might receive a great Accession to their Bulk, and divers of them incorporate into one Like as in those Icicles before mentioned.

These have been my Thoughts, occasioned by the Confideration of the furprizing Greatness of these Hail-stones, with the great Thunder and Lightning which did attend these Storms.

Yours, &c.

THE

# [1]

## Á

# SYNOPSIS

#### OF THE

# Aftronomy of Comets.

HE ancient Egyptians and Chaldeans (if we may credit Diodorus Siculus) by a long Course of Observations. were able to predict the Apparitions of Comets. But fince they are also faid, by the Help of the fame Arts, to have prognosticated Earthquakes and Tempests, 'tis past all Doubt, that their Knowledge in these Matters, was the Refult rather of meer Aftrological Cal. culation, than of any Aftronomical Theories of the Cœlestial Motions. And the Greeks, who were the Conquerors of both those People, scarce found any other fort of Learning amongst them, than this. So that 'tis to the Greeks themfelves as the Inventors (and especially to the Great Hipparchus) that we owe this Afronomy, which is now improv'd to fuch a Heigth. But yet, amongst these, the Opinion of Aristorle (who wou'd have Comets to be nothing elfe, but Sublunary Vapours, or Airy Meteors) pre-

T,

prevail'd fo far, that this most difficult Part of the Astronomical Science lay altogether neglected; for no Body thought it worth while to take Notice of, or write about, the Wandring uncertain Motions of what they esteemed Vapours floating in the *Æther*; whence it came to pass, that nothing certain, concerning the Motion of Comets, can be found transmitted from them to us.

But Seneca the Philosopher, having confider'd the Phanoinena of Two remarkable Comets of his Time, made no Scruple to place them amongst the Calefial Bodies; believing them to be Stars of equal Duration with the World tho' he owns their Motions to be govern'd by Laws not as then known or found out. And at laft (which was no untrue or vain Prediction) he foretells, that there fhould be Ages fometime hereafter, to whom Time and Diligence fhou'd unfold all these Mysteries, and who shou'd wonder that the Ancients cou'd be ignorant of them, after fome lucky Interpreter of Nature had fnewn, in what Parts of the Heavens the Comets wander'd, what, and how great they were: Yet almost all the Astronomers differ'd from this Opinion of Seneca; neither did Seneca himfelf think fit to set down those Phanomena of the Motion, by which he was enabled to maintain his Opinion : Nor the Times of those Appearances, which might be of use to Posterity, in order to the Determining these Things. And indeed, upon the Turning over very many Histories of Comets, I find nothing at all that can be of Service in this Affair, before, A. D. 1337. at which time Nicephorus Gregoras, a Con= stantinopolitan Historian and Astronomer, did pretty

Dhiesdo, Google

2

pretty accurately defcribe the Path of a Comet amongst the Fix'd Stars, but was too laxe as to the Account of the Time; fo that this most doubtful and uncertain Comet, only deferves to be inferted in our Catalogue, for the fake of its appearing near 400 Years ago.

Then the next of our Comets was in the Year t472, which being the fwifteft of all, and neareft to the Earth, was observed by *Regiomantanus*. This Comet (so frightful upon the Account both of the Magnitude of its Body, and the Tail) mov'd Forty Degrees of a great Circle in the Heavens, in the Space of one Day; and was the first, of which any proper Observations are come down to us. But all those that consider'd Comets, until the Time of *Ticho* Brabe (that great Restorer of Astronomy) believ'd them to be below the Moon, and so took but little Notice of them, reckoning them no other than Vapours.

But in the Year 1577, (Ticho fericufly purfuing the Study of the Stars, and having gotten large Instruments for the Performing Cœlestial Mensurations, with far greater Care and Certainty, than the Ancients cou'd ever hope for) there appear'd a very remarkable Comet ; to the Observation of which, Tiche vigoroufly applied himfelf; and found by many just and faithful Trials, that it had not a Diurnal Parallax that was at all perceptible: And confequently was not only no Aireal Vapour, but also much higher than the Moon ; nay, might be plac'd amongst the Planets for any thing that appear'd to the Contrary; the cavilling Opposition made by fome of the A 2 .. School-

4

School-men in the mean time, being to no Purpofe.

Next to Ticho, came the Sagacious Kepler. He having the Advantage of Ticho's Labours and Observations, found out the true Physical System of the World, and vastly improv'd the Astronomical Science.

For he demonstrated that all the Planets perform their Revolutions in Elliptick Orbits, whole Plains pals thro' the Center of the Sun, observing this Law, That the Area's (of the Elliptick Sectors, taken at the Center of the Sun, which he proved to be in the common Focus of these Ellips) are always proportional to the Times, in which the correspondent Elliptical Arches are describ'd. He discover'd alfo, That the Distances of the Planets from the Sun are in the Selquialtera Ratio of the Periodical Times, or (which is all one) That the Cubes of the Distances are as the Squares of the Times. This great Aftronomer had the Opportunity of observing Two Comets, one of which was a very remarkable one. And from the Observations of these (which afforded fufficient Indications of an Annual Parallas) he concluded, That the Comets mov'd freely thro' the Flanetary Orbs, with a Motion not much different from a Rectilinear one; but of what Kind he could not then precifely determine. Next, Hevelius (a Noble Emulator of Ticho Brahe) following in Kepler's Steps, embraced the fame Hypothesis of the Rectilinear Motion of Comets, himfelf accurately observing many of them. Yet, he complain'd, that his Calculations did not perfectly agree to the Matter of Fact in the Heavens: And was aware, that the Path of a Connet was bent into a Curve Line towards the SHT .

Sun. At length, came that prodigious Comet of the Year 1680 which defcending (as it were) from an infinite Diflance Perpendicularly towards the Sun, arole from him again with as great a Velocity.

This Comet, (which was seen for Four Months continually) by the very remarkable and peculiar Curvity of its Orbit (above all others) gave the fittest Occasion for investigating the Theory of the Motion. And the Royal Observatories at Paris and Greenwich having been for some time founded, and committed to the Care of most excellent Astronomers, the apparent Motion of this Comet was most accurately (perhaps as far as Humane Skill cou'd go) obferv'd by Mrs. Cassin and Flamssteed.

Not long after, that Great Geometrician, the Illustrious Newton, writing his Mathematical Principles of Natural Philosophy, demonstrated not only that what Kepler had found, did neceffarily obtain in the Planetary System; but alfo, that all the Phanomena of Comets wou'd naturally follow from the fame Principles; which he abundantly illustrated by the Example of the aforefaid Comet of the Year 1680. fhewing, at the fame time, a Method of Delineating the Orbits of Comets Geometrically; wherein he (not without the highest Admiration of all Men) folv'd a Problem, whofe Intricacy render'd it worthy of himfelf. This Comet he prov'd to move round the Sun in a Parabo-"lical Orb, and to describe Area's (taken at the Center of the Sun) proportional to the Times.

Where-

6

Wherefore (following the Steps of fo Great a Man) I have attempted to bring the fame Method to Arithmetical Calculation; and that with defired Succefs. For, having collected all the Obfervations of Comets I could, I fram'd this Table, the Refult of a prodigious deal of Calculation, which, tho' but fmall in Bulk, will be no unacceptable Prefent to Aftronomers. For these Numbers are capable of Representing all that has been yet observ'd about the Motion of Comets, by the Help only of the following General Table; in the making of which I spar'd no Labour, that it might come forth perfect, as a Thing confecrated to Posterity, and to last as long as Aftronomy it felf.

The Astronomical Elements of the Motions in a Parabolick Orl of all the Comets that have been hitherto duly observed.

0)						
1683 1684 1698	1677 1687 1687	1651	1618	1580	1337	An An
$\begin{array}{c} 1683 \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \$	1425 49 10 14 26 49 10 19 2. 2. 0 18 21.16.3		項 15.30.4c 時 12.12 30 日 16 月 C	125.42. ( 725.52. ( 718.57.20 8 7.42 30	П 24.21. с ур 11 46.20 И 19.25. с П 20.27. с	Ajcend.
83.11. 0 55.48.40 31.21.40	3.22.10 79 03 15 60.56. 0 17.56. 0	C 79.28. C C 32.35.5C C 21.18.30 C 76.05. C	29.40.40 55.12. 0 17. 2. 0 17. 34 0	32. 6.3c 74 32.45 54.40. 0 6 4. c	32.11. C 5.20. C 17.56. C 12.36 C	Orbitz.
П 25.29.30 П 28.52.0 П 17.00.30 И 17.01.15	€ 16.59.30 Q 17.37. 5 +>22.39.30	$ \begin{array}{c} \Pi \ 25.10, \ c \ 79.28, \ c \ \gamma \ 28.18 \ 40 \\ \Pi \ 22.30 \ 3c \ 32.35, \ 5c \ \varpi \ 25.58, \ 40 \\ \Pi \ 21.14, \ c \ 21.18, \ 30 \ \Omega \ 10.41, \ 25 \\ \Pi \ 18.02, \ c \ 76.05, \ c \ \Pi \ 11, \ 54.30 \\ \end{array} $	m 6.54 30 m 18.16. 0 ₩ 2.16. 0 Y 2.14. 0	1. 50 0	C & 7.59. 0 C & 15.33.30 C	eo ibelisa. gr. "
\$6020 96015 32500 69129		84750 44851 1025751 10649	57561 51293 58680 37975	1693 <b>5</b> 8 16955 183 <b>7</b> 5 05698	40606 54273 56700 50910	Peribela ( a Sole.
9.748343 Juli 9 982339 Mui 9.511883 Sept. 9 83966004.	69739 9.843476 Feb. 28059 9.448072 Apr. co61227.787106 Dec. 58328 9.765877 Sept.	84750 9928140 Nov. 2.15.40 44851 9.651772 Fin. 16.23.41 102575 20011044 Nov. 24.11.52 10649 9027309 Apr. 14. 5.15	9.700882 ff.n. 9.710058 ffuli 9.7684900a. 9.794980a.	2.666424 Apr. 21.20. 3 9.263447 APr. 26 18.45 9.775450 NJO.28 15.00 2.038850 S-pt. 27 19.20	9.609230 June 2. 6.2 9.734584 Feb. 28 22.2 9.734583 Aug. 24 21.1 9.753883 Aug. 24 21.1 9.768803 Or. 19 22.1	l'eribelia à Sole.
	Feb. 2 Apr. 2 Dec. Sept.	Nov. Jan. 1 Nov. 2 Abr.	Juli 3 Juli 3 Oa. 1	Apr. 2 087. 2 Nov. 2 S-pt. 2	June Feb. 2 Aug. 2 Oft. 1	Jes p Per
3. 2 50 29.10.16 6.14.33 8.16.47	20. 8.37 26.c0.37 8.00. 6 4.07 39	2.15.40 16.23.41 24.11.52 14. 5.152	29. 345 31.19.55 16. 350 29.1223		N COUS SI	Perikelii.
87.55.30 kettog 29.23 00 Dirett. 86.25.50 Dirett.	010	59.51.20 Dirett. 33.28.10 Dirett. 49.27.25 Retrog 1.66. 7.30 Retrog	51.23.50 Retrog 83.56.30 Retrog 108.05. 0 Retrog 73.47. 0 Direft	20001	2 4 4 12	gr. ,
Dirett.	9.29. o Direct. 9.12. 5 Retrog 9.22.30 Direct. 8.23 45 Retrog.	Direct. Direct. Retrog	Retrog Retrog Direft.	Direa.	o Retrog. o Retrog. o Direct.	

This Table needs little Explication, fince 'tis plain enough from the Titles, what the Numbers mean. Only it may be obferv'd, that the Perihelium Diftances, are estimated in such Parts, as the Middle Distance of the Earth from the Sun, contains 100000. A 4

8

A General Table for Calculating the Motions of Comets in a Parabolic al Orbit.

Med.	Ang. A	Logar.	Med.	Ang. a	Logar.
mor.	eribelio.	pro dift.	mot.	peribelio.	pro dift.
		à Sole.	11		à Sole.
0	gr. ' "		0	gr. "	
1	1.31.40	0.000077	31	12.55.00	0.062400
2	3. 3.15	0.000309	32	14. 3.20	0.065838
3	4.34 43	0 000694	33	45 10.29	0.069319
4	6. 6. c	0.001231	34	10.16.35	0.072839
5	7:37.1	0.001921	35	47.21.30	0.076396
6	9. 7.43	0.002759	36	48.25.3	0.079984
78	10.38. 2	0.003745	1 37	49.28.2	12.083600
8	12. 7.50	10.004876	38	50.30.1	0.087244
9	13.37.1	70.006151	39		8 0 0 9 0 9 1 0
10	15.6.	70.007564	40	52.30.5	50.094596
11	16.34 2	0.009119	41	53.29.4	40.098300
12		40.010798			0.102019
13	19.28.4	7 0.012609	43	55.242	10.1057.52
14	20.54.5	4 3.014550	44	56.20.1	2 0.109490
15	22.20.1	4001660	45	\$7.15.	60.113240
15	23.44.4	4 2.01878:	3 40	5 58. 9.	30.116995
17		2 0.02107		1 59. 2.	40.120756
18	26.31.	8 3.02347	2 48	\$ 59.54.1	10.124518
1 19	27.52.5	10.02596	9   45	60.45.2	5 0.128278
20	29.134	7 02857	0 50	61.35 4	( ).132035
21	120.33.4	10.03126	3 5	62.25.1	40135792
22	31.52.	3: 0.03404	5 5:	2 63.13.5	2 0.139544
23	33.10.	23 0 03691	6 5	3 64. 1.4	00.143291
24	34.27.	12 0.03986	4 5	4 64.48.3	80.147029
25	35.42.	55 0.04289	2 5	5 65.34.9	00.15076
26		41 2.04598		6 66.20	30,15448:
27		2010-04915		7 67 04.	10 0.15819:
28		54 3.05238		8 67.48.	12 0.161890
25	10-35.	22 2.05560	58 5	9 68.31.	50 0.16557
30	11115	4-1205900	0116	69 14.	16 0.16925

Med.	Angul. à	Logar.	Med.	Ang. à	Logar.
mot.	peribelio.	pro dift.	mot.	peribelio	pro dift
		â Sole.			à Sole.
0	gr. ' "		0	gr. ' "	
61	69.55.58	0.172914	91	86.20 24	0.274176
62	70.30.56	0.176557	92	86.46.20	0.277239
63	71.17.10	0.180188	93	87.11.43	0.280284
64	71.56.56	0.183803	94	87.36.45	0.283306
65	72.35.57	0.187404	95	88.01.27	0.286308
66	73.14.15	0.190978	96	88.25.40	0.289293
67	73.51.59	0.194540	97	88.40.48	0.292252
68	74.29. 6	0.198085	98	89.12.22	0.295201
69	75.05.30	0.201614	99	89.36.14	0.298122
70	75.41.35	0.205122	100	90.00.00	0.301030
71	76.16.56	0.208612	102		0 306782
72	76.51.43	0.212080	104	91.20.18	0.312469
73	77.25.57	0.215529	106	92.12.14	0.312409
74	77.59.41	0.218963	108	93.54. 4	0.323587
75	78.32.54	0 222378	110	93.34.52	0 329042
76	79. 5.35	0.225769	112		
77	79.37.45	0.229142	114	94-14 40	0.334424
78	80. 9.23	0.232488	116	05.21.22	0·339736 0·344979
79	30.40.34	0.235809	118	96. 8 22	0 344979
80	81.11.16	0.239127	120	96.44.30	0.355262
81		0.242416	122	07 10 .0	
82	82.11.19	0.245684	124	07 54.17	0.360306
.83	82.40.40	0.248032	126	08. 28 00	0.365284 0.370200
84	83. 9.34	0.252159	128	00.00. 47	0.370200
85	83.38. 4	0.255366	130	99.33.11	0.379842
86		0.258552	132		
87		0.261720	134	100. 4.43	0.304570
88	85. 1. 5	0.264865	136	101. 5.48	0.202860
89	85.27.58	0.264865 0.267989	138	101.35.22	0 208428
190	185.54.27		140	102 4.19	390420

Medi

9

Med.	Ang. a	Logar.	Med	Ang. a	Logar.
	peribelio.	pro. dift.	7508.	per ibelio	pro dift.
		à Sole.			à Sole.
0	gr. ' "		0	gr. '''	-
142	102.32.41	0.407380	204	113.37.25	0 52340
144	103.00.31	0.411784	208	114. 9.52	
146	103.27.47	3.410132	212	114.41.23	0.53588
148	103.54.31	0 420430	216	115.12.02	
150	104.20.43	0.424676	220	115.41.51	.54792
152	104.46.22	0.428866	224	116.10.52	0.55378
154	105.11.33	0.433012	228	116.39. 7	0.55953
156	105.36.16	0.437110	232	117. 6.38	.56519
158	106.00.32	0.441164	236	117.33.27	
160	106.24.23	0.445178	240	117.59.35	0.57623
162	106.47.47	0.449144	244	118:25. 5	0 58161
164	107.10.44	> 453060	248	118.49.57	0 58691
166	107.33.17	0.456936	252	119 14.14	0.59212
168	107.55.27	0 460772	256	119.37.56	0.59725
170	108.17.14	0.464208	260	120. 1. 6	0 60230
172	108.38.37	0.468318	264	120 23 44	0.60727
174	108.59.39	0.472030	268	120.45.52	061217
176	109.20.20	0.475705	272	121. 7.30	0.61699
		0.479340	276	121 28 39	0.62175
180	110 00.40	0.482937	280	121-49.22	0.62643
182	110.20.20	0.486498	284	122. 9 38	0 63105
184	110.39.41	0.190022	288	122.29.28	0.63560
		0.493512	292	122 48 54	
188	111.17.28	0.496965	296		0 64452
190	111.35.55	0.500384	300		0 64889
192	111.54.09	0.503769	310	124 11 40	0.65955
194	112.11.58	0.507121	320	124 54 36	3.669880
	112.29.34		330		067987
198	112.46.55	0.5137:9	340	126.14 44	
200	1113. 4.00	0.516934	1350	126.52 12	2.69897

10

ł

11

Med.	Ang. a	Logar. 1	Med.	Aug. a	Logar.
not.	peribelio.	pro dift	mot.	peribelio.	pro dist.
		à Sole.			à Sole.
0	gr.		0	gr.	
360	127.28. 6	0.708104	820	141.49.24	
270	128. 2.3	0.716970		142.10.00	0.978397
280	128.35.38	0.725000	860	142.29.56	0.985771
390	129. 7.27	0.734006	880	142.49.10	
		0.742186	900		0.100000
410	130. 7.3	0.750160	920		1.006871
420	130.36.	20.757930	940		1.013586
430	131. 3.30	0.765516	-960	144.00.10	1:020155
440	131.30.	20.772918	980	144.10.40	1.026583
		10.780148	1000	144.32.40	1.032876
460	132.20.3	00.787216	1500		1.158188
470	132.44.3	20.794122	2000	152.26.19	1.246058
480	133. 7.5	00.800002	2500		1.313703
490	133.30.2	50.807494	3000		71.368678
		00.813969	3500	157.22.4	1.414974
52	c 134.34.1	80.826522	4000	158.24.3	6 1 454950
54	c 135.14.	00.838600	4500	159.163	6 1 490125
56	0135.51.	28 2.85018.7	5000		2 1.521521
58	0 130.27.	60.861369	5500		5 1.549874
_	and a subscription of the	570.872155	6000		41 57571
		130.882575			1 599460
		580.892649			4 1 62141
		21 0.902401	7500	102.37.	41.64183
00	50,139.1	29 0.911866	8000		1.66092
-		250.921012			
7:	20 139.54	100.929907	9000		12 1 69570
7.	40140.19	. 50.938549	9500		38 1 71166
179	00140.42	560.946951	10000		20 1.72678
17	00141.05	550.955124	10000		02.19796
12	0.1141.20	0.00-0-12	1110000	/ 4).	++ 17903

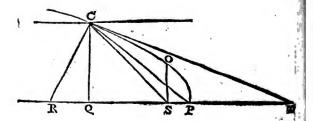
B 2

The

minized by Google

# The Construction and Use of the general Table.

'As the Planets move in Elliptick Orbs, fo do the Comets in Parabolick ones, having the Sun in their common Focus, and describe equal Area's in equal Times. But now because all Parabola's are fimilar to one another, therefore if any determinate Part of the Area of a given Parabola, be divided into any Number of Parts at Liberty, there will be a like Division made in all Parabola's, under the fame Angles, and the Distances will be proportional : And confequently this one Table of ours will ferve for all Comets. Now, the Manner of the Calculation of this Table is thus: In the Fig.



Let S be the Sun, POC the Orbit of a Comet, P the Perihelion, O the Place where the Comet is 90 gr. diftant from the Sun, C any other Place. Draw the Right Lines CP, CS, and make ST, SR, equal to CS; and then having drawn the Right Lines CR, CT, (whereof the one is a Tangent, and the other a Perpendicular to the Curvé) let fall CQ perpendicular to the Axis PSR.

Now,

12

Now, any Area, as COPS, being given, 'tis requir'd to find the Angle CSP, and the Distance CS. From the Nature of the Parabola RQ is ever = i the Parameter of the Axis and confequently if the Parameter be put = 2, then RQ = 1. Let CQ = z; then PQ shall = zz, and the Parabolick Segment COP=1zzz. But the Triangle CSP will =  $\frac{1}{4}z$ , and fo the Mixtilineal Area COPS  $= \frac{1}{2} z^3 - \frac{1}{4} z = a$ , whence z+3 z=12 a. Wherefore refolving this Cubical Equation, z or the Ordinate CQ will be known. Now, let the Area OPS be propos'd to be divided into 100 Parts; this Area is to of the Square of the Parameter, and confequently 12 a is = that Square = 4. If therefore the Roots of these Equations  $z^3 + 3 z = 0, 04:0.08:$ 0,12: 0, 16, Ge. be fuccessively extracted, there will be obtain'd fo many z or Ordinates CQ respectively, and the Area SOP will be divided into 100 Parts. And in like manner is the Calculus to be continued beyond the Place 0. Now the Root of this Equation (fince RQ is=1) is the Tabular Tangent of the Angle CRQ, or ; the Angle CSP, and fo the Angle CSP is given. And RC, the Secant of the fame Angle CRQ, is a mean Proportional between RQ or Unity, and RT, which is the Double of SC, as is plain from the Conicks. But if SP be put = 1, and fo the Latus Rectum = 4 (as in our Table) then RT will be the Diffance fought, viz. the Double of SC in the former Parabola. After this manner therefore, I compos'd the foregoing Table, which ferves to represent the Motions of all Comets : For hitherto there has been none observ'd, but comes within the Laws of the Parabola.

It

## Mifcellanea Curiofa.

14

It remains now, that we give the Rules for the Calculation, and fhew the Way of determining the Place of a Comet feen, by these Numbers. The Velocity of a Comet moving in a Parabola, is every where to the Velocity of a Planet describing a Circle about the Sun, at the same Distance from the Sun, as V 2 to 1. as appears from Cor. 7. Prop. 16. Lib. 1. of the Princip. Phil. Nat. Math. If therefore a Comet in its Perihelium were suppos'd to be as far distant from the Sun as the Earth is, then the Diurnal Area which the Comet wou'd describe, wou'd be to the Diurnal Area of the Earth, as V2 to I. And confequently, the Time of the Annual Revolution, is to the Time in which fuch a Comet wou'd describe a Quadrant of its Orbit from the Peribelium, as 3.14159, Ge. (that is the Area of the Circle) to  $\sqrt{\frac{8}{9}}$ . Therefore the Comet wou'd describe that Quadrant in 109 Days, 14 Hours, 46 Minutes; and fo that Parabolick Area (Analogous to the Area POS) being divided into 100 Parts, to each Day there wou'd be alotted 0.912280. of those Parts; the Log. of which, viz. 9.960128, is to be kept for continual Use. But then the Times in which a Comet, at a greater or lefs Difance, wou'd describe fimilar Quadrants, are as the Times of the Revolutions in Circles, that is, in the Sefquiplicate Ratio of the Distances: And to the Diurnal Area's, estimated in Centesimal Parts of the Quadrant (which Parts we put for Measures of the mean Motion, like Degrees) are in each, in the Subsefquialtera Proportion of the Distance from the Sun in the Perihelion.

Thefe

These necessary Things premis'd, let it be propos'd to compute the apparent Place of any me of the mention'd Comets, for any Given *lime*. Therefore,

I. Let the Sun's Place be had, and the Log. of ts Diftance from the Earth.

2. Let the Difference between the Time of the Petihelion, and the Time given, be gotten, in Days and Decimal Parts of Days. To the Log. of this Numher, let there be added the constant Log. 9.960128, and the Complement Arithmetical of the \$ of the Log. of the Distance in the Perihelium from the Sun : The Sum will be the Log. of the Mean Motion, to be fought in the first Column of the General Table.

3. With the Mean Motion let there be taken the correspondent Angle from the Perihelium, in the Table, and the Log. for the Distance from the Sun: Then in Comets that are Direct, add, and in Retrograde ones substract; if the Time be after the Perihelium, the Angle thus found, to or from the Place of the Perihelion; or in Direct Comets, substract; and in Retrograde ones, add; if the Time be before the Perihelion, the foresaid Angle to or from the Place of the Perihelion; and fo we shall have the Place of the Comet in its Orbit. And to the Log. found for the Distance, let there be added the Log. of the Distance in the Perihelion, and the Sum will be the Log. of the true Distance of the Comet from the Sun.

4. The Place of the Node, together with the Place of the Comet in its Orbit, being given, let the Diftance of the Comet from the Node be found; then, the Inclination of the Plane being given, there will be given alfo (from the common Rules of Trigonometry) the Comet's Place reduced to the Ecliptick, the Inclination or Heliocentrick Latitude, and the Log. of the curtage Diffance.

5. From

5. From these Things given (by the very same Rules that we find the Planets Places, from the Sun's Place and Distance given) we may obtain the Apparent or Geocentrick Place of the Comer, together with the Apparent Latitude. And this it may be worth while to illustrate by an Example or two.

#### EXAMPLE I.

Let it be requir'd to find the Place of the Comet of the Year 166⁵, March 1^d, 7^h, oo', P. M. London. That is. 96^d, 19^h 81, after the Perihelion, which happen'd Novemb. 24°, 11^h, 52'.

Log. Diff. Perihel.	0. 011044
Log. Selquialt.	0. 016566
Comp. Arith.	9 983434
Tem.L. man	9. 960128
Log. Temp.	1. 985862
Log. Med. Mot.	1. 929424
Medius Motus	85.001
Perihel. SL	10. 41. 25
Ang. Corresp.	83. 38. 05-
Comet. in Orb. 8	17. 3. 20
Afcend. Nod. II	21. 14. 00
Com. à Nodo	34. 10. 40
Red. ad Eclip.	32 19.05
Com Helioc. 8	18 54. 55
Incl. Bor.	11. 46. 50
Log. pro dift.	0. 255369
Log. Perihel.	0. 011044
Co-tin. Incl.	9 990754
Log. dift. Curt.	0. 257167
Log. dift. O	9. 997918
ΘX	21. 44. 45
Com. Visus $\gamma$	29. 18. 30
Lat. Vifa	8. 36. 19
	Ê

EXAM-

Dig and Google

#### EXAMPLE II.

Let it be requir'd to find the Place of the Comet of the Year 1683, July 23°, 13^h, 35', P.M. London: Or, 13^h, 40' Equat. Time. That is, 21^d, 10^h, 50' after the Perihelion.

Log. dift. Perihel. Log. Serquialt.	9. 748343 9. 622514
Comp. Arith.	0. 377486 9. 9 ⁵ 0128
Log. Temp.	1. 310723
Log. Med. Mot.	1. 648337
Medius Motus	44. 498
Perihel. II	25. 29. 30
Ang. Corresp.	56. 47. 20-
Comet. in Orb. $\gamma$	28. 42. 10
Nod. Descend. ¥	23. 23. 00
Com. a Nodo	35. 19. 10
Red. ad Eclip.	4. 48. 30
Com. Helioc. X	28. 11. 30
Incl. Bor.	35. 2.00
Log. pro dift.	0. 111336
Log. Perihel.	9. 748 143
Co-fin. Incl.	9 913187
Log. dift. Curt.	9. 772866
Log. dift. Q.	0. 006104
O Locus &	10. 41. 25
Com. Vilus 5	5. 11. 50
Lat. For.	28. 52.00

At the Inftant of Time fpecified in the first Example, 'twas observ'd (at London) that the Comet applied to the Second Star of Aries; fo that it was found to be 9' more Northerly, C and

Divitized by Google

and 3' to the Eaft, according to Mr. Hook's Obfervation. But at that of the Second Example, I my felf (near London, with the fame Inftruments whereby I formerly obferv'd the Southern Conftellations) found the Place of the Connet to be  $\mathfrak{B}, \mathfrak{5}^{\circ}, \mathfrak{11}'$ ; and  $28^{\circ}, \mathfrak{52}'$  North Latitude, which agreed exactly with the Obfervation made at Greenwich almost the very fame Moment.

As for the Comet of the Year 1680, which came almost to the very Sun it felf (being in its Perihelion, not above ! of the Semi-Diameter of the Sun diftant from the Surface of it) fince the Latus Rectum is fo very fmall, could hardly be contained within the Limits of the General Table, becaufe of the exceflive Velocity of the Mean Motion. Therefore in this Comet, the best Way is (after the Mean Motion is found) to get from thence (by the Help of the foregoing Equation  $z^3 + 3z = \frac{4}{10}$  of the Mean Motion) the Tangent of Half the Angle from the Perihelion, together with the Log. for the Diftance from the Sun. Which Things being given, we are to proceed by the fame Rules, as in the reft.

After this Manner therefore, the Aftronomical Reader may examine thefe Numbers, which I have calculated, with all imaginable Care, from the Obfervations I have met with. And I have not thought fit to make them publick before they have been duly examin'd, and made as accurate as 'twas poffible, by the Study of many Years. I have publifh'd this Specimen of Cometical Aftronomy, as a *Prodromus* of a defigned future Work, left, happening

Dianzed by Google

to die, these Papers might be lost, which every Man is not capable to retrieve, by reafon of the great Difficulty of the Calculation. Now, it may not be amils to put the Reader in mind, That our Five first Comets, (the Third and Fourth observ'd by Peter Apian, the Fifth by Paulus Fabricius) as also the Tenth feen by Mastin, if I miltake not, in the Year 1596. are not fo certain as the reft; for the Obfervations were made neither with fit Inftruments, nor due Care, and upon that Account are difagreeing with themfelves, and can by no means be reconcil'd with a regular Computation. The Comet which appear'd in the Year 1684. was only taken Notice of by Blanchinus, who obferved at Rome : And the last, which appear'd in the Year 1698. was feen only by those at Paris, who determin'd its Course in a very uncommon Way. This Comet was very obfcure; and, altho' it mov'd fwift, and came near enough our Earth; yet we, who are wont to be curious enough in these Matters, faw nothing of it. For want of Observations I have left out of the foregoing Catalogue, those Two remarkable Comets which have appear'd in this our Age, one in November, in the Year 1689 the other in February in the Year 1702. For they directing their Coursestowards the Sonthern Parts of the World, and being fcarce confpicuous here in Europe, met with no Observers capable of the Bulinefs. But, if any one shall bring from India, or the Southern Parts, an accurate Series of requifite Obfervations, I will willingly fall to work again, and undergo the Fatigue of reprefenting their Orbits in Numbers, as I have done the reft.

C 2

By

19

By comparing together the Accounts of the Motions of thele Comets, 'tis apparent, their Orbits are dispos'd in no manner of Order; nor can they, as the Planets are, be comprehended within a Zodiack, but move indifferently every Way, as well Retrograde as Direct; from whence it is clear, they are not carry'd about or mov'd in Vortices. Moreover, the Distances in their Perihelium's are fometimes greater, fometimes lefs; which makes me fuspect, there may be a far greater Number of them, which moving in Regions more remote from the Sun, become very obscure; and wanting Tails, pass by us unfeen:

Hitherto I have confider'd the Orbits of Comets as exactly Parabolick; upon which Supposition it wou'd follow, that Comets being impell'd towards the Sun by a Centripetal Force, defcend as from Spaces infinitely di-. ftant, and by their Falls acquire fuch a Velocity, as that they may again run off into the remotest Parts of the Universe, moving upwards with fuch a perpetual Tendency, as never to return again to the Sun. But fince they appear frequently enough, and fince none of them can be found to move with an Hyperbolick Motion, or a Motion Swifter than what the a Comet might acquire by its Gravity to the Sun, 'tis highly probable they rather move in very Excentrick Orbits, and make their Returns after long Periods of Time: For fo their Number will be determinate, and, perhaps, not fo very great. Belides, the Space between the Sun and the fix'd Stars is fo immenfe, that there is Room enough for a Comet to revolve, tho' the Period of its Revolution be vafely long. NOW.

21

- CONTRACTOR

Now, the Latus Rectum of an Ellipsis, is to the Latus Rectum of a Parabola, which has the fame Distance in its Perihelium; as the Distance in the Aphelium in the Ellipsis, is to the whole Axis of the Ellipfis. And the Velocities are in a Subduplicate Ratio of the fame: Wherefore in very Excentrick Orbits this Ratio comes very near to a Ratio of Equality; and the very fmall Difference which happens on Account of the greater Velocity in the Parabola, is eafily compensated in determining the Situation of the Orbit. The principal Use therefore of this Table of the Elements of their Motions, and that which induced me to conftruct it, is, That whenever a new Comet shall appear, we may be able to know, by comparing together the Elements, whether it be any of those which has appear'd before, and confequently to determine its Period, and the Axis of its Orbit, and to foretell its Return. And, indeed, there are many Things which make me believe that the Comet which Apian observ'd in the Year 1531. was the fame with that which Kepler and Longomontanus took Notice of and describ'd in the Year 1607. and which I my felf have feen return, and observ'd in the Year 1682. All the Elements agree, and nothing feems to contradict thismy Opinion, belides the Inequality of the Periodick Revolutions: Which Inequality is not fo great neither, as that it may not be owing to Phylical Caufes. For the Motion of Saturn is fo dilturbed by the reft of the Planets, efpecially Jupiter, that the Periodick Time of that Planet is uncertain for fome whole Days together. How much more therefore will a Comet be fubject to fuch like Errors, which rifes 1:25 .... A 3 als almost Four times higher than Saturn, and whole Velocity, tho' encrealed but a very little, would be fufficient to change its Orbit, from an Elliptical to a Parabolical one. This, moreover, confirms me in my Opinion of its being the fame; that in the Year 1456. in the Summer time, a Comet was feen paffing Retrograde between the Earth and the Sun, much after the fame Manner : Which, tho' no Body made Observations upon it, yet from its Period, and the Manner of its Transit, I cannot think dif-ferent from those I have just now mention'd. Hence I dare venture to foretell, That it will return again in the Year 1758. And, if it should then return, we shall have no Reason to doubt but the reft must return too : Therefore Altronomers have a large Field to exercife themselves in for many Ages, before they will be able to know the Number of these many and great Bodies revolving about the common Center of the Sun; and reduce their Motions to certain Rules. I thought, indeed, that the Comet which appear'd in the Year 1532. might be the fame with that observ'd by Hevelius in the Year 1661. But Apian's Observations, which are the only ones we have concerning the first of these Comets, are too rude and unskilful, for any thing of Certainty to be drawn from them, in so nice a Matter. I defign to treat of all these Things in a larger Volume, and contribute my utmost for the Promotion of this Part of Aftronomy, if it shall pleafe God to continue my Life and Health.

In the mean time, those that defire to know how to construct Geometrically the Orb of a Comet, by Three accurate Observations given, may

22

may find it at the End of the Third Book of Sir Ifaac Newton's Principles of Natural Philofophy, entituled De Systemate Mundi, in the Words of its renowned Inventor. Which have fince been more fully explained by my very worthy Collegue Dr. Gregory, in his Learned Work of Astronomia Physica & Geometrica.

One Thing more perhaps it may not be improper or unpleasant to advertise the Astronomical Reader; That fome of these Comets have their Nodes fo very near the Annual Orb of the Earth, that if it shall fo happen, that the Earth be found in the Parts of her Orb next the Node of fuch a Comet, whilft the Comet passes by; as the apparent Motion of the Comet will be incredibly fwift, fo its P. rallax will become very fenfible; and the Proportion thereof to that of the Sun will be given. Wherefore fuch Transits of Comets do afford us the very best Means, though they feldom happen, to determine the Dillance of the Sun and Earth: Which hitherto has only been attempted by Mars in his Opposition to the Sun; or elfe Venus in Perigao; whofe Parallaxes though triple to that of the Sun, are fcarce any ways to be perceived by our Inftruments; whence we are still in great Uncertainty in that Affair. This use of Comets was the ingenious Thought of that excellent Geometrician Mr. Nicolas Facio. Now the Comet of 1472, had a Farallax above Twenty times greater than the Sun's. And if the Comet of 1618, had come down, about the Middle of March, to his defcending Node: Or if that of 1684, had arrived a little fooper at its afcending Node; they would have been yet much nearcr the Earth.

22

24

Earth, and confequently have had more notable Parallaxes. But hitherto none has threaten'd the Earth with a nearer Appulse, than that of 1680. For by Calculation I find, that Novemb. 11°, 1h, 6, P. M. that Comet was not above the Semi-diameter of the the Sun to the Northwards of the Way of the Earth. At which Time, had the Earth been there. the Comet would have had a Parallax equal to that of the Moon, as I take it. This is fpoken to Aftronomers: But what might be the Confequences of fo near an Appulfe; or of a Contact; or, laftly, of a Shock of the Cælestial Bodies, (which is by no means impossible to come to pass) I leave to be difcufs'd by the Studious of Phylical Matters.

A set of a set of

readers for the sheat beneficial to the set of the set

# Geometrical Differtation

A

25

Concerning the

# **RAINBOW**:

#### IN.

Which (by a direct Method) is fhewn how to find the Diameter of each Bow, the Proportion of the Refraction being given: Together with the Solution of the Inverse Problem, or how to fend the Ratio of the Refraction, the Diameter of the Iris being given.

#### By EDM. HALLEY, F.R.S.

have particularly defcribed the Rainbow (a Meteor fo remarkable for its fair Colours) and given an Account of the Caufes of it. And the Ancient Mythologift, from its wonderful Forni and Appearance, thought

26

thought fit to give it the Title of Thaumanis, or the Child of Wonder; and placing it in the Number of the Goddeffes, attributed to it the Office of a Meffenger between the Caleftials and mortal Men; which Fable, perhaps, owes its Original to Genefis, Ch. 9. V. 13.

Those that attentively confider'd the Phanomena of the Rainbow, always found, that the Sun's Rays reflected by a Watery Cloud, came to the Eye under a certain Angle; from whence arole the Arch, or Circular Figure of But as for the Caufe and Reafon of the it. Colours, as also of the Magnitude of the Angle, by which we constantly find it distant from the Point opposite ta the Sun; these were Things, that a long while, and very greatly perlex'd, as well the Moderns, as Ancients. Neither did they do any thing to the Purpofe herein, till the Famous Monfieur Des Cartes making use of the Mathematical Sciences. fhew'd by feveral Examples, that more strict and close Methods of Reasoning might and ought to obtain, even in our Management of those Physical Speculations. Amongst other things (tho' it must be own'd that herein he had some Light, from the Learned Antonio de Dominis, Arch-bishop of Spalato) he explain'd the Theory of the Rainbow. And having difcover'd the Laws of Refraction, he clearly demonstrated, that the Primary Iris was nothing elfe, but the Sun's Image reflexed from the Concave Surface of innumerable Spherical Drops of Rain; and that with this Condition, that those Rays that were parallel at their Incidence, were not loft or diffipated by the Reflexion, and the Two Refractions (one at the THE IS IS In-

Dianced by Google

27

Ingrefs, and the other at the Egrefs) but fell (and that also parallel) on the Eye. That the Rays were ting'd with Colours by those Refractions, after the fame manner as we fee they are by a Glass Prisme. That the Secundary Iris is produced, after the fame manner, by the Rays that fall more obliquely, only here are Two Reflexions, before the Sun's Rays (which when refracted a Second time proceed parallel to the Eye) emerge out of the Drops of Wa-Further, that the Magnitude of each Iris ter. depends upon the Degrees of the Refraction, which is different according to the Nature of each transparent Solid or Liquid.

And fuppofing the Proportion of the Sines of the Angles of Incidence to the Sines of the refracted Angles, to be in Water, as 250, to 187, he determin'd the Semi-Diameter of each Iris, agreeably to Obfervations, viz. that of the Primary Iris, 41°. 30'. and that of the Secundary, 51°. 54'. By which he did not fo much confirm the Theory it felf, which was demonstrated from other Principles, as the Truth of the fore-mention'd affumed Proportion, (viz. that of the Refraction.) But for these Things, the Reader may confult the 8th Chapter of Cartes's Meteors, whither I refer him.

But now Cartes (who used an indirect and tentative Method in determining these Angles) did not seem clearly to apprehend the Eatiness of the Problem he had proposed to himself. And because none (that I know of) since him, has handled the same Argument more fully; and also since some have misunderstood what Cartes did, committing very great Paralogisms, in some Books (since his time) which particularly 28

larly pretended to explain the Phanomena of the Rainbow; I was willing to fupply what I thought was wanting in this Doctrine, and from the Proportion of the Refraction given, Geometrically to determine the Angle of its Diflance from the Point opposite to the Sun: Or contrarywife, from the Iris given, to determine the refractive Power of the Liquid.

What the Celebrated Mr. Newton has done upon this Head, the Reader will find (with much greater Advantage) in his Book of Light and Colours, when he shall think fit to beftow those excellent Discoveries upon the Publick.

But to proceed: 'Tis plain from what Cartes has demonstrated, that the Primary Iris is form'd by fuch Rays of the Sun, where the Excels of Two refracted Angles, above one Angle of Incidence, is the Greatest of all such Exceffes poffible. And that the Secondary Iris is form'd by those Rays only, where the Excess of Three refracted Angles, above one of Incidence, is in like manner the Greateft. And fo we may go on to a 3d, 4th, or any other his, which are form'd, where the Rays emerge after 3, 4, or more Reflexions. But these can never be feen in the Heavens, becaufe of the Sun's Light which is ftill more and more debilitated by each Reflexion and Refraction : Whence it comes to paisalfo, that the Secondary Iris, is painted with Colours, fo much fainter than the Primary one But in all thefe the general Rule is, that the Excels of 4, or 5, or more refracted Angles, (viz. the Number of Reflexions being increased by Unity) above one Angle of Incidence, is of all the Greateft 71111 Now

Now this greateft Excels doubled, is always the Diffance of the Iris from the Point opposite to the Sun, when the Number of Reflexions is uneven. But if that Number be even, then the Double of that greateft Angle, is the Diffance of the Iris from the Sun it felf, viz. in the 2d, 4th, 6th, &c. Iris. All these Things are either purely Cartefins's, or elfe eafily follow from his Writings in the foremention'd Place.

But now to obtain those greatest Excesses, having the Refraction of any Liquor given; 'tis to be observ'd, that the Excess of Two refracted Angles, above one of Incidence, is there the Greatest, where the momentaneous Increment of the Angle of Incidence is exactly double of the momentaneous Increment of the refracted Angle. And that the Excess of Three refracted Angles is there the Greatest, where the Increment of the Angle of Incidence is triple the Increment of the refracted Angle; and fo of the rest. And this is fufficiently evident of it felf: But as for the Angles, we may obtain them by the Help of the following Lemma, which must therefore be demonstrated.

eds 10 "

#### LEMMA.

The Legs of any plain Triangle continuing; if the Vertical Angle be augmented or diminish'd, by an Angle less than any Angle assign'd; the Momenta or Instantancous Mutations of the Angles at the Base, are to one another reciprocally, as the Segments of the Base.

At Fig. 1. Plate 3. fuppofe the Triangle ABC, whofe Vertex is A, its Legs AB, AC, and Bafe BC, upon which let fall the Perpendicular AD. Then let the Angle BAC be increafed by the Indivisible Momentum CA c, and let the Lines B c d, c D be drawn, which differ, in Imagination only, from the Lines BCD, CD. I fay, that the Momentum of the Angle ABC (viz. CBc) is to the Momentum of the Angle ACB or ACD, as CD to BD, that is reciprocally as the Segments of the Bafe.

#### DEMONSTRATION.

Because the Angle ACD, is the Sum of the Angles ABC, BAC, its Momentum also shall equal the Sum of the Momenta of those Angles; that is, it shall equal CA c + CBc. But CA c = CDc, since, because of the right Angle at D, the Points A, D, C, c, are all in the Arch of a Circle, whose Diameter is AC: By Eucl. 3.9. And consequently the Sum of the Angles CB c, CD c (that is the Angle D c d) shall be the Momentum of the Angle ACD or ACB.

13307

31

But those Angles CB c, D c d, being indefinitely fmall, are to one another as their opposite fides, that is, as c D or CD to BD, that is as the Segments of the Base reciprocally. Q: E:D.

If each of the Angles B and C be Acute, the Lemma will still (mutatis mutandis) be demonstrated after the same Manner.

#### COROLLARY.

Hence it follows that the Momenta of the Angles at the Base, are to one another directly, as the Tangents of those Angles.

By the Help of this Lemma, I will be easie to find the Diameter of any Iris whatfoever; and that either by Calculation, or a Geometrical Construction. For taking any right Line, as CA (Fig. 2.) let it be divided first of all in D, fo that CA, may be to CD, in the Ratio of the Refraction in Water, which is as 250 to 187, or more accurately, as 529 to 396. Then let CA be divided fo in E, that CE may be to AE, as Unity to the Number of Reflexions, a Ray of the Sun (fit to produce the Iris proposed) undergoes: And upon the Diameter AE describing the Semi-Circle ABE, on the Center C with the Radius CD describe the Arch BD, meeting the Semi-Circle ABE in the Point B. Laftly, Drawing the Right Lines CB, AB, let CF be let fall perpendicular upon AB produced, and EB parallel thereto. I fay then, that CBF is the Angle of Incidence, and CAB the Refracted Angle that we enquire after, and which will produce the Iris propos'd.

DE-

#### DEMONSTRATION.

Becaufe the Triangles ACF, AEB are fimilar, it will be AF: BF:: AC:EC; that is, as the Number of Reflexions encreas'd by Unity to Unity (by the Conftraction) and confequently the Momentum of the Angle CBF, will be to the Momentum of the Angle CAF, in the fame Proportion (by the foregoing Lemma.) But the Sine of the Angle CBF, is to the Sine of the Angle CAF, in the Proportion of the Sides CA, CB, that is, in the Proportion of the Refraction given (allo by the Conftraction.) Therefore CAF is the Refracted Angle, correfponding to the Angle of Incidence CBF; and their Momenta are in the Ratio propos'd, wherefore they are the Angles fought. Q. E. D.

And now, multiplying the Refracted Angle by the Number of the Reflexions encreas'd by Unity, and from the Product fubftracting the Angle of Incidence, we shall have half the Distance of the Iris from the Sun, if the Number of Reflexions be even, or from the Point opposite to the Sun, if that Number be uneven, as we have fhewn already. Hence we may exhibit (by a Construction concise and eloquent enough) the Incidencies of all the Orders of Iris's, in any Liquor whole Refraation is known. For if the Line AC (FIG. 2.) be divided into Two equal Parts at E, into Three equal Parts at e, into Four at e, into Five at n, &c. And on the Diameter AE, Ae, As, An, be describ'd, the Semi-Circles ABE, Abe, Ase, Aun, which are all interfected in the Points B, b, B, v, by the Arch DBbBu, defcrib'd

fcrib'd on the Center C with the Radius CD, which is to AC, in the given Proportion of the Refraction: I fay then that the Lines AB, Ab,  $A\beta$ , Av, will make with the Line AC, the Angles CAB, CAb, CA³, CAv, equal to the Refracted Angles; and with the refpective Rayes CB, Cb, C³, Cv, they will make Angles equal to the Angles of Incidence that are required ; viz. ABC (or rather its complement to a Semicircle) for the Primary Iris, AbC, for the Secondary, ABC, for a Third Iris, AvC, for a Fourth, If any one has a mind, to find these Angles by an accurate Calculation, 'twill follow from the fame Principle, that putting the Radius=I, and the Ratio of the Refraction as r to s, the Sine of Incidence will be  $\sqrt{\frac{4}{3} - \frac{1rr}{3^{55}}}$ , and the Sine of the Refracted Angle  $\sqrt{\frac{4ss}{4rr}}$  from  $\frac{1}{3rr}$   $\frac{1}{3}$ , which Angles proceeds the Primary Iris. For the Secondary the Sine of Incidence will be  $\sqrt{\frac{9}{8} - \frac{1rr}{8ss}}$  and the Sine of the Refracted Angle  $\sqrt{\frac{955}{8rr} - \frac{1}{8_4}}$  For a Third Iris, the Sine of Incidence will be  $\sqrt{\frac{16}{15}}$  and the Sine of the Refracted Angle  $\sqrt{\frac{16ss}{15rr} - \frac{1}{15r}}$  For a

D

Fourth

22

Fourth Iris, the Sine of Incidence will be  $\sqrt{\frac{25}{24} - \frac{1rr}{24^{55}}}$  and the Sine of the Refracted

Angle  $\sqrt{\frac{2555}{24^{\gamma}r} - \frac{1}{24}}$ : and in like manner of

the reft. Farther, 'twill be found by Calculation, that (taking Cartes's Proportion) the Primary Iris is diftant  $41^{\circ}$ .  $30^{\circ}$ . from the Point opposite to the Sun; the Secondary,  $51^{\circ}$ .  $55^{\circ}$ . from the fame. The Third,  $40^{\circ}$ .  $20^{\circ}$ . and the Fourth,  $45^{\circ}$ .  $33^{\circ}$ . from the Sun it felf; which Iris's perhaps were hardly ever feen for the reafons before mentioned.

And thus much may fuffice concerning the Magnitude of the Irides, in the perfpicuons Drops of a Fluid, whofe Refractive Power is known. It remains that nothing be faid concerning the Colours, which this Phenomenon prefents, with the orders of them in each fort of Iris, according to all the possible Variations of the Refraction.

And here we must know especially, that the Acute and Sagacious Mr. Newton, has found by most clear Experiments, that the Rays of Light are not Simple and Uniform, as they iffue out of the Luminous Body, but the pure white Light which we fee, confifts of Corpufcles of all kinds of Colours, mix'd and hurried with a violent Motion, one amongst another. And that the diversity of the Colours of things arifes, according to the various Dispositions those Objects have, to Refract or Reflect this or that peculiar kind of Light.

The Proof of which is manifest from Refractions,

25

ctions, in which thefe Species are feparated from one another, and the *Blue* or *Purple* Light, (even in the fame Diaphanous Body) is more Refracted than the *Tellow* or *Red*. But let the Reader confult this incomparable Perfon's Letters (N°. 80. and the following of the *Philofophical Tranfactions*) from which Specimen he will be able to judge, how nobly this Argument of *Light* will be managed by him.

To my purpofe 'tis fufficient, that all kinds of *Blue* Light, are fomething more refracted than *Red*, from which difference arifes the Latitude of the *Irides*, which is hardly to be determined by Obfervation, becaufe of the uncertain Limits of the Colours. But by how much the Proportion between CA and CD, is of greater Inequality, or by how much the Refraction is greater, fo much the greater is the diffance of any *Iris* from the Sun, and confequently those borders that are remoter from the Sun, fhine with a *Purple* Colour, but those that are nearer, with an intense *Red*.

This may always be feen in the Primary Iris, which vanishes in the part opposite to the Sun, if the Sine of Incidence be to the Sine of the Refracted Angle, as CA to CE, or as 2 to 1. But if that Ratio be greater, there can be no Primary Iris feen at all.

As for the Secondary Iris, 'tis to be noted, that this vanishes into a Point, in the part opposite to the Sun, when the Ratio of the Refraction is as 1 to  $\sqrt{\frac{nn1}{3}} + \sqrt{\frac{4}{27}}$ , or as 1 to 0,847487... and from thence it returns back to the Sum it felf, where it vanishes, if the faid Ratio be as 3 to 1, or as CA to Cc. D 2 But

Worldow Googl

26

But in the Ratio's between these (fuch as we have in all Fluids known to us, except the Air) by how much the greater is the Ratio, by so much is the Iris more distant from the Point opposite to the Sun, or rather from the Sun it felf, reckoning the Arch beyond a Semicircle. And confequently the Colours will seem to be in a different order from the Primary Iris, in these returnings, unless the distance of the Irides from the Sun, be taken in this Sence; which is also every where to be observed in the rest.

The Third Iris is confused in the part opposite to the Sun, the Ratio of the Refraction being as 1 to ,91855--- from thence it returns back to the Sun in the Ratio of 1 to ,68250--- whence again, the order of the Colours being reftored, in the Ratio of 4 to 1, or CA to Ce, it terminates in the part oppofite to the Sun. The Fourth Iris beginning from the Sun, in a Ratio of Equality, paffes on to the opposite Point, in the Ratio of 1 to .94895 --- and thence returns back to the Sun, if the Proportion be as 5 to 4; hence again, it difperfes to the Point opposite to the Sun in the Ratio of 1 to 56337 ---, within which compass are included the Refractions of all Fluids that are known. Laftly, The Ratio being as 5 to 1, or CA to Cn, it vanishes in the very Sun it felf; the Colours being every where inverted to the fight in its return to the Sun, as they were erect in its egress from it. Hence, in watery Clouds, the First and Fourth Iris flew deep red Colours turned towards the Sun; but the Second and Third give Purple. But perhaps I may feem too tedious in

27

in these Descriptions, the Rainbow it self being no more than a Momentary Phantasm.

But whence 'tis that the different Refraetive Power of Fluids arifes, is a Problem of the greatest Moment, and to be rank'd amongst the Secrets of Nature, not yet obvious either to our Sences or our Reafonings. For pure Water amongst all Fluids, does least of all Refract the Rays of Light. When 'tis Tinetur'd with Salts diffolved in it, according to its weight and the quantity of Salt, it increases the Refractions. And Correspoe Spirits (which are much heavier than Water) do alfo much more Refract the Rays of Nor is it any wonder, fince being Light: denfer Bodies, they may eafily be conceived fo much the more to obstruct the passage of the Rays. But why there fhould be fo great a Refraction in Burning Spirits and Oils, especially in Spirit of Turpentine, or of Wine, fince they are Fluids extreamly Light in comparison of Water, and confift very much of fubtle Ætherial Particles, does not fo eafily appear ; but feems to require (in order to the Explication of it) a more thorough knowledge of the Nature and Texture of Light.

But from the diftance (of the Iris from the Sun) given, to find the Ratio of the Refraction, is a thing that will give those that are curious, an occasion of finding the Refraction of any Fluid, accurately and with little trouble. For if a small drop of any transparent Fluid, be supposed to hang at the bottom of a small Glass Tube, and the Sun being near the Horizon and spining strongly, it be obferved under what Angle (with the Point D 3 opposite

Wolanday Goo

28

opposite to the Sun) the Colours of the Iris be feen in the drop, then the Proportion fought will be obtained with a little Calculation. It is a Cubical Equation, explicable by one only Root, by which, from the Primary Iris given, the Ratio is computed, viz. T3-3 T2t-4rrt=0, where T is the Tangent of the Angle of Incidence requisite, t the Tangent of  $\frac{1}{2}$  the diffance of the Iris from the Point opposite to the Sun, Whence (according to to the Radius r=1. Cardanus's Rules) arifes this Theorem, viz. From the Cube of t fubstract the Product of 2tr into the Excess of the Secant of the same Arch above the Radius; the difference shall be the leffer The Sum of the fame, adding Atrr, will Cube. be the greater Cube. The Sum of the fides of both Cubes, and of t, will be equal to the Tangent of the Angle of Incidence, and the half of that, will be the Tangent of the Refracted Angle. From whence the Ratio fought is manifest.

For an Example of this. In a drop of Oil of Turpentine, the diffance of the Primary Iris, from the Point opposite to the Sun, is obferved to be  $25^{\circ}$ .  $40^{\circ}$ . 'Tis required to find the Ratio of the Refraction.

t=Tang. 12°. 50'.		0, 2278063
s=Sec. of the fame.	=	1,0256197
ttt	=	0,01182217
s—r×2tr	=	0,01167265

The Difference is the leffer Cube 0,00014952 whofe fide 0,0530773

The Sum 0, 02349482 4trr 0, 91122525

Greater

Divideo by Google

39

Greater Cube 0, 93472007, whole fide 0, 9777486 t 0, 2278063 T=Tang. Incid. 51°. 32'. 1, 2586322 T=Tang. Refr. 52°. 11'. 0, 6293161

Laftly, As  $\sqrt{TT} + \frac{1}{T} + \frac{1}{TT} + 1 :: r : s :: 1 : 168026}$ . Which Proportion comes very near to that, which Experience flews to be in Glafs and most pellucid Solids. The Diamond indeed, exceeds all transparent Bodies, not only in respect of its hardness and value, but also its Restractive Power, the Proportion here being as 5 : 2, nearly, or more accurately as 100:41. But of this, perhaps more in another place.

While I was writing these things, that skillful Geometrician Mr. De Moivre, at my request, found a like Equation for determining the Ratio, from the Semidiameter of the Secondary Iris, given. By which, the Ratio is indeed fomething more exactly determined, but that Equation being a Biquadratical one, the Calculation is not fo eafily performed. This Equation is  $T^4 + \frac{4}{4}T^3 t - 2T^2 T^2 - \frac{1}{3}$  $r^4 = 0$ ; where T is the Tangent of the Refracted Angle, t, the Tangent of 1/2 the distance of the Iris from the Point opposite to the Sun, to the Radius r=1. And this Equation is of that Form, as to be always explicable, by an Affirmative and one Negative Root, the one and the lefs of which, is the Tangent of the Refracted Angle, in the Regrefs to the Sun, viz. when the Purple Colours are nearer to the Sun. The greater Root is the Tangent of the Refracted Angle in an Iris D 4

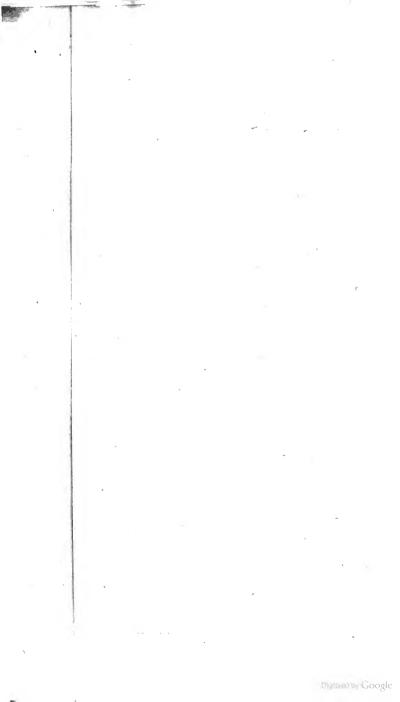
40

Iris going out from the Sun, viz. in a Fluid of a les Ratio. In Oil of Turpentine, the distance of this Ivis from the Point opposite to the Sun, is observed to be 81°. 30'. whence the curious Reader may find out the Roots, 0, 80822 -- and-2, 98131 -- the Tangents of the Refracted Angles. Hence is computed the Ratio of greater Inequality, as I to 0,67995 -fuch as is in Oil of Turpentine. But from the greater Root comes forth the leffer Ratio, as 1 to 0,9540 nearly, fuch as would be in a Fluid, exhibiting a Secondary Iris of the fame Diameter, but which (after the manner of the Primary one) fhould look towards the Sun with the Red Colours.

If any one has a mind to find these Roots by a Geometrical Construction, any Parabola being given, it is done with so much ease, that I need not repeat what I have already offered upon that Head Philosophical Transactions, N. 188.

Each of these Equations is deduced from what has been faid before, and also from the Rules for the Tangents of the Double and Treble Arch; the bare hint of which, may be instead of a Demonstration even to those that are but meanly vers'd in these things.

This Difcourfe being already in the Prefs, there came to my hands (by the means of a Friend) a certain Book, whole Title was Thaumantiadis Thaumafia, Printed at Norimterg 1699, under the Superintendency of M. Sturming. In which the skillful Author feems to have laid together whatever is to be found





of this Argument, as well amongst the Madern as the Ancient Writers ; fubjoyning and illustrating Cartes's, Eckard's, Honoratus Faber's, and Mariott's Calculus. From whence it is plain, that the reft added very little or nothing to Cartes's Inventions, building upon the fame Ungeometrical and Tentative Methods of Calculation. But that the Judicious Reader may be fenfible, what things I have performed, in the Doctrine of the Iris, I would have him read the forementioned Book, and compare it with this Difcourfe; left in putting out these things, I should seem only to have made an unpleafing Repetition of what had been done before. And of what vast use in Astronomical Matters, this Lemma of ours may be, shall be shewn upon some other occasion.

## FINIS.

Distantly Google

Books Printed for, and Sold by Jeffery Wale, at the Angel, in St. Paul's Church-Yard.

M Ifcellany Poems, as Satyrs, Epiftles, Love-Verfes, Songs, Sonners, &c. by William Wycherley, Efq; Fol.

A Supplement to Dr. Hammond's Paraphrafe and Annotations of the New Teltament, by Mr. L'Clerk. Quarto. To which is prefix'd a Letter from the Author to a Friend in England, occafioned by this Translation.

The Posthumous Works of Mr. de St. Evremont, containing variety of elegant Esfays, Letters and Poems; and other Miscellaneous Pieces on several curious Subjects. Vol. III.

The plain Man's Guide to Heaven. By Dr. Lucas. In 125.

The practice of Phyfick reduced to the ancient way of Obfervations, containing a just parallel between the Wifdom and Experience of the Ancients, and the Hypothefis of Modern Phificians, with new and curious Obfervations on the Tarantula. Octavo.

A Treatife of the Two Covenants, by J. Parker. Octavo.

An

Conditioned by Google

An impartial Account of the Affairs of Scotland, from the Death of King James V. to the Tragical Exit of the Earl of Murray, Regent of Scotland, by a Perfon of Quality. Octavo.

The Church of England proved to be Conformable to, and Approved by all the Protestant Churches in Europe. In Octavo. Price 6 d.

The Hiftory of the famous Knight Don Quixot de la Mancha, Vol. III. by Capt. John Stevens. Never before done into English.

Copernicans of all forts convicted, by the Honourable Edward Howard of Berks.

An Effay at the Machanism of the Macrosm, or the dependance upon their Causes, in a new Hypothesis, accommodated to our Modern and Experimental Philosophy, by C. Pursbal, M. D. in Octavo.

The Hiftory of Holland, in 2 Vol. illustrated with Cuts. Octavo.

A Treatife of the Difeafes of Tradefmen, fhewing the various Influence of particular Trades upon the ftate of Health, with the beft Methods to avoid or correct it. By Beru. Rammazini, Professor of Physick at Padua.

A new and accurate Description of the Coast

Coast of Guinea, divided into the Gold, the Slave, and the Ivory Coasts. Illustrated with Cuts, and an exact Map of the whole Coast. Octavo.

A brief but plain Explanation of the Church-Catechifm; by Tho. Cooke, A. M. Lecturer of the United Parishes of St. Margaret Pattons and St. Gabriel Fenchurch. Price 3 d. or 20 s. a hundred.

Synopsis Palmariorum Matheseos, or a new Introduction to the Mathematicks. containing the Principles of Arithmetick and Geometry demonstrated, in a short and easie Method, with their Application to the most useful parts thereof: As, refolving of Equations, infinite Series, ma-king the Logarithims, Interest fimple and compound; the chief Properties of the conick Sections, menfuration of Surfaces and Solids; the fundamental Precepts of perspective Trigonometry: the Laws of motion applied to mechanick Powers, Gunnery, &c. Defign'd for the benefit, and adapted to the capacities of beginners. By W .Jones

Books

To grow Google

Books and Maps fold by John Senex, next the Fleece-Tavern in Cornhil.

A Tlas Caleftis: Containing the Syftems and Theories of the Planets, the Conftellations of the Stars, and other Phenomena's of the Heavens. Price 55.

A Pocket-Book, containing feveral choice Collections in Arithmetick, Aftronomy, Geometry, Surveying, Dialing, Navigation, Altrology, Geography, Meafuring, Gauging, & c. Price 5 s.

The Theory of the Handling or working of fhips at Sea, the like never before publisted.

A New Pair of Globes, Twelve Inches Diameter. The Terrestial is laid down acdording to the newest Discoveries, and from the most exact Observations, with a general View of the Trade-Winds, and *Monson's*. The Cœlestial has the Stars Places, from the Correct Tables of M. *Hevelij*, Capt. *Halley*, and the like, never before extant, Price 3 l.

A New Syftem of Geography, defign'd in a most plain and easie Method for the better understanding that Science: Accommodated with new Maps of all the Empires, pires. Kingdoms, Principalities, Dukedoms, Provinces and Countries in the whole World; with Geographical Tables, explaining the Divisions in each Map. The Third Edition. To which is added, An Introduction to that Science. Price 6 s.

10,000

The General Laws of Nature and Motion, with their Application to Mechanicks. Alfo, The Doctrine of Centripetal Forces, and Velocities of Bodies, defcribing any of the Conick Sections, being a Part of the Great Mr. Newton's Principles. The Whole illustrated with Variety of ufeful Theorems and Problems, and accomadated to the Ufe of Youger Mathematicians. By H Ditton. Price 4 s.

A Hexagon, fortified with all Sorts of Out-Works, according to the modern Method of Fortification; together with all the Infruments ufed in Attacking and Defending of Places, ufeful for all that read the Publick News. Curioufly Engraven. Price 15.

A Treatife of Trigonometry, Plain and Spherical, Theoretical and Practical; in which the feveral Cafes of Plain and Spherical Triangles are folv'd Inftrumentally and Arithmetically: As likewife a Treatife of Steriographick and Orthographick Projection of the Sphere. By Sam. Haynes,

2

Digreeder Google

Haynes, late Reader of the Mathematicks to His Majesty's Engineers. Price 3 s.

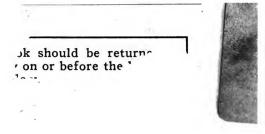
A New Map of the Seat of War in Flanders, on the Rhine, the Meuse, the Mosfel, &c. Price 6 d.

A Plan of the Town, Caffle, Moles, and Bay of Gibraltar, as Attack'd by the French, 1705. Price 6 d.

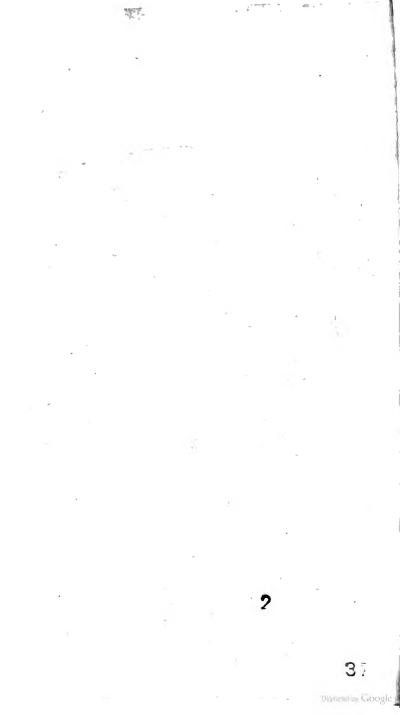
Alfo a Plan of the City and Cittadel of Liege. Price 3 d.

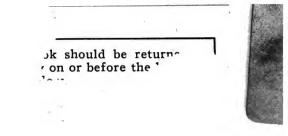
Where may be had all Sorts of Mathematical Books, Maps and Instruments, for the Sea or Land.



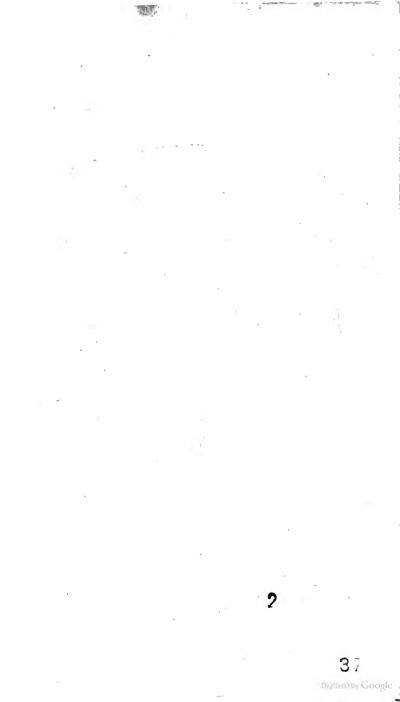


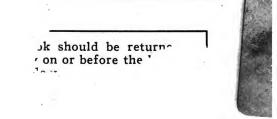




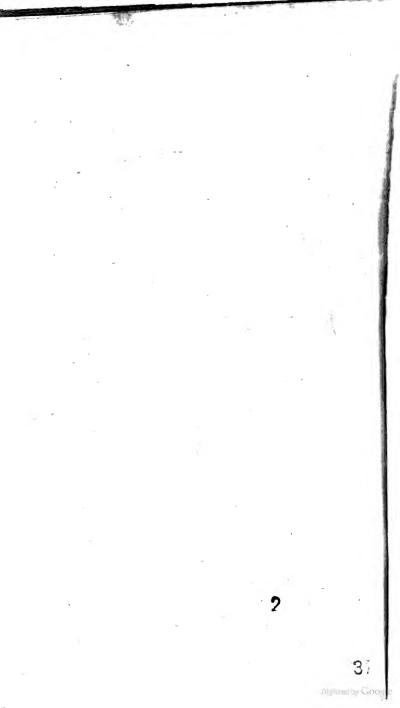


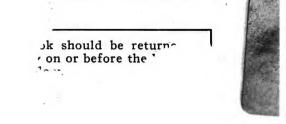




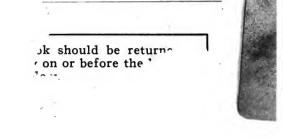




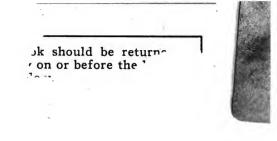






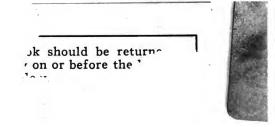








and by Google





-----

1 - 3⁰⁰ 0.1 28 - 1030

* * *

the information of

Bineria ai Ciarty)